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Eco-friendly and Non Vacuum Electrostatic Spray-Assisted Vapour Deposition of Cu(In,Ga)(S,Se)₂ Thin Film Solar Cells.

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ABSTRACT: Chalcopyrite Cu(In,Ga)(S.Se)₂ (CIGSSe) thin films have been deposited by a novel, non-vacuum and cost-effective Electrostatic Spray-Assisted Vapor Deposition (ESAVD) method. The generation of a fine aerosol of precursor solution, and their controlled deposition onto molybdenum substrate, results in adherent, dense and uniform $Cu(In,Ga)S₂ (CIGS)$ films. This is an essential tool to keep the interfacial area of thin film solar cells to a minimum value for efficient charge separation as it helps to achieve the desired surface smoothness uniformity for subsequent cadmium sulfide and window layer deposition. This non-vacuum aerosol based approach for making the CIGSSe film uses environmentally benign precursor solution, and it is cheaper for producing solar cells than that of the vacuum-based thin film solar technology. An optimized CIGSSe thin film solar cell with a device configuration of molybdenum-coated sodalime glass substrate/CIGSSe/CdS/i-ZnO/AZO shows the photovoltaic (*j-V*) characteristics of *V*oc $= 0.518$ V, $j_{\rm sc} = 28.79$ mA cm⁻², *fill factor* = 64.02%, and a promising power conversion efficiency of η = 9.55% under simulated AM 1.5 100 mW cm⁻² illuminations, without the use of an antireflection layer. This demonstrates the potential of ESAVD deposition as a promising alternative approach for making thin film CIGSSe solar cells at a lower cost.

KEYWORDS: Cu(In,Ga)(S,Se)₂, ESAVD deposition, selenization, thin film, solar cells.

1. INTRODUCTION

Chalcopyrite $Cu(In, Ga)Se₂$ (CIGS) thin film solar cells are advantageous over conventional crystalline silicon solar cells in terms of materials processing and the associated cost of production with respect to crystalline silicon solar cells. Its tunable optical properties (band gap of 1.05 - 1.5 eV) and excellent light harvesting characteristics (absorption coefficient of 10^5 cm⁻ ¹) allow low usage of precursor materials in order to obtain light harvesting efficiency of close to unity. The research on CIGS thin films and their decent efficiency solar cells leads to the discovery of several emerging light absorbing materials, such as kesterite $Cu₂ZnSnSe₄,^{1, 2}$ cadmium telluride (CdTe),^{3, 4} and lead sulfide (PbS) quantum dots,⁵ which have all shown their potential in the conversion of solar energy into electricity at low cost.

While research on CIGS solar cell began decades ago, efforts have mostly focused on the optimization of vacuum-based thin films which has recently resulted in unprecedented power conversion efficiency (PCE) of 21.7% among all thin film solar cells including CdTe (21.4%) and amorphous silicon (10.5%) ^{3, 6} The vacuum methods, such as multistage coevaporation/sputtering processes, are well-optimized for industrial production lines, however, they are costly because of expensive instrumentation and use of sophisticated vacuum systems.⁷⁻⁹ It is also challenging to maintain compositional uniformity for large area CIGS films and a H_2 Se gas treatment step is often required, making the CIGS solar cells vulnerable in terms of associated cost and safety in production.^{7, 10, 11} Therefore, it is crucially important to prepare and optimize CIGS thin films by non-vacuum methods so that solar cells can be prepared more cheaply than the existing vacuum methods.

approaches; including spin-coating, $^{12-19}$ spray-pyrolysis, $^{20-24}$ electrodeposition, 25 and electrostatic spray-assisted vapor deposition $(ESAVD)^{26-28}$. These techniques mostly use a molecular precursor solution consisting of either precursors containing sources of copper (Cu), indium (In), and gallium (Ga) with non-metallic sulfur (S) and selenium (Se), or pre-synthesized nanoparticles. These solution-based approaches allow selected precursor materials to be prepared into a CIGS absorber film and have similar potential to those of vacuum methods. Among these, the CIGS film prepared by non-vacuum hydrazine (N_2H_4) -based molecular precursor solution led to the highest 15.2% efficiency solar cells by Mitzi *et al.*, however, the N_2H_4 is toxic and explosive in nature, making it undesirable in industrial production lines.^{29, 30} Solar cells with an efficiency of 8.75% have been reported by dissolving metal oxides into relatively less toxic solvents, such as carbon disulfide (CS_2) , to prepare the molecular precursor materials.³¹ These

