

# ExoMol line lists – XLVIII. High-temperature line list of thioformaldehyde (H<sub>2</sub>CS)

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## ABSTRACT

A comprehensive rotation–vibration (ro–vibrational) line list of thioformaldehyde (<sup>1</sup>H<sub>2</sub><sup>12</sup>C<sup>32</sup>S) that is applicable for elevated temperatures (2000 K) is presented. The new MOTY line list covers the 0–8000 cm<sup>−1</sup> range (wavelengths  $\lambda > 1.3 \mu\text{m}$ ) and contains around 43.5 billion transitions between 52.3 million states with rotational excitation up to  $J = 120$ . Line list calculations utilize a newly determined empirically refined potential energy surface (PES) – the most accurate H<sub>2</sub>CS PES to date – a previously published high-level *ab initio* dipole moment surface, and the use of an exact kinetic energy operator for solving the ro–vibrational Schrödinger equation. Post-processing of the MOTY line list is performed by replacing calculated energy levels with empirically derived values, vastly improving the accuracy of predicted line positions in certain spectral windows and making the line list suitable for high-resolution applications. The MOTY line list is available from the ExoMol data base at [www.exomol.com](http://www.exomol.com) and the CDS astronomical data base.

**Key words:** molecular data – opacity – planets and satellites: atmospheres – stars: atmospheres – ISM: molecules.

## 1 INTRODUCTION

Since its original interstellar detection (Sinclair et al. 1973), thioformaldehyde (main isotopologue <sup>1</sup>H<sub>2</sub><sup>12</sup>C<sup>32</sup>S, henceforth referred to as H<sub>2</sub>CS) has been observed in a variety of astronomical environments. For example, in massive star-forming regions such as Sagittarius B2 (Gardner, Whiteoak & Höglund 1980), where it can be efficiently formed through bimolecular reactions (Doddipatla et al. 2020), in galaxies which are nearby (Heikkilä, Johansson & Olofsson 1999; Martín et al. 2005) and more distant (Muller et al. 2011), molecular clouds (Esplugues et al. 2022), and in the comet Hale–Bopp (Woodney et al. 1997), which was the first detection of thioformaldehyde in a comet. Given that it is a simple organosulphur molecule, thioformaldehyde can be expected in the atmospheres of exoplanets, where sulphur chemistry is known to play a significant role in atmospheric composition (He et al. 2020; Hobbs et al. 2021). The need for accurate and complete molecular opacity data of H<sub>2</sub>CS is a major driving force behind this work and the production of a new hot line list for the ExoMol data base (Tennyson et al. 2020).

Of interest here is the infrared (IR) rotation–vibration (ro–vibrational) spectrum of H<sub>2</sub>CS, which has been the subject of numerous computational and experimental studies. Very recently, we performed a comprehensive analysis of the published experimental spectroscopic literature of H<sub>2</sub>CS (Mellor et al. 2023) and extracted all meaningful transition data from 11 literature sources (Johnson & Powell 1970; Johns & Olson 1971; Beers et al. 1972; Fabricant, Krieger & Muentner 1977; Bedwell & Duxbury 1980; Turner, Halonen & Mills 1981; Mcnaughton & Bruget 1993; Clouthier et al. 1994;

Flaud et al. 2008; Maeda et al. 2008; Müller et al. 2019). These data covered the  $\nu_1$ ,  $\nu_2$ ,  $\nu_3$ ,  $\nu_4$ ,  $\nu_5$ ,  $\nu_6$ , and  $2\nu_2$  vibrational bands of H<sub>2</sub>CS and included transitions with rotational excitation up to  $J = 54$ , where  $J$  is the total angular momentum quantum number. Using the robust MARVEL (Measured Active Rotational–Vibrational Energy Levels) procedure (Császár et al. 2007; Furtenbacher, Császár & Tennyson 2007; Furtenbacher & Császár 2012; Tóbiás et al. 2019), we were able to validate 11 638 transitions to produce a highly accurate set of 4254 ro–vibrational energy levels, all labelled with unique quantum numbers and uncertainties. Such a data set of energy levels has been instrumental in producing a new line list of thioformaldehyde that is suitable for studying exoplanet atmospheres at high spectral resolution, where there are considerable demands on the accuracy of the line positions (Snellen 2014; Birkby 2018).

A previous room temperature line list of H<sub>2</sub>CS was generated by Yachmenev, Polyak & Thiel (2013) and the computations utilized a highly accurate potential energy surface (PES) and dipole moment surface (DMS) constructed using state-of-the-art *ab initio* theory (Yachmenev et al. 2011). Just under 548 million transitions for ro–vibrational states up to  $J = 30$  with band origins below 5000 cm<sup>−1</sup> were computed using the variational nuclear motion code TROVE (Yurchenko, Thiel & Jensen 2007). Because the PES utilized was essentially *ab initio*, the accuracy of the predicted fundamental wavenumbers was limited to around 1–2 cm<sup>−1</sup>, which is not sufficient for spectroscopic applications. To substantially improve the accuracy of the computed line positions, it is necessary to refine the PES to a set of empirical-quality energy levels such as the MARVEL H<sub>2</sub>CS data set (Mellor et al. 2023) discussed above. Doing so can lead to orders-of-magnitude improvements in the accuracy of the transition wavenumbers and this procedure has been undertaken for this work. It is not necessary to perform any kind of

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refinement to high-quality *ab initio* DMSs such as the H<sub>2</sub>CS DMS of Yachmenev et al. (2013). Experience has shown that transition intensities computed using *ab initio* DMSs are often comparable to, and occasionally more reliable than experiment (Tennyson 2014; Yurchenko 2014; Bielska et al. 2022).

In this work, we present a newly computed comprehensive line list for H<sub>2</sub>CS, named MOTY, that is suitable for elevated temperatures up to  $T = 2000$  K. Although, using a very similar computational set-up to the room temperature line list calculations of Yachmenev et al. (2013), we have made two major developments to improve the accuracy and spectroscopic coverage of the new MOTY line list. The most significant improvement is the use of an empirically refined PES, the details of which are discussed in Section 2.3. Secondly, our variational ro–vibrational calculations of H<sub>2</sub>CS employed an exact kinetic energy operator (KEO), which considerably improves convergence making the computations more tractable, thus allowing for greater energy level coverage and improved accuracy. The use of an exact KEO is the first application of its kind in TROVE to a polyatomic molecule and the details are reported in Section 2.1. Tailoring the MOTY line list towards high-resolution applications is detailed in Section 4.1, where we follow the ExoMolHD strategy (Yurchenko & Tennyson 2021) by replacing calculated energy levels with the empirically derived MARVEL values if available. Results and simulated spectra are presented in Section 5 and we conclude in Section 6.

## 2 VARIATIONAL NUCLEAR MOTION CALCULATIONS

Variational calculations used the nuclear motion programme TROVE, whose general methodology is well documented (Yurchenko et al. 2007, 2009; Yachmenev & Yurchenko 2015; Tennyson & Yurchenko 2017; Yurchenko, Yachmenev & Ovsyannikov 2017). TROVE calculations on H<sub>2</sub>CS have been carried out previously when generating a room temperature line list (Yachmenev et al. 2013) and for validating a high-level *ab initio* PES of H<sub>2</sub>CS (Yachmenev et al. 2011). We therefore summarize only the key steps in the calculation process and expand on the aspects that are novel to this work, namely the empirically refined PES and the use of an exact KEO. Generally speaking, ro–vibrational calculations were performed in the range  $J = 0–120$  for each of the four irreducible representations  $A_1$ ,  $A_2$ ,  $B_1$ , and  $B_2$  of the  $C_{2v}(M)$  molecular symmetry group (Bunker & Jensen 1998) to obtain ro–vibrational energies (as eigenvalues) and wavefunctions (as eigenfunctions) of H<sub>2</sub>CS. The Hamiltonian and basis functions used to construct the eigenfunctions were expressed in terms of the valence coordinates:  $r_{CS}$  (the C–S bond),  $r_{CH_1}$ , and  $r_{CH_2}$  (the C–H bonds),  $\alpha_{SCH_1}$  and  $\alpha_{SCH_2}$  (the H–C–S angles), and  $\tau$  (the dihedral angle between the H–C–S planes). The DMS was also expressed in terms of these coordinates, and along with the eigenfunctions, was used to compute ro–vibrational line intensities using the GPU-code MPI-GAIN (Al-Refaie, Yurchenko & Tennyson 2017).

### 2.1 Kinetic energy operator

TROVE usually employs an approximate KEO that is constructed as a power series expansion in terms of linearized coordinates when solving the ro–vibrational Schrödinger equation. Here, an exact analytic expression in terms of the valence coordinates has been derived and implemented as an external input for the TROVE calculations. Benefits of using an exact KEO are more accurate results and a refined PES that can be used in other nuclear motion codes.

Expressing the Hamiltonian in terms of valence coordinates should also be expected to result in faster convergence in the computed energy levels, especially at high  $J_s$ , reducing the computational burden and allowing for the use of a more compact ro–vibrational basis set (Yachmenev & Yurchenko 2015).

