

## Inclusive Jet Cross Section in $\bar{p}p$ Collisions at $\sqrt{s} = 1.8$ TeV

B. Abbott,<sup>40</sup> M. Abolins,<sup>37</sup> V. Abramov,<sup>15</sup> B. S. Acharya,<sup>8</sup> I. Adam,<sup>39</sup> D. L. Adams,<sup>48</sup> M. Adams,<sup>24</sup> S. Ahn,<sup>23</sup> H. Aihara,<sup>17</sup> H. Alves,<sup>2</sup> N. Amos,<sup>36</sup> E. W. Anderson,<sup>30</sup> R. Astur,<sup>42</sup> M. M. Baarmand,<sup>42</sup> V. V. Babintsev,<sup>15</sup> L. Babukhadia,<sup>16</sup> A. Baden,<sup>33</sup> V. Balamurali,<sup>28</sup> B. Baldin,<sup>23</sup> S. Banerjee,<sup>8</sup> J. Bantly,<sup>45</sup> E. Barberis,<sup>17</sup> P. Baringer,<sup>31</sup> J. F. Bartlett,<sup>23</sup> A. Belyaev,<sup>14</sup> S. B. Beri,<sup>6</sup> I. Bertram,<sup>26</sup> V. A. Bezzubov,<sup>15</sup> P. C. Bhat,<sup>23</sup> V. Bhatnagar,<sup>6</sup> M. Bhattacharjee,<sup>42</sup> N. Biswas,<sup>28</sup> G. Blazey,<sup>25</sup> S. Blessing,<sup>21</sup> P. Bloom,<sup>18</sup> A. Boehnlein,<sup>23</sup> N. I. Bojko,<sup>15</sup> F. Borchering,<sup>23</sup> C. Boswell,<sup>20</sup> A. Brandt,<sup>23</sup> R. Breedon,<sup>18</sup> R. Brock,<sup>37</sup> A. Bross,<sup>23</sup> D. Buchholz,<sup>26</sup> V. S. Burtovoi,<sup>15</sup> J. M. Butler,<sup>34</sup> W. Carvalho,<sup>2</sup> D. Casey,<sup>37</sup> Z. Casilum,<sup>42</sup> H. Castilla-Valdez,<sup>11</sup> D. Chakraborty,<sup>42</sup> S.-M. Chang,<sup>35</sup> S. V. Chekulaev,<sup>15</sup> L.-P. Chen,<sup>17</sup> W. Chen,<sup>42</sup> S. Choi,<sup>10</sup> S. Chopra,<sup>36</sup> B. C. Choudhary,<sup>20</sup> J. H. Christenson,<sup>23</sup> M. Chung,<sup>24</sup> D. Claes,<sup>38</sup> A. R. Clark,<sup>17</sup> W. G. Cobau,<sup>33</sup> J. Cochran,<sup>20</sup> L. Coney,<sup>28</sup> W. E. Cooper,<sup>23</sup> C. Cretsinger,<sup>41</sup> D. Cullen-Vidal,<sup>45</sup> M. A. C. Cummings,<sup>25</sup> D. Cutts,<sup>45</sup> O. I. Dahl,<sup>17</sup> K. Davis,<sup>16</sup> K. De,<sup>46</sup> K. Del Signore,<sup>36</sup> M. Demarteau,<sup>23</sup> D. Denisov,<sup>23</sup> S. P. Denisov,<sup>15</sup> H. T. Diehl,<sup>23</sup> M. Diesburg,<sup>23</sup> G. Di Loreto,<sup>37</sup> P. Draper,<sup>46</sup> Y. Ducros,<sup>5</sup> L. V. Dudko,<sup>14</sup> S. R. Dugad,<sup>8</sup> A. Dyshkant,<sup>15</sup> D. Edmunds,<sup>37</sup> J. Ellison,<sup>20</sup> V. D. Elvira,<sup>42</sup> R. Engelmann,<sup>42</sup> S. Eno,<sup>33</sup> G. Eppley,<sup>48</sup> P. Ermolov,<sup>14</sup> O. V. Eroshin,<sup>15</sup> V. N. Evdokimov,<sup>15</sup> T. Fahland,<sup>19</sup> M. K. Fatyga,<sup>41</sup> S. Feher,<sup>23</sup> D. Fein,<sup>16</sup> T. Ferbel,<sup>41</sup> G. Finocchiaro,<sup>42</sup> H. E. Fisk,<sup>23</sup> Y. Fisyak,<sup>43</sup> E. Flattum,<sup>23</sup> G. E. Forden,<sup>16</sup> M. Fortner,<sup>25</sup> K. C. Frame,<sup>37</sup> S. Fuess,<sup>23</sup> E. Gallas,<sup>46</sup> A. N. Galyaev,<sup>15</sup> P. Gartung,<sup>20</sup> V. Gavrilov,<sup>13</sup> T. L. Geld,<sup>37</sup> R. J. Genik II,<sup>37</sup> K. Genser,<sup>23</sup> C. E. Gerber,<sup>23</sup> Y. Gershtein,<sup>13</sup> B. Gibbard,<sup>43</sup> B. Gobbi,<sup>26</sup> B. Gómez,<sup>4</sup> G. Gómez,<sup>33</sup> P. I. Goncharov,<sup>15</sup> J. L. González