In addition, more than 90% of molecular precursor materials and/or nanoparticles are being wasted on the surrounding area as spin-off material during the spin-coating process,, which is also believed to be less suitable for large scale production. The usage of electrolytes during electrochemical methods is also high which would pose environmental issues for the treatment and safe disposal of the large volumes used for the electrochemical bath, especially with the increasing tighter environmental regulations. Very recently 10.7% efficiency CIGSSe solar cells have been reported by spray-pyrolysis approach using sodium (Na) incorporation, and with an antireflection coating.²³ Other non-vacuum precursor solution and nanocrystal inks-based approaches also require a post-selenization or sulfurization steps, such as spray-pyrolysis-based CIGSSe solar cells have shown efficiency of 10.54% ,³² electrodeposited CIGS solar cells with

efficiency of 10.93%,³³ and sol-gel approach CIGS solar cell showed 7.04% efficiency.³⁴ One of the key advantages of the ESAVD approach over other non-vacuum deposition, such as spray pyrolysis, is the use of an electric field during deposition which helps to direct the atomized aerosol chemical precursor to the substrate, thus minimizing the loss of precursor to the surrounding and increase the deposition yield. In fact, the deposition efficiency could be reached close to unity, which is higher than spray pyrolysis, and it is an ideal approach for the large area fabrication of CIGS films. $28, 35$

Based upon these observations, we adopted the ESAVD approach to prepare the CIGS films using environmental friendly precursor materials. A non-toxic alcohol-based solvent, which can dissolve precursor materials was utilized to prepare good quality CIGS films. A post-selenization step is carried out to remove any residual solvent from the as-deposited CIGS films and to densify the films so that photovoltaic grade CIGSSe films can be obtained. An optimized CIGSSe solar cell with a device configuration of SLG/Mo/CIGSSe/CdS/i-ZnO/AZO/Ni/Al shows the *j*-*V* characteristics of $V_{\text{oc}} = 0.518$ V, $j_{\text{sc}} = 28.79$ mA cm⁻², *fill factor* = 64.02%, and power conversion efficiency (PCE) of $η = 9.55%$ under simulated AM 1.5 100 mW cm⁻² illuminations without Na incorporation into CIGSSe or an anti-reflection coating. Our results show the potential of non-vacuum ESAVD for the fabrication of high efficiency CIGSSe thin films using an environmental-friendly, highly reproducible, sustainable and energy efficient nonvacuum approach with a much lower cost, and there is still plenty of room for further optimization. In overall, the application of non-vacuum ESAVD approach is a very promising way to increase the CIGS deposition rate and to remove the need for expensive high-vacuum systems.

2. EXPERIMENTAL SECTION

Deposition of CIGS films. The Molybdenum (Mo)-coated soda-lime glass (SLG) substrate was sequentially cleaned by sonication in an aqueous solution of 5% Decon D90 solution and DI water, followed by dipping into 10% ammonium hydroxide aqueous solution and drying in nitrogen (N_2) stream prior to the deposition of CIGS films. Metal salt precursors of copper(II) chloride (CuCl₂), indium(III) chloride (InCl₃), gallium(III) nitrate (Ga(NO₃)₃), and thiourea $(SC(NH₂)₂)$ with molar concentration within the range 0.2 to 0.7 M were prepared .The precursor salts dissolved in a mixture of ethanol-water based solution and atomized and directed to the heated substrate; as schematically shown in Figure 1. Details of the ESAVD method can be found in the literature.^{28, 35} In brief, an applied voltage of 5-15 kV was used and the precursor flow rate was maintained up to 85 ml/h. The atomizer-to-substrate distance was maintained up to 5 cm. During the ESAVD process, the precursor mixture is atomized to form an aerosol which is charged and subsequently directed towards a heated substrate where it undergoes decomposition and chemical reactions to deposit as a stable solid film of CIGS onto the SLG/Mo substrate. The composition of the CIGS films absorber can be finely controlled by adjusting the ratios of metallic precursors of Cu, In and Ga in the precursor solution. To prepare the as-deposited CIGS-1, CIGS-2, and CIGS-3 thin films, the precursor ratio of $Cu/(In+Ga)$ was varied from 0.6, 0.75 to 0.78, respectively, while the ratio of $Ga/(In+Ga)$ was maintained at 0.25 for all films. A ratio of $Cu/(In+Ga) = 0.62$ was found for an optimized solar cells.