For the KEO, a (right-handed) body-fixed frame was chosen such that the  $x$ -axis bisected the dihedral angle  $\tau$ , the  $z$ -axis pointed along the C–S bond, and the frame’s origin was at the centre of mass of the nuclei. The KEO was represented in a sum-of-products form which enabled it to be used by the program TROVE with only minor modifications. In this representation, each  $G$  matrix element in the KEO operator (and the pseudo-potential part, see Yurchenko et al. 2007 for details) has the form

$$\sum_{ijklmn} a_{ijklmn} f_{CS}^i(r_{CS}) f_{CH_1}^j(r_{CH_1}) f_{CH_2}^k(r_{CH_2}) f_{SCH_1}^l(\alpha_{SCH_1}) \times f_{SCH_2}^m(\alpha_{SCH_2}) f_{\tau}^n(\tau), \quad (1)$$

where the  $a_{ijklmn}$  are the coefficients and the  $f_s$  label functions of the different coordinates. For each mode, there is a finite set of functions. For the bond lengths, they are

$$\frac{1}{r}, \frac{1}{r^2}, \frac{1}{r^3}, \frac{1}{r^4}. \quad (2)$$

For the bond angles, they are

$$\cos \alpha, \cot \alpha, \csc \alpha, \sin \alpha, \cot^2 \alpha, \cot \alpha \csc \alpha, \csc^2 \alpha. \quad (3)$$

Finally, for the dihedral angle, they are

$$\cos(\tau/2), \sin(\tau/2), \cos^2(\tau/2), \cos(\tau/2)\sin(\tau/2), \sin^2(\tau/2). \quad (4)$$

The KEO was constructed using the Sørensen (1979) method that assumes the so-called Wilson integration volume with a Jacobian equals to 1. The full KEO is reproduced in Appendix A.

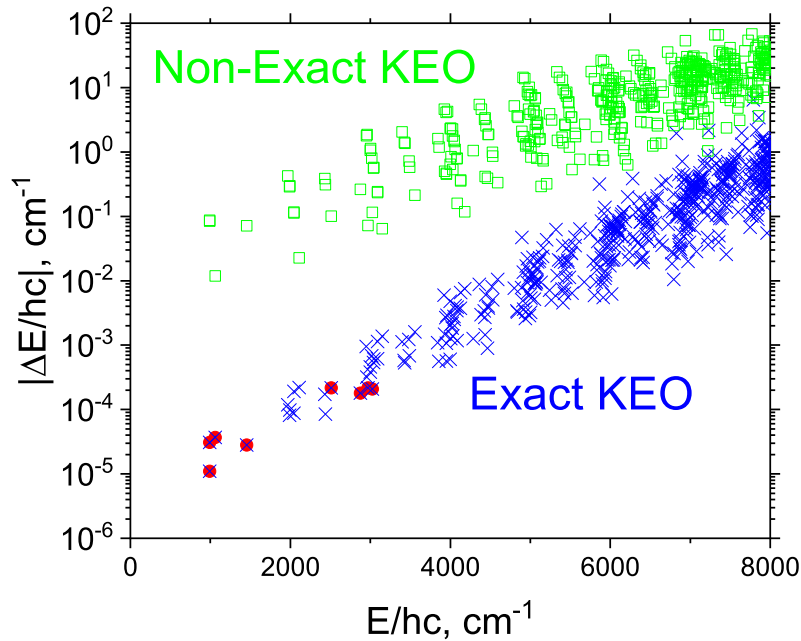
### 2.2 Ro–vibrational basis set

TROVE uses a three-step construction process to build the symmetry-adapted ro–vibrational basis set. In step 1, one-dimensional primitive basis functions  $\phi_{v_i}$ , where  $v_i$  corresponds to the  $i$ th function for the  $v$ th vibrational mode, are calculated numerically for each vibrational degrees of freedom (dof; six in total for H<sub>2</sub>CS) by integrating the respective one-dimensional vibrational Schrödinger equations using the Numerov–Cooley method (Noumerov 1924; Cooley 1961). The one-dimensional Hamiltonian is generated by freezing all vibrational modes except the one of interest.

At step 2, the symmetry-adapted vibrational basis set ( $J = 0$ ) is formed from four independent subsets of the Numerov–Cooley primitive basis functions  $\phi_{v_i}$  for the (i) C–S stretch, (ii) C–H<sub>1</sub> and C–H<sub>2</sub> stretches, (iii) H<sub>1</sub>–C–S and H<sub>2</sub>–C–S bending modes, and (iv)  $\tau$  dihedral out-of-plane mode by solving the corresponding reduced-mode vibrational Schrödinger equations. The reduced-mode Hamiltonian is formed by integrating out the modes for the other subsets using their ground Numerov–Cooley eigenfunctions. These reduced-mode eigenfunctions are symmetrized according to the irreducible representations of the molecular symmetry group  $C_{2v}(M)$  and form the basis functions for solving the full six-dimensional  $J = 0$  problem. For H<sub>2</sub>CS, a vibrational basis set cut-off was applied based on a polyad number truncation scheme,

$$P = v_{CS} + 2(v_{CH_1} + v_{CH_2}) + v_{H_1CS} + v_{H_2CS} + v_{\tau} \leq P_{\max}, \quad (5)$$

which used the polyad cut-off  $P_{\max} = 14$ , chosen to ensure that full ro–vibrational calculations were computationally tractable when



**Figure 1.** Convergence of the vibrational ( $J = 0$ ) TROVE calculations as a difference between  $P = 14$  and  $16$  energy term values H<sub>2</sub>CS (cm<sup>-1</sup>). The blue crosses correspond to the calculations using the new exact KEO implementation in TROVE; the green squares correspond to calculations using the approximate, linearized KEO version of TROVE as in Yachmenev et al. (2011, 2013). The red circles indicate states used in the refinement of the spectroscopic model.

generating a line list up to very high  $J$  (as discussed by Yachmenev et al. 2013).

The final ro–vibrational basis set for  $J > 0$  computations is then formed as a contracted product of these vibrational functions  $\Phi_{i_{\text{vib}}}^{J=0, \Gamma_{\text{vib}}}$  (7912 functions) and symmetrized rigid rotor wavefunctions  $|J, K, \Gamma_{\text{rot}}, m\rangle$  as given by

$$\Psi_{i_{\text{vib}}, K}^{J, \Gamma} = \left\{ \Phi_{i_{\text{vib}}}^{J=0, \Gamma_{\text{vib}}} \times |J, K, \Gamma_{\text{rot}}, m\rangle \right\}^{\Gamma}, \quad (6)$$

where  $\Gamma$ ,  $\Gamma_{\text{vib}}$ , and  $\Gamma_{\text{rot}}$  are the total, vibrational, and rotational symmetries in  $\mathcal{C}_{2v}(\text{M})$ ,  $K = |k|$ ,  $k$  and  $m$  is the projection of the angular momentum on the molecular  $z$  and laboratory  $Z$ -axis, respectively, and  $i_{\text{vib}}$  is a TROVE vibrational index to count the  $\Phi_{i_{\text{vib}}}^{J=0, \Gamma_{\text{vib}}}$  functions regardless of their symmetry. For further details of the TROVE symmetry adaptation and contraction procedure, see Yurchenko et al. (2017).

For H<sub>2</sub>CS, the TROVE wavefunctions were assigned the following ro–vibrational quantum numbers:  $v_1, v_2, v_3, v_4, v_5, v_6, J, K, \Gamma, \Gamma_{\text{vib}}$ , and  $\Gamma_{\text{rot}}$ , which can be correlated to the spectroscopic (normal mode) quantum numbers commonly encountered. Details of the mapping between the two, which is essential for ‘MARVELizing’ the final MOTY line list, i.e. replacing the computed energy levels with the more accurate MARVEL values, will be discussed in Section 4.1.

In order to assess the convergence of the TROVE  $P = 14$  vibrational basis set used in calculations, a larger basis set with  $P = 16$  was used to compute the vibrational  $J = 0$  energies of H<sub>2</sub>CS using the same refined PES (see below). Fig. 1 shows the differences between the  $P = 14$  and  $16$  term values. Below 6000 cm<sup>-1</sup> the differences are within 0.1 cm<sup>-1</sup>, which then gradually increase to up to 1 cm<sup>-1</sup> below 8000 cm<sup>-1</sup> with a few states reaching up to 10 cm<sup>-1</sup>. This convergence should be better than the expected ‘detuning’ of the empirical PES, considering that the experimentally determined vibrational states do not extend beyond 3050 cm<sup>-1</sup>, indicated in this figure by red circles.