Solís,<sup>11</sup> H. Gordon,<sup>43</sup> L. T. Goss,<sup>47</sup> K. Gounder,<sup>20</sup> A. Goussiou,<sup>42</sup> N. Graf,<sup>43</sup> P. D. Grannis,<sup>42</sup> D. R. Green,<sup>23</sup> H. Greenlee,<sup>23</sup> S. Grinstein,<sup>1</sup> P. Grudberg,<sup>17</sup> S. Grünendahl,<sup>23</sup> G. Guglielmo,<sup>44</sup> J. A. Guida,<sup>16</sup> J. M. Guida,<sup>45</sup> A. Gupta,<sup>8</sup> S. N. Gurzhiev,<sup>15</sup> G. Gutierrez,<sup>23</sup> P. Gutierrez,<sup>44</sup> N. J. Hadley,<sup>33</sup> H. Haggerty,<sup>23</sup> S. Hagopian,<sup>21</sup> V. Hagopian,<sup>21</sup> K. S. Hahn,<sup>41</sup> R. E. Hall,<sup>19</sup> P. Hanlet,<sup>35</sup> S. Hansen,<sup>23</sup> J. M. Hauptman,<sup>30</sup> D. Hedin,<sup>25</sup> A. P. Heinson,<sup>20</sup> U. Heintz,<sup>23</sup> R. Hernández-Montoya,<sup>11</sup> T. Heuring,<sup>21</sup> R. Hirosky,<sup>24</sup> J. D. Hobbs,<sup>42</sup> B. Hoeneisen,<sup>4,\*</sup> J. S. Hoftun,<sup>45</sup> F. Hsieh,<sup>36</sup> Ting Hu,<sup>42</sup> Tong Hu,<sup>27</sup> T. Huehn,<sup>20</sup> A. S. Ito,<sup>23</sup> E. James,<sup>16</sup> J. Jaques,<sup>28</sup> S. A. Jeger,<sup>37</sup> R. Jesik,<sup>27</sup> T. Joffe-Minor,<sup>26</sup> K. Johns,<sup>16</sup> M. Johnson,<sup>23</sup> A. Jonckheere,<sup>23</sup> M. Jones,<sup>22</sup> H. Jöstlein,<sup>23</sup> S. Y. Jun,<sup>26</sup> C. K. Jung,<sup>42</sup> S. Kahn,<sup>43</sup> G. Kalbfleisch,<sup>44</sup> D. Karmanov,<sup>14</sup> D. Karmgard,<sup>21</sup> R. Kehoe,<sup>28</sup> M. L. Kelly,<sup>28</sup> S. K. Kim,<sup>10</sup> B. Klima,<sup>23</sup> C. Klopfenstein,<sup>18</sup> W. Ko,<sup>18</sup> J. M. Kohli,<sup>6</sup> D. Koltick,<sup>29</sup> A. V. Kostitskiy,<sup>15</sup> J. Kotcher,<sup>43</sup> A. V. Kotwal,<sup>39</sup> A. V. Kozelov,<sup>15</sup> E. A. Kozlovsky,<sup>15</sup> J. Krane,<sup>38</sup> M. R. Krishnaswamy,<sup>8</sup> S. Krzywdzinski,<sup>23</sup> S. Kuleshov,<sup>13</sup> S. Kunori,<sup>33</sup> F. Landry,<sup>37</sup> G. Landsberg,<sup>45</sup> B. Lauer,<sup>30</sup> A. Leflat,<sup>14</sup> J. Li,<sup>46</sup> Q. Z. Li-Demarteau,<sup>23</sup> J. G. R. Lima,<sup>3</sup> D. Lincoln,<sup>23</sup> S. L. Linn,<sup>21</sup> J. Linnemann,<sup>37</sup> R. Lipton,<sup>23</sup> F. Lobkowicz,<sup>41</sup> S. C. Loken,<sup>17</sup> A. Lucotte,<sup>42</sup> L. Lueking,<sup>23</sup> A. L. Lyon,<sup>33</sup> A. K. A. Maciel,<sup>2</sup> R. J. Madaras,<sup>17</sup> R. Madden,<sup>21</sup> L. Magaña-Mendoza,<sup>11</sup> V. Manankov,<sup>14</sup> S. Mani,<sup>18</sup> H. S. Mao,<sup>23,†</sup> R. Markeloff,<sup>25</sup> T. Marshall,<sup>27</sup> M. I. Martin,<sup>23</sup> K. M. Mauritz,<sup>30</sup> B. May,<sup>26</sup> A. A. Mayorov,<sup>15</sup> R. McCarthy,<sup>42</sup> J. McDonald,<sup>21</sup> T. McKibben,<sup>24</sup> J. McKinley,<sup>37</sup> T. McMahon,<sup>44</sup> H. L. Melanson,<sup>23</sup> M. Merkin,<sup>14</sup> K. W. Merritt,<sup>23</sup> C. Miao,<sup>45</sup> H. Miettinen,<sup>48</sup> A. Mincer,<sup>40</sup> C. S. Mishra,<sup>23</sup> N. Mokhov,<sup>23</sup> N. K. Mondal,<sup>8</sup> H. E. Montgomery,<sup>23</sup> P. Mooney,<sup>4</sup> M. Mostafa,<sup>1</sup> H. da