Figure 1. Showing a schematic diagram of Electrostatic Spray-Assisted Vapor Deposition (ESAVD) method to prepare the as deposited CIGS thin films onto SLG/Mo substrate.

The as-deposited CIGS films were selenized in a quartz tube furnace (Elite Furnace TSH12/75/750) to allow grain growth and to densify the CIGS films in order to obtain highly crystalline, dense and pure photovoltaic grade CIGSSe films. A Selenium (Se) pellet of 55 mg was taken into a quartz tube at an optimized temperature of 550 ºC for 12 minutes. A preannealing step at 300 ºC is required to remove any source of carbon contaminants prior to finally selenize at 550 ºC to incorporate selenium into CIGS films and to enhance PCE of the CIGSSe solar cells. The selenization converts as-deposited CIGS films into CIGSSe in which various defect states at grain-boundary regions, and surface states are greatly suppressed. After selenization, the CIGSSe films were cooled down to room temperature and transferred immediately into a chemical bath for CdS deposition.

Characterization of thin films. Crystal structure of the as-deposited CIGS and post-selenized CIGSSe films were characterized by a grazing angle incidence X-ray diffraction (XRD) technique with a Bruker D8 discovery diffractometer using Cu Ka1 radiation ($\lambda = 0.15406$ nm). Raman spectroscopy measurement was carried out with a micro-Raman spectrometer (Renishaw

inVia Raman Microscope) using a diode-pumped solid-state laser with excitation wavelength of 514 nm. The microstructure of the CIGSSe films was characterized using a variable pressure analytical scanning electron microscope (SEM, JEOL JSM-6480LV). Prior to observation under electron microscopy, a thin Au layer was deposited to prevent charging effects. The elemental analysis of the CIGSSe films was carried out using an Oxford Link system Energy-Dispersive Xray Spectroscopy (EDS) fitted with the SEM. An X-ray fluorescence analyzer system (Fischer, XRF-250) was utilized to quantify elemental composition of the as-deposited CIGS films.

Fabrication of solar cells. An approximately 50 nm thick CdS buffer layer was deposited onto CIGSSe films to obtain the CIGSSe/CdS heterojunction required for the separation of photogenerated charge carriers. While a few alternative buffer layers to the CdS is emerging, CdS is still normally used because of its small lattice mismatch with the CIGS films, and it has shown beneficial band energetics with the CIGS and transparent conducting oxide (TCO). In this study, a water bath at 80 ºC was utilized as a heat-source for the CIGSSe films containing aqueous solution of 4 mM cadmium acetate hydrate $(Cd(CH_3COO)_2, 99\%)$, 4 mM ammonium acetate (NH₄(CH₃COO), 99.99%), 2 mM thiourea (SC(NH₂)₂, 99%), and 0.4 ml ammonium hydroxide solution (NH4OH, 28-30% solution). After the deposition of CdS for 10 minutes, blue colored CIGSSe films were rinsed with DI water and dried in a N_2 stream. Subsequently, an approximately 50 nm thick intrinsic zinc oxide (i-ZnO) film was deposited to protect the CdS buffer layer from plasma damage during the subsequent deposition of aluminium-doped ZnO (AZO) as the TCO layer. The i-ZnO and approximately 600 nm thick AZO layers were deposited by HHV sputtering system with RF power of 100 W for 15 minutes and 220 W for 80 minutes, respectively, to efficiently collect the separated charge carriers. The AZO layer was utilized as TCO because of its more than 90% transparency to the visible solar spectrum and it has a low

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electrical resistivity of $\leq 10^{-4} \Omega$ cm which allows light to enter into the CIGSSe films and collect the photogenerated charge carrier efficiently before being recombined, respectively.^{36, 37}

Finally, in order to collect photogenerated charge carriers and transport them with a minimum resistive losses, a patterned bilayer top electrode of Ni/Al (75 nm /400 nm) was thermally evaporated onto the above electrodes using a HHV thermal evaporation system.³⁶ Subsequently, they were mechanically scribed to define the active area of individual pixel solar cells to be 0.15 cm². Solder wire containing 60% tin and 40% lead was deposited onto Mo substrate to make a better contact with measuring cables during device characterization.