In Fig. 1, we also demonstrate that the basis set convergence of the current ro–vibrational calculations based on the new implementations of the exact KEO of H<sub>2</sub>CS in TROVE is a significant improvement compared to the TROVE calculation by Yachmenev et al. (2011, 2013), where an approximate, linearized Hamiltonian operator was used. Here, we show analogous energy differences between the  $P = 14$  and  $16$  vibrational calculations of H<sub>2</sub>CS using the approximate linearized TROVE model in conjunction with the refined PES from this work. The improvement in the basis set convergence of 2–3 orders of magnitude is due to the new KEO in TROVE.

### 2.3 Potential energy surface and empirical refinement

The PES used in this work is based on an *ab initio* surface originally generated by Yachmenev et al. (2011), which utilized state-of-the-art electronic structure methods in the calculation process. The PES was represented as an expansion in terms of six valence coordinates,

$$\xi_i = 1 - \exp[-b_i(r_i - r_i^e)], \quad i = \text{CS}, \text{CH}_1, \text{CH}_2, \quad (7)$$

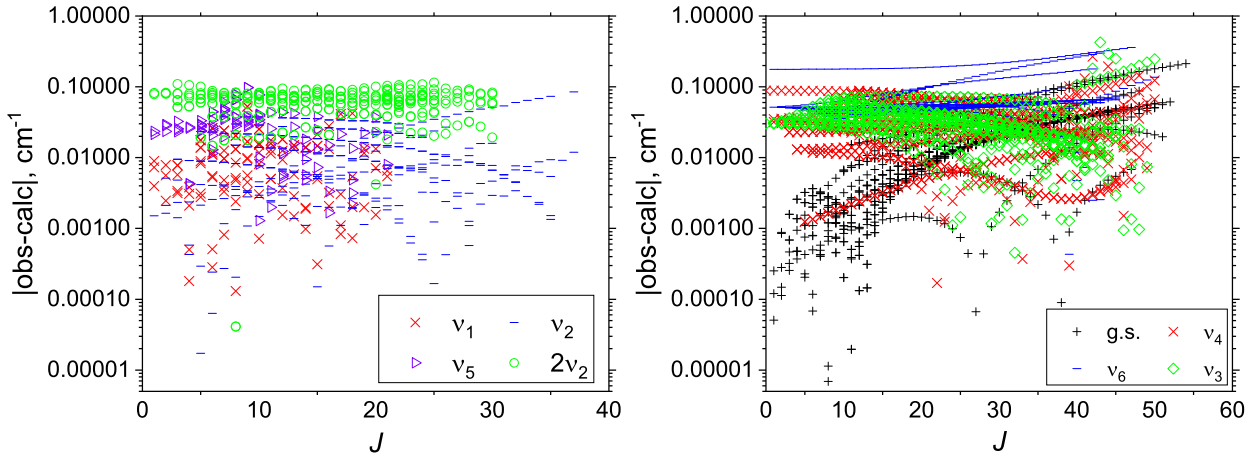
$$\xi_j = \alpha_j - \alpha_j^e, \quad j = \text{SCH}_1, \text{SCH}_2, \quad (8)$$

$$\xi_6 = 1 + \cos \tau. \quad (9)$$

Here,  $r_{\text{CS}}$ ,  $r_{\text{CH}_1}$ , and  $r_{\text{CH}_2}$  are the bond lengths,  $\alpha_{\text{SCH}_1}$  and  $\alpha_{\text{SCH}_2}$  are the bond angles,  $\tau$  is the dihedral angle between the SCH<sub>1</sub> and SCH<sub>2</sub> planes, and  $b_i$  is a Morse oscillator parameter. The PES was expressed analytically as,

$$V = \sum_{ijklmn} a_{ijklmn} \xi_{\text{CS}}^i \xi_{\text{CH}_1}^j \xi_{\text{CH}_2}^k \xi_{\text{SCH}_1}^l \xi_{\text{SCH}_2}^m \xi_{\tau}^n, \quad (10)$$

where  $a_{ijklmn}$  are the expansion parameters with maximum expansion order  $i + j + k + l + m + n = 6$  with the linear expansion parameters fixed to zero. A total of 413 parameters were utilized including the three equilibrium ( $r_{\text{CS}}^e$ ,  $r_{\text{CH}}^e$ ,  $\alpha_{\text{SCH}}^e$ ) and two Morse parameters ( $b_{\text{CH}}$ ,  $b_{\text{CS}}$ ). The expansion parameters assumed the values of the *ab*



**Figure 2.** The fitting residual errors (O–C), i.e. the energy difference (in  $\text{cm}^{-1}$ ) between the empirically derived MARVEL energies and computed TROVE values, as a function of the total angular momentum quantum number  $J$ . Residual errors are shown for seven vibrational bands and have been computed using the newly refined PES of  $\text{H}_2\text{CS}$ .

**Table 1.** Extract from a `.trans` file of the  $\text{H}_2\text{CS}$  MOTY line list.

$f$	$i$	$A_{fi}$ ( $\text{s}^{-1}$ )
1967530	2157952	5.6990e–01
9281842	9684141	2.1589e–11
21248596	21941059	2.5861e–12
5252014	5149230	9.6459e–04
12481388	12623346	8.6482e–11
1781054	1841556	2.4679e–16

*Notes.*  $f$ : upper state counting number;  $i$ : lower state counting number;  $A_{fi}$ : Einstein-A coefficient (in  $\text{s}^{-1}$ ).

*initio* PES from Yachmenev et al. (2011), with only the quadratic expansion parameters (14 in total) varied in the refinement to the empirically derived MARVEL energy levels. The refinement used  $\text{H}_2\text{CS}$  MARVEL energies for  $J = 0, 1, 2, 3, 4, 5, 8, 10$  (448 in total) covering the ground,  $v_1, v_2, v_3, v_4, v_5, v_6$ , and  $2v_2$  vibrational bands. For the equilibrium structural parameters, we used the optimized geometry values from Yachmenev et al. (2013):  $r_{\text{CS}}^e = 1.608\,952\text{ \AA}$ ,  $r_{\text{CH}}^e = 1.086\,848\text{ \AA}$ , and  $\alpha_{\text{SCH}}^e = 121.750^\circ$ . To ensure the refined PES maintained a realistic shape and gave reliable results in regions not sampled by the MARVEL data, the PES was constrained to the *ab initio* PES of Yachmenev et al. (2011) in the fitting.

The refinement showed very quick convergence. The quality of the fit is demonstrated in Fig. 2, where we have plotted the fitting residuals, i.e. the energy difference (in  $\text{cm}^{-1}$ ) between the MARVEL and computed TROVE  $\text{H}_2\text{CS}$  values, for the seven vibrational states used in the refinement. The majority of the fitting residuals are all below  $0.1\text{ cm}^{-1}$ , see Fig. 2, notably for the ground state which is to be expected since it corresponds to the lowest part of the PES. The errors are substantially smaller than those of the original *ab initio*  $\text{H}_2\text{CS}$  PES of Yachmenev et al. (2011), which would be in the region of  $1\text{--}5\text{ cm}^{-1}$ , demonstrating the improvements in accuracy that can be achieved with a refined PES. A FORTRAN version of the PES is given in the supporting information.

## 2.4 Dipole moment surfaces

A previously published *ab initio* DMS was utilized for intensity calculations. The DMS of Yachmenev et al. (2013) was computed

using coupled cluster theory CCSD(T) in conjunction with the augmented correlation-consistent basis set aug-cc-pVQZ(+d for S) (Dunning Jr 1989; Dunning Jr, Peterson & Wilson 2001). This DMS has already been used in room temperature line list calculations of  $\text{H}_2\text{CS}$  and with the *ab initio* PES was able to accurately reproduce complicated resonance effects, such as intensity borrowing, when modelling several vibrational bands affected by strong Coriolis coupling. The three dipole components making up the DMS were represented analytically using a sum-of-products expansion in terms of linear expansion variables for the six vibrational coordinates (further details can be found in Yachmenev et al. 2013).

## 3 THE MOTY LINE LIST OF $\text{H}_2\text{CS}$

The newly computed line list for  $\text{H}_2\text{CS}$ , called MOTY, contains 43 56 116 660 transitions between 52 292 454 states and covers the  $0\text{--}8000\text{ cm}^{-1}$  range for ro–vibrational states with rotational excitation up to  $J = 120$ . The lower and upper state energy thresholds were chosen to be  $8000$  and  $18000\text{ cm}^{-1}$ , respectively. The MOTY line list is provided in the ExoMol data format (Tennyson, Hill & Yurchenko 2013) and an extract from one of the `.trans` transition files and from the `.states` states file is shown in Tables 1 and 2, respectively. The `.trans` files are divided into  $1000\text{ cm}^{-1}$  wavenumber ranges to make them more manageable and they contain upper and lower state ID numbers along with the Einstein A coefficient (in  $\text{s}^{-1}$ ) of the transition between the states. The `.states` file contains a list of the ro–vibrational states of  $\text{H}_2\text{CS}$  with state ID numbers, energies (in  $\text{cm}^{-1}$ ), uncertainties (in  $\text{cm}^{-1}$ ), state lifetimes, and unique quantum numbers. Both the standard spectroscopic normal mode quantum numbers and TROVE quantum numbers are included in the `.states` file and the mapping between the two sets is discussed in Section 4, with ‘MARVELization’ of the MOTY line list detailed in Section 4.1.