Motta,<sup>2</sup> C. Murphy,<sup>24</sup> F. Nang,<sup>16</sup> M. Narain,<sup>23</sup> V. S. Narasimham,<sup>8</sup> A. Narayanan,<sup>16</sup> H. A. Neal,<sup>36</sup> J. P. Negret,<sup>4</sup> P. Nemethy,<sup>40</sup> D. Norman,<sup>47</sup> L. Oesch,<sup>36</sup> V. Oguri,<sup>3</sup> E. Oliveira,<sup>2</sup> E. Oltman,<sup>17</sup> N. Oshima,<sup>23</sup> D. Owen,<sup>37</sup> P. Padley,<sup>48</sup> A. Para,<sup>23</sup> Y. M. Park,<sup>9</sup> R. Partridge,<sup>45</sup> N. Parua,<sup>8</sup> M. Paterno,<sup>41</sup> B. Pawlik,<sup>12</sup> J. Perkins,<sup>46</sup> M. Peters,<sup>22</sup> R. Piegai,<sup>1</sup> H. Piekarczyk,<sup>21</sup> Y. Pischalnikov,<sup>29</sup> B. G. Pope,<sup>37</sup> H. B. Prosper,<sup>21</sup> S. Protopopescu,<sup>43</sup> J. Qian,<sup>36</sup> P. Z. Quintas,<sup>23</sup> R. Raja,<sup>23</sup> S. Rajagopalan,<sup>43</sup> O. Ramirez,<sup>24</sup> S. Reucroft,<sup>35</sup> M. Rijssenbeek,<sup>42</sup> T. Rockwell,<sup>37</sup> M. Roco,<sup>23</sup> P. Rubinov,<sup>26</sup> R. Ruchti,<sup>28</sup> J. Rutherford,<sup>16</sup> A. Sánchez-Hernández,<sup>11</sup> A. Santoro,<sup>2</sup> L. Sawyer,<sup>32</sup> R. D. Schamberger,<sup>42</sup> H. Schellman,<sup>26</sup> J. Sculli,<sup>40</sup> E. Shabalina,<sup>14</sup> C. Shaffer,<sup>21</sup> H. C. Shankar,<sup>8</sup> R. K. Shivpuri,<sup>7</sup> M. Shupe,<sup>16</sup> H. Singh,<sup>20</sup> J. B. Singh,<sup>6</sup> V. Sirotenko,<sup>25</sup> E. Smith,<sup>44</sup> R. P. Smith,<sup>23</sup> R. Snihur,<sup>26</sup> G. R. Snow,<sup>38</sup> J. Snow,<sup>44</sup> S. Snyder,<sup>43</sup> J. Solomon,<sup>24</sup> M. Sosebee,<sup>46</sup> N. Sotnikova,<sup>14</sup> M. Souza,<sup>2</sup> A. L. Spadafora,<sup>17</sup> G. Steinbrück,<sup>44</sup> R. W. Stephens,<sup>46</sup> M. L. Stevenson,<sup>17</sup> D. Stewart,<sup>36</sup> F. Stichelbaut,<sup>42</sup> D. Stoker,<sup>19</sup> V. Stolin,<sup>13</sup> D. A. Stoyanova,<sup>15</sup> M. Strauss,<sup>44</sup> K. Streets,<sup>40</sup> M. Strovink,<sup>17</sup> A. Sznajder,<sup>2</sup> P. Tamburello,<sup>33</sup> J. Tarazi,<sup>19</sup> M. Tartaglia,<sup>23</sup> T. L. T. Thomas,<sup>26</sup> J. Thompson,<sup>33</sup> T. G. Trippe,<sup>17</sup> P. M. Tuts,<sup>39</sup> V. Vaniev,<sup>15</sup> N. Varelas,<sup>24</sup> E. W. Varnes,<sup>17</sup> D. Vititoe,<sup>16</sup> A. A. Volkov,<sup>15</sup> A. P. Vorobiev,<sup>15</sup> H. D. Wahl,<sup>21</sup> G. Wang,<sup>21</sup> J. Warchol,<sup>28</sup> G. Watts,<sup>45</sup> M. Wayne,<sup>28</sup> H. Weerts,<sup>37</sup> A. White,<sup>46</sup> J. T. White,<sup>47</sup> J. A. Wightman,<sup>30</sup> S. Willis,<sup>25</sup> S. J. Wimpenny,<sup>20</sup> J. V. D. Wirjawan,<sup>47</sup> J. Womersley,<sup>23</sup> E. Won,<sup>41</sup> D. R. Wood,<sup>35</sup> Z. Wu,<sup>23,†</sup> H. Xu,<sup>45</sup> R. Yamada,<sup>23</sup> P. Yamin,<sup>43</sup> T. Yasuda,<sup>35</sup> P. Yepes,<sup>48</sup> K. Yip,<sup>23</sup> C. Yoshikawa,<sup>22</sup> S. Youssef,<sup>21</sup> J. Yu,<sup>23</sup> Y. Yu,<sup>10</sup> B. Zhang,<sup>23,†</sup> Y. Zhou,<sup>23,†</sup> Z. Zhou,<sup>30</sup> Z. H. Zhu,<sup>41</sup> M. Zielinski,<sup>41</sup> D. Zieminska,<sup>27</sup> A. Zieminski,<sup>27</sup> E. G. Zverev,<sup>14</sup> and A. Zylberstejn<sup>5</sup>