Characterization of solar cells. A solar simulator (Oriel Sol 1A) was utilized to measure photovoltaic, j -V characteristics of the CIGSSe solar cells under simulated AM 1.5 100 mW cm⁻² illumination using a Keithley Source (2400) meter and Oriel IV test station software package. The solar simulator was equipped with a 500 W Xenon lamp and the light intensity was calibrated using a silicon (Si) reference solar cell (Oriel instruments). External quantum efficiency (EQE) spectra were obtained using a Spequest quantum efficiency photovoltaic system (Rera) equipped with a xenon/quartz lamp. The incident photon flux was determined using the calibrated silicon (Si) and germanium (Ge) photodiodes which allowed us to measure the EQE up to photon wavelength of 1800 nm. The recording of photocurrent spectra were performed using a Photor software package.

3. RESULTS and DISCUSSION

Structural characterization of the thin films. The thin films were characterized using a grazing angle incidence XRD with an optimized incident angle of 1.25 degrees. As shown in Figure 2, the XRD patterns show characteristic diffraction peaks of CIGSSe-1, CIGSSe-2,

CIGSSe-3 and optimized CIGSSe films with various elemental compositions. The XRD patterns of the as-deposited CIGS-1, CIGS-2, and CIGS-3 films show that they consist of crystallites of CIGS (Supporting information, Figure S1).

Figure 2. XRD patterns of CIGSSe thin films with various Cu/(In+Ga) ratios fabricated onto Mo-coated SLG substrate. The Cu/(In+Ga) ratio in the CIGSSe-1, CIGSSe-2, CIGSSe-3, and optimized CIGSSe films are observed to be 0.77, 0.79, 0.89 and 0.78, respectively.

The narrow full width at half maximum (FWHM) and intense XRD peaks suggest the presence of well-crystalized grains in the selenized polycrystalline CIGSSe films (Figure 2). The XRD patterns confirmed the chalcopyrite structure and phase purity of the CIGSSe material. No apparent XRD peaks corresponding to unwanted secondary phases are observed. As all the films consisted of almost equal amounts of Ga (supported by observed same $Ga/(In+Ga)$ ratio of ~ 0.25 from EDS result, Table 1), a shift of the XRD peaks among the CIGSSe films, which occurs due to Ga incorporation and subsequent change of crystal lattice, was not observed.³² A broad and

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low intensity peak at ca., $2\theta = 31^{\circ}$ for all samples implies the existence of a thin MoSe₂ layer formed due to the diffusion of Se^{2} vapour through as-deposited CIGS films.³⁸ The absence of XRD peaks corresponding to $Cu_{2-x}Se$ confirms that these CIGSSe films are phase pure.³⁹

Morphological characterization. Figure 3 shows typical morphological characteristics of the optimized CIGSSe films investigated by SEM. The top-view image, as shown in Figure 3(a), illustrates that the materials exhibits interconnected large grains of CIGSSe with a size of 0.8 – 1.5 μ m which ultimately forms a dense, thin film. The corresponding cross-sectional image indicates the thickness of the CIGSSe film (Figure 3(b)) to be an average of 1 μ m, and covers the entire Mo substrate. Such an interconnected grainy structure and compact CIGSSe film also indicates an improved selenization step when process temperature, time, and amount of selenium, which assists grain growth are considered. As demonstrated by Figure 3(b), the dense CIGSSe film may lead to a significantly thin $Mose₂$ layer onto the Mo substrate because of suppressed diffusion of reactive Se^{2} through inter-grains of the CIGSSe films to reach the substrate. An insignificant variation of grain size of 0.8-1.4, 0.8-1.4, and 0.85-1.6 μ m was observed for the CIGSSe-1, CIGSSe-2 and CIGSSe-3 thin films, respectively (Figure S2, Supporting information).

Figure 3. (a) Surface morphology and (b) cross-sectional image of the optimized CIGSSe thin film.