As is standard now for ExoMol line lists, all molecular states possess an uncertainty. The uncertainties were defined either as the MARVEL uncertainty if available, or estimated via the expression (in  $\text{cm}^{-1}$ ):

$$\text{unc} = 0.4n_1 + 0.2(n_2 + n_3 + n_4 + n_5 + n_6) + 0.002J(J + 1), \quad (11)$$

where  $n_1\text{--}n_6$  are the TROVE quantum numbers. These uncertainties are only approximate and grow steadily with increasing rotational and vibrational excitation, where we tend to be conservative in estimating

**Table 2.** Extract from the `.states` file of the H<sub>2</sub>CS MOTY line list.

$i$	$\tilde{E}$ (cm <sup>-1</sup> )	$g$	$J$	$\delta$ (cm <sup>-1</sup> )	$\tau$ (s <sup>-1</sup> )	$\Gamma_{\text{tot}}$	$v_1$	$v_2$	$v_3$	$v_4$	$v_5$	$v_6$	$\Gamma_{\text{vib}}$	$J$	$K_a$	$K_c$	$i_v$	$ C_i^2 $	$n_1$	$n_2$	$n_3$	$n_4$	$n_5$	$n_6$	Ca/Ma	$\tilde{E}_T$ (cm <sup>-1</sup> )
1	0.000 000	1	0	0.000 001	Inf	A1	0	0	0	0	0	0	A1	0	0	0	1	1.0	0	0	0	0	0	0	Ma	0.000 000
2	1059.204 475	1	0	0.000 460	8.3703E-01	A1	0	0	1	0	0	0	A1	0	0	0	4	1.0	1	0	0	0	0	0	Ma	1059.234 861
3	1455.496 493	1	0	0.100 000	1.6832E+00	A1	0	1	0	0	0	0	A1	0	0	0	5	1.0	0	0	0	1	0	0	Ca	1455.496 493
4	1966.570 835	1	0	0.200 000	2.3271E-01	A1	0	0	0	0	0	2	A1	0	0	0	6	1.0	0	0	0	1	1	0	Ca	1966.570 835
5	1992.435 721	1	0	0.200 000	1.6450E-01	A1	0	0	0	2	0	0	A1	0	0	0	8	1.0	0	0	0	0	2	0	Ca	1992.435 721
6	2109.341 721	1	0	0.200 000	4.2379E-01	A1	0	0	2	0	0	0	A1	0	0	0	11	1.0	2	0	0	0	0	0	Ca	2109.341 721
7	2510.397 221	1	0	0.200 000	5.6662E-01	A1	0	1	1	0	0	0	A1	0	0	0	14	1.0	1	0	0	1	0	0	Ca	2510.397 221
8	2877.032 488	1	0	0.200 000	1.3711E-01	A1	0	2	0	0	0	0	A1	0	0	0	15	1.0	0	0	0	1	1	0	Ca	2877.032 488
9	2971.028 217	1	0	0.200 000	3.3900E-02	A1	1	0	0	0	0	0	A1	0	0	0	18	1.0	0	0	1	0	0	0	Ca	2971.028 217
10	3017.401 926	1	0	0.300 000	1.8251E-01	A1	0	0	1	0	0	2	A1	0	0	0	21	1.0	1	0	0	1	1	0	Ca	3017.401 926
11	3043.141 281	1	0	0.300 000	1.4094E-01	A1	0	0	1	2	0	0	A1	0	0	0	24	1.0	1	0	0	0	0	2	Ca	3043.141 281
12	3150.274 608	1	0	0.300 000	2.8533E-01	A1	0	0	3	0	0	0	A1	0	0	0	27	1.0	3	0	0	0	0	0	Ca	3150.274 608

*Notes.*  $i$ : state identifier;  $\tilde{E}$ : state term value;  $g$ : state degeneracy;  $J$ : state rotational quantum number;  $\delta$ : energy uncertainty;  $\tau$ : lifetime;  $\Gamma_{\text{tot}}$ : total symmetry in  $C_{2v}(M)$ ;  $v_1$ – $v_6$ : normal mode vibrational quantum numbers;  $\Gamma_{\text{vib}}$ : symmetry of vibrational contribution in  $C_{2v}(M)$ ;  $J$ : state rotational quantum number;  $K_a$  and  $K_c$ : state oblate and prolate quantum numbers;  $i_v$ : vibrational state ID;  $|C_i^2|$ : largest coefficient used in the assignment;  $n_1$ – $n_6$ : TROVE vibrational quantum numbers; Ca/Ma: Label indicating if the term value is based on the MARVEL (Ma) or the MOTY energy list (Ca);  $\tilde{E}_{\text{AVTY}}$ : original MOTY state term value.

values for highly excited states. It is important for users of the MOTY line list to be aware that transitions that are suitable for high-resolution applications possess a defined (MARVEL) uncertainty, and those have uncertainties devined by equation (11) are not for high-resolution studies. States with an estimated uncertainty may well be highly accurate, however, without experimental verification this is difficult to verify.

The temperature-dependent partition function  $Q(T)$  of H<sub>2</sub>CS has been computed on a 1 K grid in the 1–2000 K range and is available as a `.p.f` file from the ExoMol website. Partition function values were computed by summing over all the computed ro–vibrational states of the MOTY line list and used nuclear spin statistical weights of  $g_{\text{ns}} = 1, 1, 3, 3$  for states of symmetry  $A_1, A_2, B_1,$  and  $B_2$ , respectively. The dependence of  $Q(T)$  on temperature is illustrated in Fig. 3, where we also show the partition function values from the Cologne Database for Molecular Spectroscopy (CDMS; Endres et al. 2016) data base. The latter only contains ground vibrational transitions which explains the lower value at  $T = 500$  K.

As part of the ExoMol services, cross-sections of H<sub>2</sub>CS are provided via the ExoMol cross-sections app. The app can be used, for example, to generate absorption cross-sections of H<sub>2</sub>CS with the Doppler broadening line profile for temperatures from 100 to 2000 K covering the wavenumber range from 0 to 8000 cm<sup>-1</sup> on a grid with a resolution of up to 0.01 cm<sup>-1</sup> (Hill, Yurchenko & Tennyson 2013). We also provide the H<sub>2</sub>CS spectroscopic model used in TROVE calculations in the form of a TROVE input file, containing the potential energy and DMS parameters as well as the basis set specifications. This input file can be used with the FORTRAN code TROVE freely available from GitHub via [www.github.com/exomol](http://www.github.com/exomol).

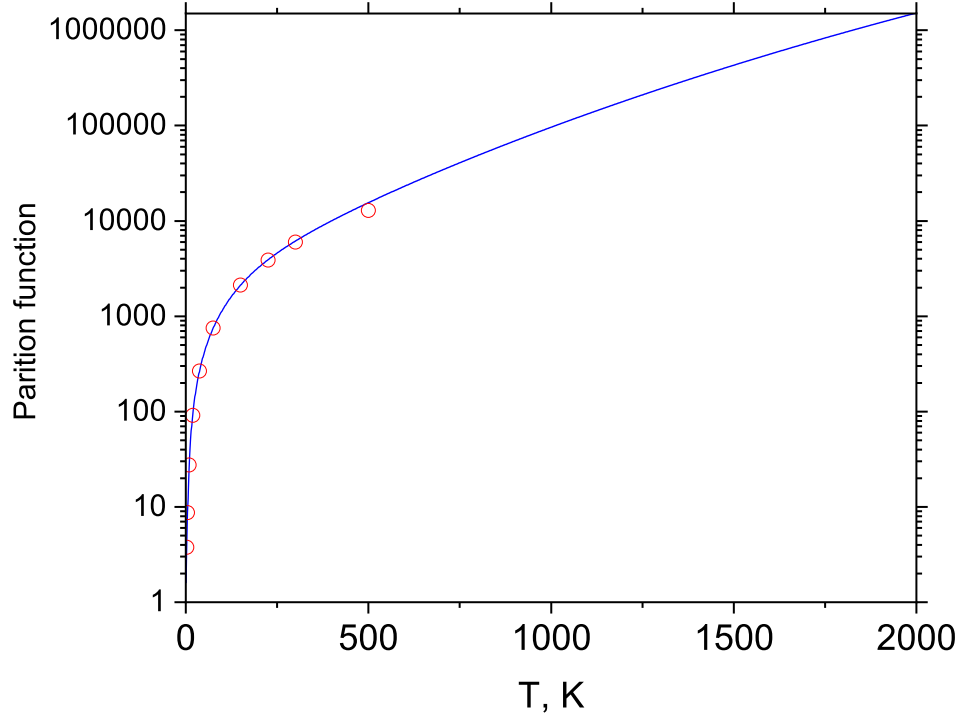
#### 4 MAPPING BETWEEN TROVE AND NORMAL MODE QUANTUM NUMBERS

To provide the line list with both the TROVE quantum numbers and the normal mode vibrational quantum numbers, a mapping between the two is necessary. The TROVE local mode quantum numbers  $v_i$  are related to the primitive basis functions  $\phi_{v_i}$  that are used to build the full symmetry-adapted basis set in ro–vibrational calculations. Here,  $v_{\text{CS}}, v_{\text{CH}_1},$  and  $v_{\text{CH}_2}$  correspond to the stretching dof,  $v_{\text{H}_1\text{CS}}$  and  $v_{\text{H}_2\text{CS}}$  to the bending dof, and  $v_{\tau}$  to the dihedral angle. The TROVE local mode quantum numbers  $v_i^{\text{Ca}}$  are used to label the calculated ro–vibrational states based on the highest contribution from the different primitive basis functions. To be of use, they must be correlated to the standard spectroscopic normal mode quantum numbers  $n_i$ . The following relations between the normal mode  $n_i$  and TROVE local mode quantum numbers  $v_i$  apply:

$$\begin{aligned} n_1 &= v_2^{\text{Ca}}, \\ n_2 + n_5 &= v_2^{\text{Ca}} + v_3^{\text{Ca}}, \\ n_3 + n_6 &= v_4^{\text{Ca}} + v_5^{\text{Ca}}, \\ n_4 &= v_6^{\text{Ca}}, \end{aligned}$$

these are also detailed in Table 3, which shows the mapping between the local and normal modes for the vibrational modes in H<sub>2</sub>CS.

To correlate the TROVE rotational quantum numbers  $J, K,$  and  $\Gamma_{\text{tot}}$  to the standard rotational quantum numbers for H<sub>2</sub>CS  $J^{K_a, K_c}$ , we first note that the TROVE quantum number  $K$  coincides with  $K_a$ , while  $K_c$  is reconstructed using the standard spectroscopic rules as described by Bunker & Jensen (1998), see their table 12-9. The relationship between  $K_a$  and  $K_c$  with rotational symmetry is listed in Table 4.



**Figure 3.** Temperature dependence of the partition function  $Q(T)$  of  $\text{H}_2\text{CS}$  computed using the MOTY line list (solid line) and compared to the CDMS values (Endres et al. 2016; circles).

**Table 3.** Mapping between the TROVE local modes and the standard normal mode notation  $\nu_i$ . Here,  $\Gamma$  is the symmetry of the fundamental mode in  $\mathcal{C}_{2v}(\text{M})$ .

$\Gamma$	Vibrations	Normal mode, $\nu$	TROVE modes
$A_1$	C–H symm. stretch	$\nu_1$	2 and 3
$A_1$	S–C–H symm. bend	$\nu_2$	4 and 5
$A_1$	S–C stretch	$\nu_3$	1
$B_1$	Out-of-plane	$\nu_4$	6
$B_2$	C–H asymm. stretch	$\nu_5$	2 and 3
$B_2$	S–C–H asymm. bend	$\nu_6$	4 and 5

**Table 4.** Mapping between the rotational symmetry  $\Gamma_{\text{rot}}$  in  $\mathcal{C}_{2v}(\text{M})$  and rotational quantum numbers  $K_a, K_c$  used for  $\text{H}_2\text{CS}$  (Bunker & Jensen 1998).

$K_a, K_c$	$\Gamma_{\text{rot}}$	$K_a, K_c$	$\Gamma_{\text{rot}}$
Even, even	$A_1$	Odd, even	$B_2$
Even, odd	$A_2$	Odd, odd	$B_1$

#### 4.1 MARVELization of the line list

Following the ExoMolHD strategy (Yurchenko & Tennyson 2021), in order to improve the accuracy of the MOTY line list and tailor it to high-resolution applications, the calculated energy levels are replaced with the empirically derived MARVEL values, if available. To this end, the MARVEL and TROVE computed energies have to be matched by correlating their quantum numbers.

Table 5 lists the previously described correlation for the eight vibrational bands used in the MARVELization of the  $\text{H}_2\text{CS}$  line list. To be efficient, the correlation is built via the TROVE vibrational quantum index  $i_{\text{vib}}$  as indicated in Table 5, see also equation (6), where  $i_{\text{vib}}$  is used to identify the vibrational basis functions  $\Phi_{i_{\text{vib}}}^{J=0, \Gamma_{\text{vib}}}$ .

**Table 5.** Mapping between the TROVE local mode quantum numbers  $\nu_i$  and the standard normal mode notations  $n \nu_i$  for the states used in the MARVELization. Here,  $\Gamma$  is the symmetry of the fundamental mode in  $\mathcal{C}_{2v}(\text{M})$  and  $i_{\text{vib}}$  is a TROVE counting index identifying the vibrational ( $J = 0$ ) states, see equation (6).

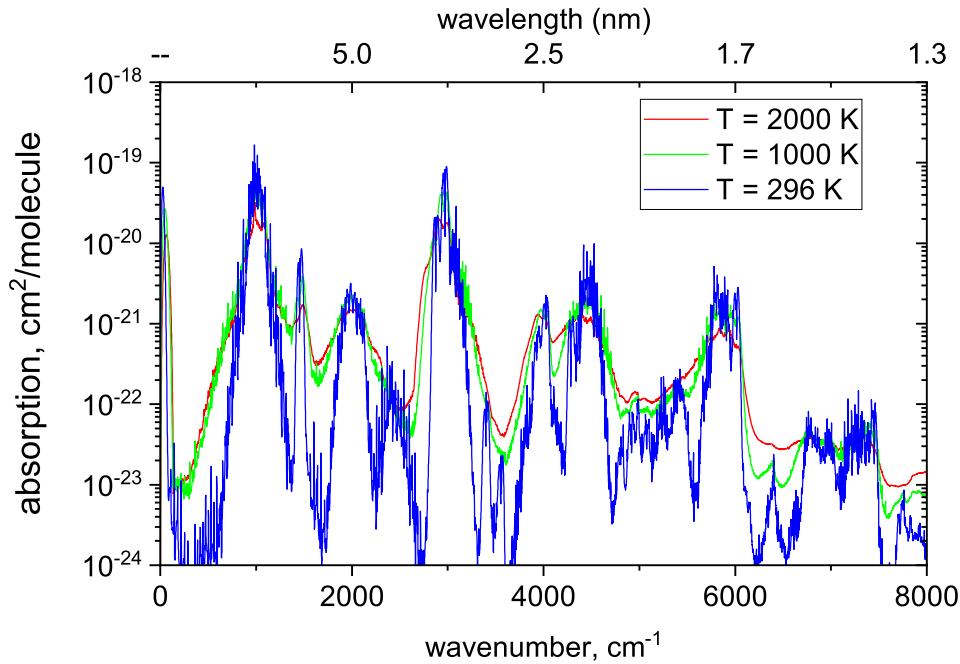
State	$\Gamma$	$\nu_1$	$\nu_2$	$\nu_3$	$\nu_4$	$\nu_5$	$\nu_6$	$i_{\text{vib}}$
g.s.	$A_1$	0	0	0	0	0	0	1
$\nu_4$	$B_1$	0	0	0	0	0	1	2
$\nu_6$	$B_2$	0	0	0	0	1	0	3
$\nu_3$	$A_1$	1	0	0	0	0	0	4
$\nu_2$	$A_1$	0	0	0	1	0	0	5
$2\nu_2$	$A_1$	0	0	0	1	1	0	15
$\nu_1$	$A_1$	0	0	1	0	0	0	18
$\nu_5$	$B_2$	0	1	0	0	0	0	22