## (D0 Collaboration)

- <sup>1</sup>Universidad de Buenos Aires, Buenos Aires, Argentina  
<sup>2</sup>LAFEX, Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, Brazil  
<sup>3</sup>Universidade do Estado do Rio de Janeiro, Rio de Janeiro, Brazil  
<sup>4</sup>Universidad de los Andes, Bogotá, Colombia  
<sup>5</sup>DAPNIA/Service de Physique des Particules, CEA, Saclay, France  
<sup>6</sup>Panjab University, Chandigarh, India  
<sup>7</sup>Delhi University, Delhi, India  
<sup>8</sup>Tata Institute of Fundamental Research, Mumbai, India  
<sup>9</sup>Kyungsung University, Pusan, Korea  
<sup>10</sup>Seoul National University, Seoul, Korea  
<sup>11</sup>CINVESTAV, Mexico City, Mexico  
<sup>12</sup>Institute of Nuclear Physics, Kraków, Poland  
<sup>13</sup>Institute for Theoretical and Experimental Physics, Moscow, Russia  
<sup>14</sup>Moscow State University, Moscow, Russia  
<sup>15</sup>Institute for High Energy Physics, Protvino, Russia  
<sup>16</sup>University of Arizona, Tucson, Arizona 85721  
<sup>17</sup>Lawrence Berkeley National Laboratory and University of California, Berkeley, California 94720  
<sup>18</sup>University of California, Davis, California 95616  
<sup>19</sup>University of California, Irvine, California 92697  
<sup>20</sup>University of California, Riverside, California 92521  
<sup>21</sup>Florida State University, Tallahassee, Florida 32306  
<sup>22</sup>University of Hawaii, Honolulu, Hawaii 96822  
<sup>23</sup>Fermi National Accelerator Laboratory, Batavia, Illinois 60510  
<sup>24</sup>University of Illinois at Chicago, Chicago, Illinois 60607  
<sup>25</sup>Northern Illinois University, DeKalb, Illinois 60115  
<sup>26</sup>Northwestern University, Evanston, Illinois 60208  
<sup>27</sup>Indiana University, Bloomington, Indiana 47405  
<sup>28</sup>University of Notre Dame, Notre Dame, Indiana 46556  
<sup>29</sup>Purdue University, West Lafayette, Indiana 47907  
<sup>30</sup>Iowa State University, Ames, Iowa 50011  
<sup>31</sup>University of Kansas, Lawrence, Kansas 66045  
<sup>32</sup>Louisiana Tech University, Ruston, Louisiana 71272  
<sup>33</sup>University of Maryland, College Park, Maryland 20742  
<sup>34</sup>Boston University, Boston, Massachusetts 02215  
<sup>35</sup>Northeastern University, Boston, Massachusetts 02115  
<sup>36</sup>University of Michigan, Ann Arbor, Michigan 48109  
<sup>37</sup>Michigan State University, East Lansing, Michigan 48824  
<sup>38</sup>University of Nebraska, Lincoln, Nebraska 68588  
<sup>39</sup>Columbia University, New York, New York 10027  
<sup>40</sup>New York University, New York, New York 10003  
<sup>41</sup>University of Rochester, Rochester, New York 14627  
<sup>42</sup>State University of New York, Stony Brook, New York 11794  
<sup>43</sup>Brookhaven National Laboratory, Upton, New York 11973  
<sup>44</sup>University of Oklahoma, Norman, Oklahoma 73019  
<sup>45</sup>Brown University, Providence, Rhode Island 02912  
<sup>46</sup>University of Texas, Arlington, Texas 76019  
<sup>47</sup>Texas A&M University, College Station, Texas 77843  
<sup>48</sup>Rice University, Houston, Texas 77005

(Received 17 July 1998)

We have made a precise measurement of the central inclusive jet cross section at  $\sqrt{s} = 1.8$  TeV. The measurement is based on an integrated luminosity of  $92 \text{ pb}^{-1}$  collected at the Fermilab Tevatron  $\bar{p}p$  Collider with the D0 detector. The cross section, reported as a function of jet transverse energy ( $E_T \geq 60$  GeV) in the pseudorapidity interval  $|\eta| \leq 0.5$ , is in good agreement with predictions from next-to-leading order quantum chromodynamics. [S0031-9007(99)08800-6]

PACS numbers: 13.87.Ce, 12.38.Qk

Within the framework of quantum chromodynamics (QCD), inelastic scattering between a proton and an antiproton can be described as an elastic collision between

a single proton constituent and a single antiproton constituent. These constituents are often referred to as partons. After the collision, the outgoing partons manifest

themselves as localized streams of particles or “jets.” Predictions for the inclusive jet cross section are given by the folding of parton scattering cross sections with experimentally determined parton distribution functions (pdf’s). These predictions have recently improved with next-to-leading order (NLO) QCD scattering calculations [1–3] and new, accurately measured pdf’s [4,5]. We measure the cross section for the production of jets as a function of the jet energy in the plane transverse to the incident beams,  $E_T$ . The measurement is based on an integrated luminosity of  $92 \text{ pb}^{-1}$  [6] of  $\bar{p}p$  collisions collected with the D0 detector [7] at the Fermilab Tevatron Collider. Measurements of inclusive jet production with smaller integrated luminosity have been performed previously by the UA2 and CDF Collaborations [8,9]. The cross section measurement presented here allows a stringent test of QCD, with a total uncertainty substantially reduced relative to previous results.

Jet detection in the D0 detector utilizes primarily the uranium-liquid argon calorimeters which have full coverage for pseudorapidity  $|\eta| \leq 4.1$  ( $\eta = -\ln[\tan(\theta/2)]$ , where  $\theta$  is the polar angle relative to the proton beam). Initial event selection occurred in two hardware trigger stages and a software stage. The first hardware trigger selected an inelastic  $\bar{p}p$  collision—indicated by signals from the trigger hodoscopes located near the beams on either side of the interaction region. The next stage required transverse energy above a preset threshold in calorimeter trigger tiles of  $\Delta\eta \times \Delta\phi = 0.8 \times 1.6$ , where  $\phi$  is the azimuthal angle. Selected events were digitized and sent to an array of processors. Jet candidates were then reconstructed with a cone algorithm and the entire event recorded if any jet  $E_T$  exceeded a specified threshold. For software jet thresholds of 30, 50, 85, and 115 GeV, integrated luminosities of 0.34, 4.6, 55, and  $92 \text{ pb}^{-1}$ , respectively, were accumulated in a 1994–1995 data run.