Compositional analysis. Compositional analysis of the as-deposited CIGS films was conducted by XRF which shows calculated values of $Cu/(In+Ga)$ as 0.60, 0.75, 0.78 and 0.62, and for Ga/(In+Ga) as 0.24, 0.23, 0.23 and 0.22 for the as-deposited CIGS-1, CIGS-2, CIGS-3, and optimized CIGS films, respectively (supporting information, Table S1).

Table 1. Compositional study of the CIGSSe thin films by EDS elemental analysis.

Thin Films								Cu (at%) In (at%) Ga (at%) S (at%) Se (at%) Cu/(In+Ga) Ga/(In+Ga) Se/(S+Se)
CIGSSe-1	18.65	17.82	6.15	10.56	46.82	0.77	0.25	0.81
CIGSSe-2	16.51	15.85	5.06	13.29	49.29	0.79	0.24	0.78
CIGSSe-3	21.69	18.01	6.26	9.93	44.11	0.89	0.25	0.81
Optimised								
	20.07	18.10	7.44	3.61	50.78	0.78	0.29	0.93
CIGSSe								

The Cu deficiency is required in the films to eliminate any chance of formation of secondary phases of $Cu_{2-x}S$, $Cu_{2-x}Se$, $Cu_{3}Se_{2}$, and CuSe. However, as shown in Table 1, after selenization, the Cu/(In+Ga) ratio has increased to 0.77, 0.79, 0.89, and 0.78 in the CIGSSe-1, CIGSSe-2,

CIGSSe-3, and optimized CIGSSe film, respectively. These values are close to the known optimum value of 0.85 for high efficiency CIGS solar cells. The increase of the Cu/(In+Ga) and Ga/(In+Ga) ratios suggests a loss of In during the selenization step. These Cu-deficient CIGSSe films were utilized to fabricate the solar cells.

Raman spectroscopy. The crystallographic data alone is not sufficient to accurately determine the phase purity of the CIGSSe films as they often produce non-stoichiometric phases during selenization where the XRD peaks could overlap in the diffraction patterns of CIGSSe, leading to an incorrect conclusion. Therefore, the films were studied by Raman spectroscopy to correctly probe the insight of the CIGSSe films and deduce whether impurity phases of $Cu_{2-x}Se$, $Cu_{3}Se_{2}$, CuSe, were present in the films, despite the fact that the XRD patterns, as shown in Figure 2, confirmed their absence. The films selenized at 550 ºC show the dominant Raman bands of A1 mode at 178 cm⁻¹ with narrow FWHM and also show mixed B2/E modes at around 218 cm⁻¹ (Figure 4).^{40, 41}

Figure 4. The Raman spectra of the CIGSSe thin films deposited onto Mo substrate with various Cu/(In+Ga) ratios.

The vibrational mode at around 220-260 cm⁻¹ corresponds to a mixed mode of $B_{2/F}$ of Cu-poor CIGSSe films.⁴² It also implies that there are no apparent impurity phases of $Cu_{2-x}Se$, CuSe and Cu3Se2 in the CIGSSe films and the films are good quality. The intensities of Raman bands at around 295 cm-1, correspond to CIGS films, they become insignificant with respect to that of 177 cm⁻¹, correspond to CIGSSe films, confirming a high Se/S in the films. The Raman spectra do not reveal a broad vibrational mode at around 150 cm^{-1} , implying that there are no apparent order defects of Cu(In,Ga)₃(S,Se)₅ or Cu₂(In,Ga)₄(S,Se)₇ which tend to occur in Cu-poor CIGS films.⁴³

Photovoltaic *j-V* **characteristics.** Solar cells were fabricated using CIGSSe thin films with a conventional device structure of SLG/Mo/CIGSSe/CdS/i-ZnO/AZO/Ni/Al. The decent *j-V* characteristics of solar cells are tabulated in Table 2 and the *j-V* curves are shown in Figure 5 (a). As compared with the CIGSSe-2, the efficiency of the solar cells is found to be improved when using the film of CIGSSe-1, which has slightly higher $Ga/(In+Ga) = 0.25$ and lower Cu/(In+Ga) $= 0.77$, this may suggest its excellent electronic properties (charge carrier mobilities, densities, *etc.*). For the CIGSSe-3 film, despite the fairly similar Ga/(In+Ga) ratio of 0.25 with CIGSSe-1, it exhibited poor rectification behavior because of high value of Cu/(In+Ga) ratio of 0.89. The optimized solar cell showed the best efficiency of 9.55% with $V_{\text{oc}} = 0.518 \text{ V}, j_{\text{sc}} = 28.79 \text{ mA cm}^2$, and $FF = 64.02\%$.