## 5 SPECTRA SIMULATIONS

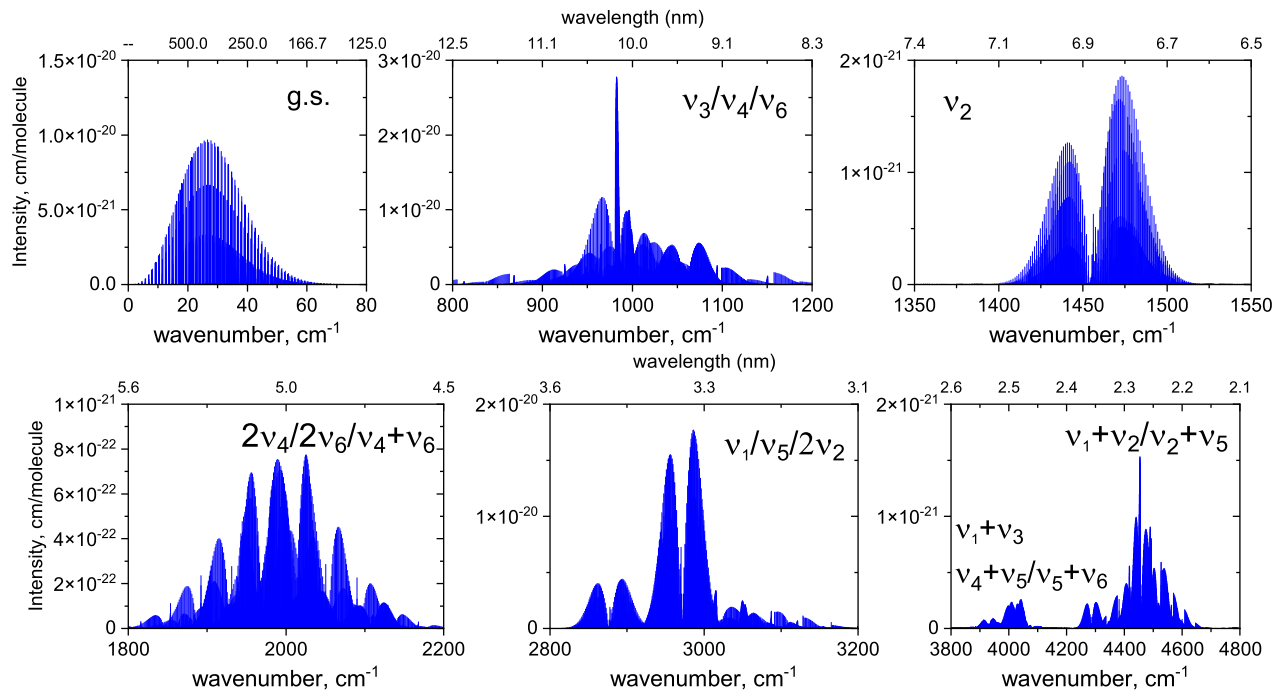
All spectra simulations used the open source program EXOCROSS (Yurchenko, Al-Refaie & Tennyson 2018; see [github.com/exomol](https://github.com/exomol)). In Fig. 4, a general overview of the  $\text{H}_2\text{CS}$  spectrum at three different temperatures (296, 1000, and 2000 K) is plotted. Absolute absorption cross-sections were computed at a resolution of  $1 \text{ cm}^{-1}$  using a Gaussian line profile with a half width at half-maximum (HWHM) of  $1 \text{ cm}^{-1}$ . As expected, the higher temperature spectra exhibit spectral flattening as the previously weaker features gain intensity.

A more detailed band-by-band illustration of the room temperature spectrum of  $\text{H}_2\text{CS}$  is shown in Fig. 5.

In Fig. 6 (left-hand panel), a comparison of a synthetic microwave spectrum of  $\text{H}_2\text{CS}$  between the MOTY ExoMol line list and the CDMS (Müller et al. 2001, 2005) is given. The good agreement of the intensities indicate that the equilibrium dipole moments used in CDMS and the MOTY line list agree well. The CDMS line list is



**Figure 4.** Absorption cross-sections of  $H_2CS$  at  $T = 296, 1000,$  and  $2000$  K. A Gaussian line profile with a HWHM of  $1\text{ cm}^{-1}$  was used on a grid of resolution  $1\text{ cm}^{-1}$ .



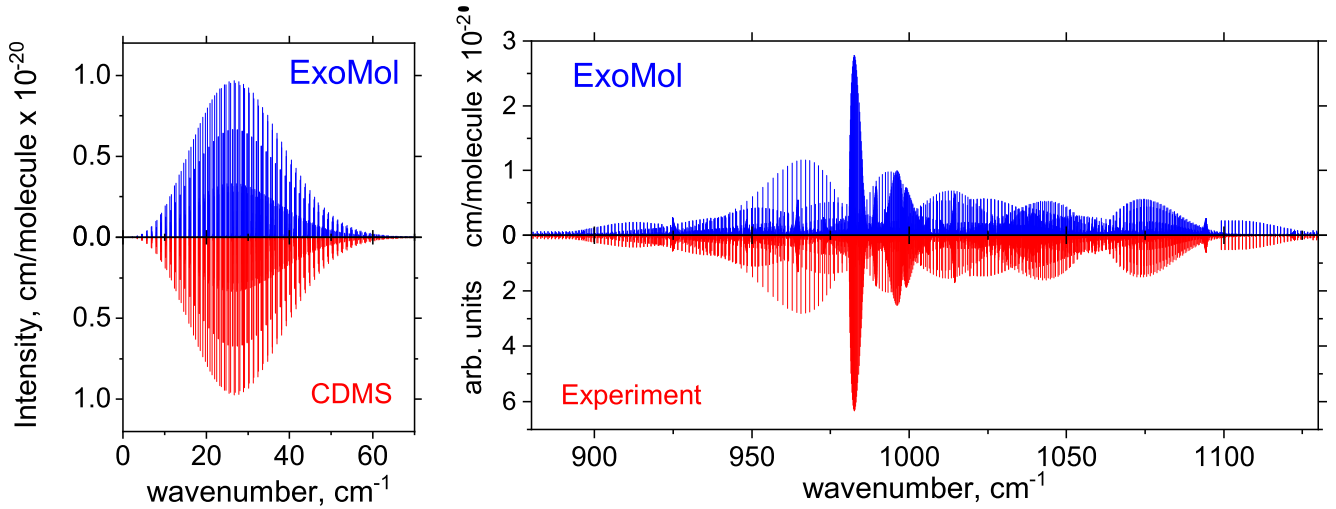
**Figure 5.** Overview of the strongest bands of  $H_2CS$  at  $T = 296$  K.

based on a dipole moment value of  $1.6491$  D due to Fabricant et al. (1977). Our *ab initio* dipole moment gives an expectation value of  $1.64872$  D for the ground vibrational state. These values can also be compared to an experimental laser-Stark value of  $1.6483$  D obtained by Cox, Hubbard & Kato (1982).

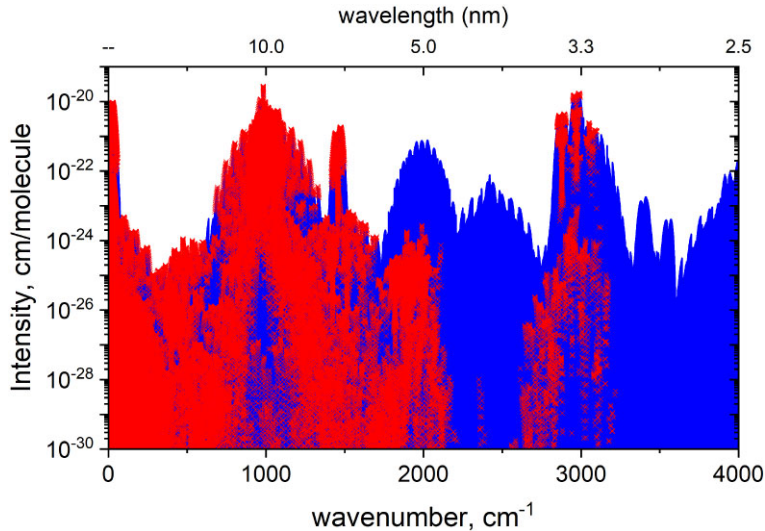
The right-hand panel of Fig. 6 shows a comparison with a high-resolution spectrum of  $H_2CS$  by Flaud et al. (2008) for the  $\nu_3/\nu_4/\nu_6$  region, although the experimental data only give relative intensities. The agreement is excellent. Further confirmation of the reliability

of the MOTY line list intensities can be seen when comparing to recent theoretical calculations by Erfort, Tschöpe & Rauhut (2020), which gave an intensity of  $1.01 \times 10^{-20}\text{ cm molecule}^{-1}$  (Erfort et al. 2020; using coupled cluster theory, CCSD(T)/cc-pVTZ) for the  $\nu_4$  transition  $3_{03} \leftarrow 3_{13}$ , compared to our value of  $1.02 \times 10^{-20}\text{ cm molecule}^{-1}$  at  $T = 296$  K.

In Fig. 7, we illustrate the impact of the MARVELization procedure for the MOTY line list of  $H_2CS$ . Here, we compare room temperature spectra of  $H_2CS$  computed using the entire



**Figure 6.** Left: Comparison of ExoMol and CDMS (Endres et al. 2016) microwave absorption cross-sections of H<sub>2</sub>CS at  $T = 296$  K. Right: Comparison of ExoMol (cm molecule<sup>-1</sup>) and experimental (arb.units) absorption cross-sections by Flaud et al. (2008) of the  $\nu_3/\nu_4/\nu_6$  region of H<sub>2</sub>CS at  $T = 296$  K.



**Figure 7.** Room temperature ( $T = 296$  K) spectra of H<sub>2</sub>CS showing the coverage of the MARVELized transitions (red crosses) compared to the total MOTY spectrum.

‘unMARVELized’ MOTY line list and using the MARVELized (upper and lower) states only. We obtained 81 446 MARVELized transitions of H<sub>2</sub>CS with  $T = 296$  K intensities below the HITRAN threshold of  $10^{-30}$ /cm<sup>-1</sup> molecule<sup>-1</sup>. The total number of H<sub>2</sub>CS transitions at  $T = 296$  K above this threshold (0–8000 cm<sup>-1</sup>) is 32 094 935. For example, the weak transitions in the region of 2000 cm<sup>-1</sup> belong to the hot bands formed from  $\nu_1/\nu_5$  (upper) and  $\nu_4/\nu_6$  (lower) states, which have not been experimentally characterized. This figure provides some indication of the H<sub>2</sub>CS spectral regions that are suitable for high-resolution applications.