Jets were reconstructed off-line using an iterative fixed-cone algorithm with a cone radius of  $\mathcal{R} = 0.7$  in  $\eta$ - $\phi$  space [10]. Background from isolated noisy calorimeter cells and accelerator beam losses which mimicked jets were eliminated with quality cuts [11]. Background events from cosmic ray bremsstrahlung or misvertexed events were eliminated by requiring the missing transverse energy in each event to be less than the larger of 30 GeV or  $0.3E_T^{\text{max}}$ , where  $E_T^{\text{max}}$  is the  $E_T$  of the leading jet. Residual jet contamination is less than 1% at all  $E_T$ , based on event simulations with superimposed calorimeter noise distributions and on visual scanning of jet candidates with  $E_T$  greater than 350 GeV. The jet selection efficiency for  $|\eta| \leq 0.7$  has been measured as a function of jet  $E_T$  and found to be  $(97 \pm 1)\%$  below 250 GeV and decreasing smoothly to  $(95 \pm 2)\%$  at 400 GeV.

At high instantaneous luminosity, more than one interaction in a single beam crossing is probable ( $\sim 20\%$  for this data set). The event vertex was reconstructed using data from the central tracking system. For events with

multiple vertices, the two vertices with the largest number of tracks were retained. Because of the fluctuations of jet charged-particle multiplicity, an additional parameter was used to select the vertex. If an event had more than one vertex, the quantity  $S_T = |\sum \vec{E}_T^{\text{jet}}|$  was calculated for both vertices. The vertex with the smaller  $S_T$  was selected as the event vertex and used to calculate jet  $E_T$  and  $\eta$ . The selected vertex was required to be within 50 cm of the detector center. This last requirement retained  $(90 \pm 1)\%$  of the events, independent of jet  $E_T$ .

The transverse energy of each jet was corrected for the underlying event, additional interactions, noise from uranium decay, the fraction of particle energy showered outside of the reconstruction cone, detector uniformity, and detector hadronic response. A complete discussion of the jet energy scale calibration can be found in Ref. [12]. For  $|\eta| \leq 0.5$ , the mean total correction factor for jet  $E_T$  is  $1.154 \pm 0.017$  [ $1.118 \pm 0.023$ ] at 100 GeV [400 GeV].

The inclusive jet cross section was computed in contiguous  $E_T$  ranges using data from the four trigger sets. The spectrum includes data from the 30 GeV trigger between 60 and 90 GeV, from the 50 GeV  $E_T$  trigger between 90 and 130 GeV, from the 85 GeV trigger between 130 and 170 GeV, and above 170 GeV from the 115 GeV trigger. A single interaction (per beam crossing) requirement on the two lowest- $E_T$  triggers introduced an inefficiency corrected by matching the 50 GeV trigger cross section to the 85 GeV trigger cross section above 130 GeV, where both triggers are fully efficient. This introduces an additional 1.1% luminosity uncertainty to the 50 GeV trigger set. A similar matching between the lowest- $E_T$  trigger and the 50 GeV trigger introduces another 1.4% uncertainty for the lower set, which is added in quadrature to the 1.1% matching uncertainty.

The steep  $E_T$  spectrum is distorted by jet energy resolution. At all  $E_T$ , the resolution (measured by balancing  $E_T$  in jet events) is well described by a Gaussian distribution; at 100 GeV the standard deviation is 7 GeV. The distortion was corrected by assuming an ansatz function  $(AE_T^{-B})(1 - 2E_T/\sqrt{s})^C$ , smearing it with the measured resolution and comparing the smeared result with the measured cross section. The procedure was repeated by varying parameters  $A$ ,  $B$ , and  $C$  until the best fit was found between the observed cross section and the smeared trial spectrum. The ratio of the initial ansatz to the smeared ansatz was used to correct the cross section on a bin-by-bin basis [13]. The resolution correction reduces the observed cross section by  $(13 \pm 3)\%$  [ $(8 \pm 2)\%$ ] at 60 GeV [400 GeV].

The resulting inclusive jet cross section for  $|\eta| \leq 0.5$ , shown in Fig. 1, has been averaged over each  $E_T$  bin ( $\Delta E_T$ ) and over the central unit of rapidity ( $\Delta\eta = 1$ ). This bin-averaged double differential cross section,  $\langle d^2\sigma/(dE_T d\eta) \rangle$ , was calculated as  $NC/(\Delta E_T \Delta\eta \epsilon \mathcal{L})$  where  $N$  is the total number of jets observed in a bin,

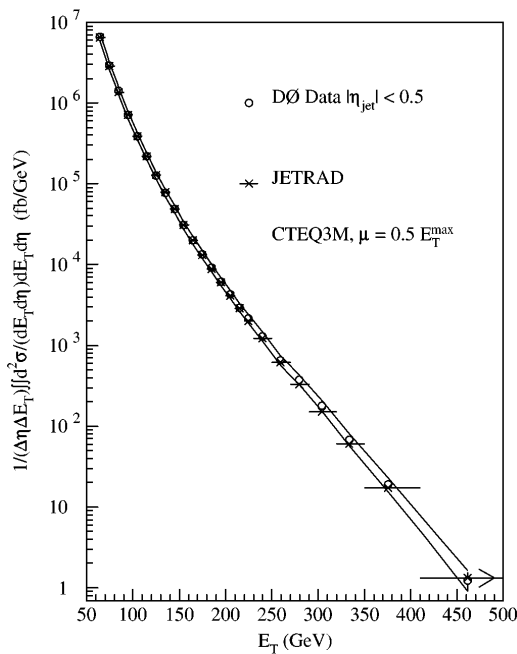


FIG. 1. The  $|\eta| \leq 0.5$  inclusive cross section. Statistical uncertainties are invisible on this scale. The solid curves represent the  $\pm 1\sigma$  systematic uncertainty band on the data.