The PCE of these solar cells can be improved further by optimizing the thickness of CIGSSe films and using a TCO layer with reduced optical and resistive losses. In general, the V_{oc} value should increase with CIGSSe film thickness, provided that they are pure and crystalline with suppressed trap states for carrier recombination. The fill factor of the optimized solar cells is 64.02%, implying that the CIGSSe/CdS interface needs to be optimized along with reducing

series resistance (R_s) and raising shunt resistance (R_{sh}) . A much higher PCE of CIGSSe solar cell is expected if these parasitic resistances are reduced to a greater extent.

Figure 5. (a) *j-V* characteristics of CIGSSe thin film solar cells measured under simulated AM 1.5 100 mW cm⁻² illuminations, (b) the EQE spectra of corresponding solar cells. The inset of Figure 5(b) shows the Tauc-plot using EQE spectra and the approximate band gap values of CIGSSe-1, CIGSSe-2, CIGSSe-3 and optimized CIGSSe films. The top section of Figure 5(b) shows the EQE ratio, [EQE(-1 V)/EQE(0 V)], for the CIGSSe-3 solar cell.

EQE measurements, as depicted in Figure 5(b), show conversion of photons beyond 1100 nm in the solar spectrum. The highest photocurrent response is observed to be approximately 75% for the CIGSSe-1 solar cells. However, this peak EQE is actually much lower than typical vacuum-based solar cells, usually above 90% , suggesting the major loss in $j_{\rm sc}$ could be due to comparatively poor carrier collection efficiency, pronounced recombination, short lifetime of minority carriers resulting from defect states at various interfaces, and insufficiently thick CIGSSe films with low light harvesting efficiency. In addition, the EQE curve of the CIGSSe-3 solar cell is found to be more deficient in the long wavelength region, near the band edge, which

indicates a pathway for recombination loss which is deep in the CIGSSe layer, or a shorter minority carrier diffusion length, which limits the carrier collection efficiency. The EQE bias ratio plot, the ratio of EQE measured at -1 V and 0 V, $[EQE(-1 V)/EQE(0 V)]$, as shown in the top section of Figure 5(b) shows an increasing $[EQE(-1 V)/EQE(0 V)]$ ratio with photon wavelength, especially at longer wavelength, suggesting a voltage-dependent collection efficiency which occurs in solar cells with shorter minority carrier diffusion length. Therefore, further improvement of films quality (*i.e.,* larger grain size, passivation of grain boundaries incorporating alkaline materials, and optimizing the heterojunction interfaces) is crucial to obtain enhanced carrier collection efficiency in solar cells.

Provided that the carrier collection is efficient in device, an improvement of $j_{\rm sc}$ is expected by increasing the light harvesting efficiency using a relatively thicker CIGSSe absorber layer that may also prevent the shunting pathways in device. The *j*sc of the solar cells were relatively lower than that of the reported high efficiency vacuum-based CIGS solar cell ($>$ 32 mA cm⁻²). This is mainly due to the use of a thinner CIGSSe film (average thickness 1 *µm*), which was much lower than the optimal value $(2 \mu m)$. In addition, lowering bulk resistance of the different stack layers, *i.e.*, Mo, CIGSSe, CdS and AZO layers, and improving the charge transport characteristics at these adjacent interfaces, are alternative ways to further improve the $j_{\rm sc}$. Overall, the j -V characteristic of these solar cells suggest that there are still several paths to improve their efficiency, *i.e.*, photon management by incorporating an antireflection coating to enhance light harvesting efficiency of these solar cells, and consequently increasing their power conversion efficiency. The reduced EQE below a photon wavelength of 500 nm is due to the light absorption by CdS, buffer (i-ZnO), and window (AZO) layers, which can be reduced by using their optimized layer thickness.44 The reduced EQE in near-infrared region (1050 nm onwards) could

also be due to incomplete absorption of incident light by the CIGSSe films and/or due to reflectivity of free-carrier absorption from the AZO contact and various recombination losses in this long wavelength region.⁴⁴

Table 2. The *j*-*V* characteristics of CIGSSe solar cells under simulated AM 1.5, 100 mW cm⁻² illumination.