## 6 CONCLUSION

A hot comprehensive ro–vibrational line list of H<sub>2</sub>CS has been presented. The MOTY line list contains over 43.5 billion transitions and covers the 0–8000 cm<sup>-1</sup> range (wavelengths  $\lambda > 1.3$  μm) for states with rotational excitation up to  $J = 120$ . One of the major improvements over previous H<sub>2</sub>CS line lists is the use of a new

empirically refined PES. The H<sub>2</sub>CS PES is the most accurate to date and was determined by refining to 218 empirically derived MARVEL energy levels up to  $J = 5$ , leading to orders-of-magnitude improvements in the accuracy of the calculated energy levels, and thus transition wavenumbers. Variational calculations also utilized an exact KEO and a Hamiltonian operator expressed in valence coordinates which greatly improved the convergence of computed energy levels. This enabled a more accurate and extensive line list to be computed as higher  $J$  states could be reliably considered. Another significant feature of the MOTY line list is that it has been MARVELized by replacing the calculated energy levels with more accurate empirically derived MARVEL values therefore tailoring the line list to high-resolution applications in certain spectral windows. The prevalence of thioformaldehyde in a variety of astronomical environments means the MOTY line list will be useful for future detection of this molecule. Notably, we expect the MOTY line list to be of relevance to exoplanet atmospheric studies concerned with sulphur chemistry, since H<sub>2</sub>CS is a simple organosulphur molecule.



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## DATA AVAILABILITY

The data underlying this article are available as part of the supporting information, from the ExoMol data base at [www.exomol.com](http://www.exomol.com) and the CDS data base. The MOTY line list (states, transition, partition function files, and a TROVE input specifying the spectroscopic model of H<sub>2</sub>CS) can be downloaded from [www.exomol.com](http://www.exomol.com)<sup>1</sup> and cdsarc.u-strasbg.fr. The open access programmes EXOCROSS and TROVE are available via [github.com/exomol](https://github.com/exomol).

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## SUPPORTING INFORMATION

A *Mathematica* script used to generate the KEO are the refined PES of H<sub>2</sub>CS used to compute the line list and the refined PES are provided.

Supplementary data are available at *MNRAS* online.

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<sup>1</sup><https://exomol.com/data/molecules/H2CS/1H2-12C-32S/MOTY/>

**APPENDIX: KINETIC ENERGY OPERATOR**

We use reduced masses

$$\frac{1}{\mu_{CS}} = \frac{1}{m_C} + \frac{1}{m_S} \quad (\text{A1})$$

and similarly for the others. Translational components:

$$G_{XX} = G_{YY} = G_{ZZ} = \frac{1}{m_C + 2m_H + m_S}. \quad (\text{A2})$$

Rotational components:

$$\begin{aligned} G_{XX} = G_{YY} &= \frac{1}{\mu_{CS} r_{CS}^2}, \\ G_{XZ} &= \frac{\cot \alpha_{SCH_1} \cos(\tau/2)}{2\mu_{CS} r_{CS}^2} + \frac{\cot \alpha_{SCH_2} \cos(\tau/2)}{2\mu_{CS} r_{CS}^2} - \frac{\csc \alpha_{SCH_1} \cos(\tau/2)}{2m_C r_{CS} r_{CH_1}} - \frac{\csc \alpha_{SCH_2} \cos(\tau/2)}{2m_C r_{CS} r_{CH_2}}, \\ G_{YZ} &= \frac{\cot \alpha_{SCH_2} \sin(\tau/2)}{2\mu_{CS} r_{CS}^2} - \frac{\cot \alpha_{SCH_1} \sin(\tau/2)}{2\mu_{CS} r_{CS}^2} + \frac{\csc \alpha_{SCH_1} \sin(\tau/2)}{2m_C r_{CS} r_{CH_1}} - \frac{\csc \alpha_{SCH_2} \sin(\tau/2)}{2m_C r_{CS} r_{CH_2}}, \\ G_{ZZ} &= \frac{\cot \alpha_{SCH_1} \cot \alpha_{SCH_2} \cos^2(\tau/2)}{2\mu_{CS} r_{CS}^2} - \frac{\cot \alpha_{SCH_1} \cot \alpha_{SCH_2} \sin^2(\tau/2)}{2\mu_{CS} r_{CS}^2} + \frac{\cot^2 \alpha_{SCH_1}}{4\mu_{CS} r_{CS}^2} \\ &+ \frac{\cot^2 \alpha_{SCH_2}}{4\mu_{CS} r_{CS}^2} - \frac{\csc \alpha_{SCH_1} \cot \alpha_{SCH_2} \cos^2(\tau/2)}{2m_C r_{CS} r_{CH_1}} + \frac{\csc \alpha_{SCH_1} \cot \alpha_{SCH_2} \sin^2(\tau/2)}{2m_C r_{CS} r_{CH_1}} \\ &- \frac{\cot \alpha_{SCH_1} \csc \alpha_{SCH_1}}{2m_C r_{CS} r_{CH_1}} - \frac{\cot \alpha_{SCH_1} \csc \alpha_{SCH_2} \cos^2(\tau/2)}{2m_C r_{CS} r_{CH_2}} - \frac{\cot \alpha_{SCH_2} \csc \alpha_{SCH_2}}{2m_C r_{CS} r_{CH_2}} \\ &+ \frac{\cot \alpha_{SCH_1} \csc \alpha_{SCH_2} \sin^2(\tau/2)}{2m_C r_{CS} r_{CH_2}} + \frac{\csc^2 \alpha_{SCH_1}}{4\mu_{CH} r_{CH_1}^2} + \frac{\csc \alpha_{SCH_1} \csc \alpha_{SCH_2} \cos^2(\tau/2)}{2m_C r_{CH_1} r_{CH_2}} \\ &- \frac{\csc \alpha_{SCH_1} \csc \alpha_{SCH_2} \sin^2(\tau/2)}{2m_C r_{CH_1} r_{CH_2}} + \frac{\csc^2 \alpha_{SCH_2}}{4\mu_{CH} r_{CH_2}^2}. \end{aligned} \quad (\text{A3})$$

Coriolis components:

$$\begin{aligned} G_{Xr_{CH_1}} &= \frac{\sin \alpha_{SCH_1} \sin(\tau/2)}{m_C r_{CS}}, \\ G_{Xr_{CH_2}} &= -\frac{\sin \alpha_{SCH_2} \sin(\tau/2)}{m_C r_{CS}}, \\ G_{X\alpha_{SCH_1}} &= \frac{\cos \alpha_{SCH_1} \sin(\tau/2)}{m_C r_{CS} r_{CH_1}} - \frac{\sin(\tau/2)}{\mu_{CS} r_{CS}^2}, \\ G_{X\alpha_{SCH_2}} &= \frac{\sin(\tau/2)}{\mu_{CS} r_{CS}^2} - \frac{\cos \alpha_{SCH_2} \sin(\tau/2)}{m_C r_{CS} r_{CH_2}}, \\ G_{X\tau} &= \frac{\cot \alpha_{SCH_2} \cos(\tau/2)}{\mu_{CS} r_{CS}^2} + \frac{\csc \alpha_{SCH_1} \cos(\tau/2)}{m_C r_{CS} r_{CH_1}} - \frac{\csc \alpha_{SCH_2} \cos(\tau/2)}{m_C r_{CS} r_{CH_2}} - \frac{\cot \alpha_{SCH_1} \cos(\tau/2)}{\mu_{CS} r_{CS}^2}, \\ G_{Yr_{CH_1}} &= \frac{\sin \alpha_{SCH_1} \cos(\tau/2)}{m_C r_{CS}}, \\ G_{Yr_{CH_2}} &= \frac{\sin \alpha_{SCH_2} \cos(\tau/2)}{m_C r_{CS}}, \\ G_{Y\alpha_{SCH_1}} &= \frac{\cos \alpha_{SCH_1} \cos(\tau/2)}{m_C r_{CS} r_{CH_1}} - \frac{\cos(\tau/2)}{\mu_{CS} r_{CS}^2}, \\ G_{Y\alpha_{SCH_2}} &= \frac{\cos \alpha_{SCH_2} \cos(\tau/2)}{m_C r_{CS} r_{CH_2}} - \frac{\cos(\tau/2)}{\mu_{CS} r_{CS}^2}, \\ G_{Y\tau} &= \frac{\cot \alpha_{SCH_1} \sin(\tau/2)}{\mu_{CS} r_{CS}^2} + \frac{\cot \alpha_{SCH_2} \sin(\tau/2)}{\mu_{CS} r_{CS}^2} - \frac{\csc \alpha_{SCH_1} \sin(\tau/2)}{m_C r_{CS} r_{CH_1}} - \frac{\csc \alpha_{SCH_2} \sin(\tau/2)}{m_C r_{CS} r_{CH_2}}, \\ G_{Zr_{CH_1}} &= \frac{\sin \alpha_{SCH_1} \cot \alpha_{SCH_2} \sin(\tau/2) \cos(\tau/2)}{m_C r_{CS}} - \frac{\sin \alpha_{SCH_1} \csc \alpha_{SCH_2} \sin(\tau/2) \cos(\tau/2)}{m_C r_{CH_2}}, \\ G_{Zr_{CH_2}} &= \frac{\csc \alpha_{SCH_1} \sin \alpha_{SCH_2} \sin(\tau/2) \cos(\tau/2)}{m_C r_{CH_1}} - \frac{\cot \alpha_{SCH_1} \sin \alpha_{SCH_2} \sin(\tau/2) \cos(\tau/2)}{m_C r_{CS}}, \end{aligned}$$