$C$  the smearing correction,  $\epsilon$  the selection efficiency, and  $\mathcal{L}$  the integrated luminosity associated with the trigger set. The cross section is consistent with a preliminary measurement from a smaller 1992–1993 data set [11].

Figure 1 also shows a theoretical prediction for the cross section from the NLO event generator JETRAD [3]. There is good agreement over 7 orders of magnitude. Inputs to the NLO calculation are the renormalization scale  $\mu$  (equal to the factorization scale), the pdf, and the parton clustering algorithm. For the calculation shown here,  $\mu = 0.5E_T^{\max}$  and the pdf is CTEQ3M [4]. Partons separated by less than  $R_{\text{sep}} = 1.3R$  were clustered if they were also within  $R = 0.7$  of their  $E_T$ -weighted  $\eta$ - $\phi$  centroid. This choice of  $R_{\text{sep}}$  is discussed in Ref. [10]. Variations in the predicted cross section due to the input choices are about 30% [14].

The data in Fig. 1 have an overall luminosity uncertainty of 6.1%, and are plotted at the  $E_T$  value for which a smooth function describing the cross section is equal to the average cross section in each bin. The band shows the total systematic uncertainty as a function of  $E_T$ . Listed in Table I are the plotted values of  $E_T$ ,  $E_T$  ranges, cross section, and statistical and systematic uncertainty. The systematic uncertainties include jet and event selection, unsmearing, relative luminosity, and energy scale uncertainties added in quadrature. The 6.1% luminosity uncertainty is not included.

Figure 2 shows the various uncertainties for the  $|\eta| \leq 0.5$  cross section. Each curve represents the average of the nearly symmetric upper and lower uncertainties. The energy scale uncertainty varies from 8% at low  $E_T$  to

TABLE I. The  $|\eta| < 0.5$  cross section (overall luminosity uncertainty not included).

Plotted $E_T$ (GeV)	Bin range (GeV)	Cross sec. $\pm$ stat. (fb/GeV)	Syst. Uncer. (%)
64.6	60–70	$(6.59 \pm 0.04) \times 10^6$	$\pm 8$
74.6	70–80	$(2.89 \pm 0.03) \times 10^6$	$\pm 8$
84.7	80–90	$(1.41 \pm 0.02) \times 10^6$	$\pm 8$
94.7	90–100	$(7.07 \pm 0.04) \times 10^5$	$\pm 8$
104.7	100–110	$(3.88 \pm 0.03) \times 10^5$	$\pm 8$
114.8	110–120	$(2.21 \pm 0.02) \times 10^5$	$\pm 8$
124.8	120–130	$(1.27 \pm 0.02) \times 10^5$	$\pm 8$
134.8	130–140	$(7.70 \pm 0.04) \times 10^4$	$\pm 8$
144.8	140–150	$(4.86 \pm 0.03) \times 10^4$	$\pm 8$
154.8	150–160	$(3.07 \pm 0.02) \times 10^4$	+9, -8
164.8	160–170	$(2.00 \pm 0.02) \times 10^4$	$\pm 9$
174.8	170–180	$(1.34 \pm 0.01) \times 10^4$	$\pm 9$
184.8	180–190	$(9.12 \pm 0.10) \times 10^3$	$\pm 9$
194.8	190–200	$(6.15 \pm 0.09) \times 10^3$	+10, -9
204.8	200–210	$(4.29 \pm 0.07) \times 10^3$	$\pm 10$
214.8	210–220	$(2.93 \pm 0.06) \times 10^3$	+11, -10
224.8	220–230	$(2.14 \pm 0.05) \times 10^3$	+11, -10
239.4	230–250	$(1.30 \pm 0.03) \times 10^3$	$\pm 11$
259.4	250–270	$(6.54 \pm 0.20) \times 10^2$	+12, -11
279.5	270–290	$(3.77 \pm 0.15) \times 10^2$	+13, -12
303.9	290–320	$(1.79 \pm 0.08) \times 10^2$	+15, -13
333.9	320–350	$(6.82 \pm 0.52) \times 10^1$	+17, -15
375.5	350–410	$(1.89 \pm 0.19) \times 10^1$	+20, -17
461.1	410–560	$(1.24 \pm 0.31) \times 10^0$	+30, -26

30% at 450 GeV. This contribution dominates all other sources of uncertainty, except at low  $E_T$ , where the 6.1% luminosity uncertainty is of a comparable magnitude.