Solar cells	$V_{oc}(V)$	$j_{\rm sc}$ (mA cm ⁻²)	FF(%)	η (%)	$R_{\rm s}$ (Ω cm ²)	$R_{\rm sh}$ (Ω cm ²)
CIGSSe-1	0.540	26.12	63.46	8.95	3.16	264.76
CIGSSe-2	0.549	25.69	53.93	7.61	4.72	120.27
CIGSSe-3	0.475	28.23	41.95	5.62	5.87	89.85
Optimized CIGSSe	0.518	28.79	64.02	9.55	2.83	408.95

The CIGSSe-3 solar cells showed comparatively lower EQE which could be the result of inefficient charge separation in Cu-rich CIGSSe-3 materials and its heterojunction interface with CdS. The E_g of the CIGSSe-1, CIGSSe-2, and CIGSSe-3 were calculated to be 1.2 eV, 1.19 eV, and 1.18 eV, respectively from the corresponding Tauc-plot of the EQE spectra [inset of Figure 5(b)]. These are surprisingly similar because of similar Se/(S+Se) ratios of 0.78 - 0.81 (Table 1). In this context, the optimized CIGSSe solar cells show slightly lower E_g values of approximately 1.15 eV because of higher Se/(Se+S) ratio of 0.93 (Table 1), implying incorporation of more Se during the selenization step. It is worth noting that a high efficiency co-evaporation-based CIGS device requires an approximately 2 *µm* thick absorber layer with an *E*g values of 1.15 - 1.18 eV.⁴⁵ The wavelength dependent quantum efficiencies for the these solar cells, in the range of 500 to 850 nm, could arise from optical losses from AZO layer, which requires further studies for better understanding of different stack layers.

Figure 6. Cross-sectional image of the optimized CIGSSe solar cell with stack layers of Mo, CIGSSe, CdS, i-ZnO and AZO.

Figure 6 shows a cross-sectional SEM image of the optimized CIGSSe solar cell made onto Mo-coated SLG substrate. The CIGSSe film has a thickness of circa, 1 *µm* consisting of large crystal grains of more than 1 *µm*. There was no void in between the Mo and CIGSSe layers. An amorphous conformal CdS layer of approximately 50 nm thick can be seen on top of the CIGSSe film. The SEM image suggests that the M_0Se_2 layer, which in generally formed during the selenization step, is actually very thin (Figure 6). We believe that the highly crystalline and pure CIGSSe thin films prepared with environmental friendly precursors, and the cheaper non vacuum ESAVD method, will allow us to fabricate solar cells with promising efficiency in more costeffective ways. This is very encouraging for further development of alternative low-cost and high throughput non-vacuum processed CIGSSe thin film solar cells.

4. CONCLUSIONS

In summary, CIGSSe thin films have been successfully fabricated using an environmental friendly, non-vacuum ESAVD approach with subsequent selenization. The composition of CIGSSe films was varied using different amounts of precursor solution containing Cu, In, Ga and S. The as-deposited uniform CIGS films were nanocrystalline and adhered well onto Mo substrates. After the selenization step, they were transformed into highly crystalline CIGSSe thin films. The XRD and Raman study revealed the phase purity of the CIGSSe thin films. The solar cells fabricated with an optimized and highly crystalline CIGSSe films showed power conversion efficiencies of 9.55% under simulated AM1.5 100 mW cm⁻² solar irradiation. These results show that the ESAVD is a promising approach for the fabrication of cost-effective chalcogenide thin film solar cells.

5. ASSOCIATED CONTENT

Electronic Supplementary Information (ESI) available. XRD patterns (Figure S1) and composition (Table S1) of the as-deposited CIGS thin films. SEM images of the CIGSSe thin films (Figure S2). This material is available free of charge via the Internet at http://pubs.acs.org.

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All authors have given approval to the final version of the manuscript. $!$ These authors contributed equally.

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