The  $|\eta| \leq 0.5$  region provides our optimum test for departures of data from NLO QCD. In this region, the detector is uniformly thick (seven or more interaction lengths with no gaps) and both jet resolution and calibration are precise. Also, jet production from the scattering of possible constituents within quarks is largest for  $\eta = 0$ , relative to standard QCD predictions [15]. For comparison to Ref. [9], we have also carried out a similar

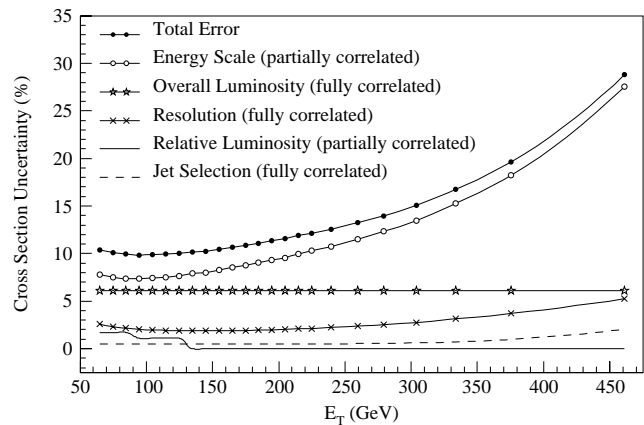


FIG. 2. Contributions to the  $|\eta| \leq 0.5$  cross section uncertainty plotted by component.

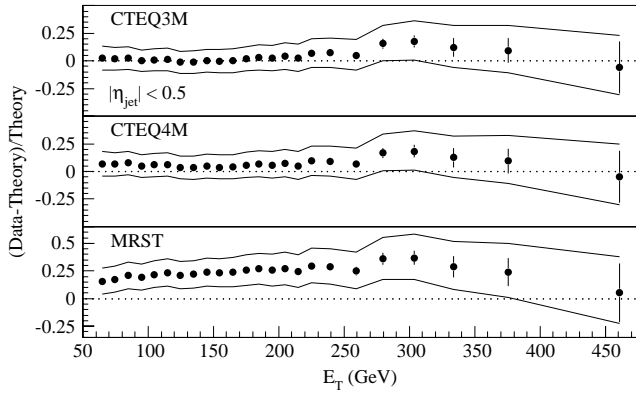


FIG. 3. The difference between data and JETRAD QCD predictions normalized to predictions. The bands represent the total experimental uncertainty.

analysis for  $0.1 \leq |\eta| \leq 0.7$ . Figure 3 shows the ratios  $(D - T)/T$  for the data ( $D$ ) and JETRAD NLO theoretical ( $T$ ) predictions based on the CTEQ3M, CTEQ4M, and MRST pdf's [4,5] for  $|\eta| \leq 0.5$ . Given the experimental and theoretical uncertainties, the predictions are in agreement with the data; in particular, the data above 350 GeV show no indication of an excess relative to QCD.

The data and theory can be compared quantitatively with a  $\chi^2$  test incorporating the uncertainty covariance matrix. The matrix elements are constructed from the statistical and systematic uncertainties and by analyzing the mutual correlation of the uncertainties in Fig. 2 at each pair of  $E_T$  values. As indicated by the figure the overall systematic uncertainty is highly correlated from bin to bin. Table II shows that the bin-to-bin correlations in the full uncertainty for representative  $E_T$  bins are greater than 40% and positive. (The full matrix can be found in Ref. [16].)

Table III lists  $\chi^2$  values for several JETRAD predictions incorporating various parton distribution functions [4,5]. Each comparison has 24 degrees of freedom. The JETRAD predictions have been fit to a smooth function of  $E_T$ . All five predictions describe the  $|\eta| \leq 0.5$  cross section very well (the probabilities for  $\chi^2$  to exceed the listed values are between 47% and 90%). The  $0.1 \leq |\eta| \leq 0.7$  cross section is also well described (probabilities between 24% and 72%). We have also made comparisons between the  $|\eta| \leq 0.5$  data and Ellis-Kunszt-Soper (EKS) [1] calculations using CTEQ3M,  $\mathcal{R}_{\text{sep}} = 1.3\mathcal{R}$ , and with renormalization scales  $\mu = 0.25E_T^{\text{max}}$ ,  $0.50E_T^{\text{max}}$ , and  $1.00E_T^{\text{max}}$

TABLE II. Cross section total uncertainty correlations.

$E_T$ (GeV)	64.6	104.7	204.8	303.9	461.1
64.6	1.00	0.96	0.85	0.71	0.40
104.7	0.96	1.00	0.92	0.79	0.46
204.8	0.85	0.92	1.00	0.91	0.61
303.9	0.71	0.79	0.91	1.00	0.67
461.1	0.40	0.46	0.61	0.67	1.00

TABLE III.  $\chi^2$  comparisons between JETRAD and  $|\eta| \leq 0.5$  and  $0.1 \leq |\eta| \leq 0.7$  data for  $\mu = 0.5E_T^{\text{max}}$ ,  $\mathcal{R}_{\text{sep}} = 1.3\mathcal{R}$  and various pdfs. There are 24 degrees of freedom.

pdf	$ \eta  \leq 0.5$	$0.1 \leq  \eta  \leq 0.7$
CTEQ3M	23.9	28.4
CTEQ4M	17.6	23.3
CTEQ4HJ	15.7	20.5
MRSA'	20.0	27.8
MRST	17.0	19.5

and  $\mu = 0.25E_T^{\text{jet}}$ ,  $0.50E_T^{\text{jet}}$ , and  $1.00E_T^{\text{jet}}$ . These calculations also describe the data very well (better than 57% probability) at all renormalization scales.

The top panel in Fig. 4 shows  $(D - T)/T$  for our data in the  $0.1 \leq |\eta| \leq 0.7$  region relative to an EKS calculation using the CTEQ3M pdf,  $\mu = 0.5E_T^{\text{jet}}$ , and  $\mathcal{R}_{\text{sep}} = 2.0\mathcal{R}$ . (The tabulated data can be found in Ref. [16].) Also shown are the data of Ref. [9] relative to the same EKS prediction. For this rapidity region, we have carried out a  $\chi^2$  comparison between our data and the nominal curve describing the central values of the data of Ref. [9]. Comparing our data to the nominal curve, as though it were theory, we obtain a  $\chi^2$  of 63.2 for 24 degrees of freedom (probability of 0.002%). Thus our data cannot be described with this parametrization. As illustrated in the bottom panel of Fig. 4, our data and the curve differ at low and high  $E_T$ ; such differences cannot be accommodated by the highly correlated uncertainties of our data. If we include the systematic uncertainties of the data of Ref. [9] in the covariance matrix, the  $\chi^2$  is reduced to 24.7 (probability of 42%).

In conclusion, we have made the most precise measurement to date of the inclusive jet cross section for  $E_T \geq 60$  GeV. QCD predictions are in good agreement with the observed cross section for standard parton distribution functions and different renormalization scales. This is consistent with our previous measurements of dijet angular distributions [15], which are also in good

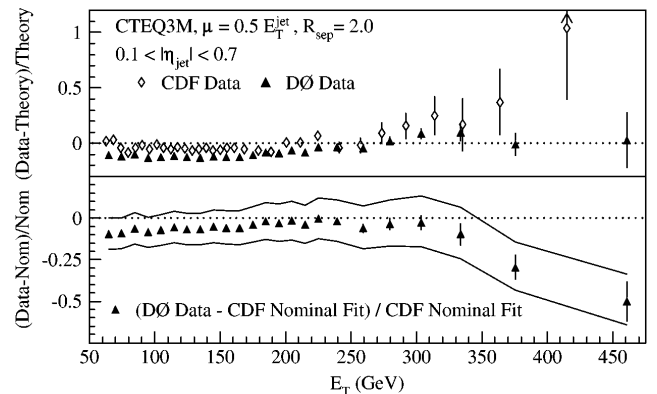


FIG. 4. Top: Comparisons of our data to EKS and of the data in Ref. [9] to EKS. See text for details. Bottom: Our data minus smoothed results of Ref. [9] divided by the latter. The band represents the uncertainty on our data.

agreement with QCD and show no indication of physics beyond the standard model.

We thank the staffs at Fermilab and collaborating institutions for their contributions to this work, and acknowledge support from the Department of Energy and National Science Foundation (U.S.A.), Commissariat à L'Energie Atomique (France), Ministry for Science and Technology and Ministry for Atomic Energy (Russia), CAPES and CNPq (Brazil), Departments of Atomic Energy and Science and Education (India), Colciencias (Colombia), CONACyT (Mexico), Ministry of Education and KOSEF (Korea), and CONICET and UBACyT (Argentina). We thank W. T. Giele, E. W. N. Glover, and D. A. Kosower for help with JETRAD.

---

\*Visitor from Universidad San Francisco de Quito, Quito, Ecuador.

†Visitor from IHEP, Beijing, China.

- [1] S. D. Ellis, Z. Kunszt, and D. E. Soper, *Phys. Rev. Lett.* **64**, 2121 (1990).
- [2] F. Aversa *et al.*, *Phys. Rev. Lett.* **65**, 401 (1990).
- [3] W. T. Giele, E. W. N. Glover, and D. A. Kosower, *Phys. Rev. Lett.* **73**, 2019 (1994).
- [4] H. L. Lai *et al.*, *Phys. Rev. D* **51**, 4763 (1995); **55**, 1280 (1997).
- [5] A. D. Martin *et al.*, *Eur. Phys. J. C* **4**, 463 (1998).
- [6] J. Bantly *et al.*, Report No. Fermilab-TM-1998, 1997.
- [7] D0 Collaboration, S. Abachi *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **338**, 185 (1994).
- [8] UA2 Collaboration, J. Alitti *et al.*, *Phys. Lett. B* **257**, 232 (1991); *Z. Phys. C* **49**, 17 (1991); CDF Collaboration, F. Abe *et al.*, *Phys. Rev. Lett.* **70**, 1376 (1993).
- [9] CDF Collaboration, F. Abe *et al.*, *Phys. Rev. Lett.* **77**, 438 (1996).
- [10] B. Abbott *et al.*, Report No. Fermilab-Pub-97/242-E, 1997.
- [11] V. D. Elvira, Ph.D. thesis, Universidad de Buenos Aires, 1995 (unpublished).
- [12] D0 Collaboration, B. Abbott *et al.*, Report No. Fermilab-Pub-97/330-E, 1998 (to be published).
- [13] M. Bhattacharjee, Ph.D. thesis, University of Delhi, 1996 (unpublished).
- [14] B. Abbott *et al.*, *Eur. Phys. J. C* **5**, 687 (1998).
- [15] D0 Collaboration, B. Abbott *et al.*, *Phys. Rev. Lett.* **80**, 666 (1998).
- [16] See AIP Document No. E-PAPS: E-PRLTAO-82-047913 for files with cross sections. E-PAPS document files may be retrieved free of charge from our FTP server (<http://www.aip.org/pubservs/paps.html>) or from <ftp.aip.org> in the director /epaps/. For further information, email: [paps@aip.org](mailto:paps@aip.org); or fax: 516-576-2223.