$$\begin{aligned}
 G_{Z\alpha_{\text{SCH}_1}} &= -\frac{\cot \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{\cos \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
 &+ \frac{\csc \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} - \frac{\cos \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}}, \\
 G_{Z\alpha_{\text{SCH}_2}} &= \frac{\cot \alpha_{\text{SCH}_1} \sin(\tau/2) \cos(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{\csc \alpha_{\text{SCH}_1} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
 &- \frac{\cot \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \sin(\tau/2) \csc(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\csc \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}}, \\
 G_{Z\tau} &= -\frac{\cot^2 \alpha_{\text{SCH}_1}}{2\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{\cot^2 \alpha_{\text{SCH}_2}}{2\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{\cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_1}}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} - \frac{\cot \alpha_{\text{SCH}_2} \csc \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} \\
 &- \frac{\csc^2 \alpha_{\text{SCH}_1}}{2\mu_{\text{CH}} r_{\text{CH}_1}^2} + \frac{\csc^2 \alpha_{\text{SCH}_2}}{2\mu_{\text{CH}} r_{\text{CH}_2}^2}.
 \end{aligned}$$

Vibrational components:

$$\begin{aligned}
 G_{r_{\text{CS}} r_{\text{CS}}} &= \frac{1}{\mu_{\text{CS}}}, \\
 G_{r_{\text{CS}} r_{\text{CH}_1}} &= \frac{\cos \alpha_{\text{SCH}_1}}{m_{\text{C}}}, \\
 G_{r_{\text{CS}} r_{\text{CH}_2}} &= \frac{\cos \alpha_{\text{SCH}_2}}{m_{\text{C}}}, \\
 G_{r_{\text{CS}} \alpha_{\text{SCH}_1}} &= -\frac{\sin \alpha_{\text{SCH}_1}}{m_{\text{C}} r_{\text{CH}_1}}, \\
 G_{r_{\text{CS}} \alpha_{\text{SCH}_2}} &= -\frac{\sin \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CH}_2}}, \\
 G_{r_{\text{CH}_1} r_{\text{CH}_1}} &= \frac{1}{\mu_{\text{CH}}}, \\
 G_{r_{\text{CH}_1} r_{\text{CH}_2}} &= -\frac{\sin \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}}} + \frac{\sin \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}}} + \frac{\cos \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2}}{m_{\text{C}}}, \\
 G_{r_{\text{CH}_1} \alpha_{\text{SCH}_1}} &= -\frac{\sin \alpha_{\text{SCH}_1}}{m_{\text{C}} r_{\text{CS}}}, \\
 G_{r_{\text{CH}_1} \alpha_{\text{SCH}_2}} &= \frac{\sin \alpha_{\text{SCH}_1} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CS}}} - \frac{\sin \alpha_{\text{SCH}_1} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CS}}} + \frac{\sin \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_2}} \\
 &- \frac{\sin \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_2}} - \frac{\cos \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CH}_2}}, \\
 G_{r_{\text{CH}_1} \tau} &= \frac{2 \sin \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}}} - \frac{2 \sin \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CH}_2}}, \\
 G_{r_{\text{CH}_2} r_{\text{CH}_2}} &= \frac{1}{\mu_{\text{CH}}}, \\
 G_{r_{\text{CH}_2} \alpha_{\text{SCH}_1}} &= \frac{\sin \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CS}}} - \frac{\sin \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CS}}} + \frac{\cos \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_1}} \\
 &- \frac{\cos \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_1}} - \frac{\sin \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CH}_1}}, \\
 G_{r_{\text{CH}_2} \alpha_{\text{SCH}_2}} &= -\frac{\sin \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CS}}}, \\
 G_{r_{\text{CH}_2} \tau} &= \frac{2 \cot \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}}} - \frac{2 \csc \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CH}_1}}, \\
 G_{\alpha_{\text{SCH}_1} \alpha_{\text{SCH}_1}} &= \frac{1}{\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{2 \cos \alpha_{\text{SCH}_1}}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} + \frac{1}{\mu_{\text{CH}} r_{\text{CH}_1}^2}, \\
 G_{\alpha_{\text{SCH}_1} \alpha_{\text{SCH}_2}} &= -\frac{\sin^2(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{\cos^2(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{\cos \alpha_{\text{SCH}_1} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} + \frac{\cos \alpha_{\text{SCH}_1} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
 &- \frac{\cos \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\cos \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\sin \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} \\
 &- \frac{\cos \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} + \frac{\cos \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}}.
 \end{aligned}$$

$$\begin{aligned}
G_{\alpha_{\text{SCH}_1} \tau} &= -\frac{2 \cot \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{2 \cos \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
&+ \frac{2 \csc \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} - \frac{2 \cos \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}}, \\
G_{\alpha_{\text{SCH}_2} \alpha_{\text{SCH}_2}} &= \frac{1}{\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{2 \cos \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{1}{\mu_{\text{CH}} r_{\text{CH}_2}^2}, \\
G_{\alpha_{\text{SCH}_2} \tau} &= -\frac{2 \cot \alpha_{\text{SCH}_1} \sin(\tau/2) \cos(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{2 \cot \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} \\
&- \frac{2 \csc \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} + \frac{2 \csc \alpha_{\text{SCH}_1} \sin(\tau/2) \cos(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}}, \\
G_{\tau \tau} &= -\frac{2 \cot \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{2 \cot \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CS}}^2} + \frac{\cot^2 \alpha_{\text{SCH}_1}}{\mu_{\text{CS}} r_{\text{CS}}^2} \\
&+ \frac{\cot^2 \alpha_{\text{SCH}_2}}{\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{2 \csc \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} - \frac{2 \csc \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
&- \frac{2 \cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_1}}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} + \frac{2 \cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\csc^2 \alpha_{\text{SCH}_1}}{\mu_{\text{CH}} r_{\text{CH}_1}^2} \\
&- \frac{2 \cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{\mu_{\text{CS}} r_{\text{CS}} r_{\text{CH}_2}} - \frac{2 \cot \alpha_{\text{SCH}_2} \csc \alpha_{\text{SCH}_2}}{m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\csc^2 \alpha_{\text{SCH}_2}}{\mu_{\text{CH}} r_{\text{CH}_2}^2} \\
&- \frac{2 \csc \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} + \frac{2 \csc \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}}.
\end{aligned}$$

Pseudo-potential:

$$\begin{aligned}
U &= -\frac{\cot \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{4\mu_{\text{CS}} r_{\text{CS}}^2} + \frac{\cot \alpha_{\text{SCH}_2} \csc \alpha_{\text{SCH}_1} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} + \frac{\cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} \\
&- \frac{\csc \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} + \frac{\cot \alpha_{\text{SCH}_2} \sin \alpha_{\text{SCH}_1} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} - \frac{\csc \alpha_{\text{SCH}_2} \sin \alpha_{\text{SCH}_1} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} \\
&+ \frac{\cot \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} - \frac{\csc \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} - \frac{\sin \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \cos^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} \\
&+ \frac{\cot^2 \alpha_{\text{SCH}_1}}{32\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{\cot^2 \alpha_{\text{SCH}_1}}{16\mu_{\text{CH}} r_{\text{CH}_1}^2} + \frac{\cot^2 \alpha_{\text{SCH}_2}}{32\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{\cot^2 \alpha_{\text{SCH}_2}}{16\mu_{\text{CH}} r_{\text{CH}_2}^2} - \frac{5 \csc^2 \alpha_{\text{SCH}_1}}{32\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{\csc^2 \alpha_{\text{SCH}_1}}{16\mu_{\text{CH}} r_{\text{CH}_1}^2} - \frac{5 \csc^2 \alpha_{\text{SCH}_2}}{32\mu_{\text{CS}} r_{\text{CS}}^2} \\
&- \frac{\csc^2 \alpha_{\text{SCH}_2}}{16\mu_{\text{CH}} r_{\text{CH}_2}^2} + \frac{\cot \alpha_{\text{SCH}_1} \cot \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{4\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{\cot \alpha_{\text{SCH}_2} \csc \alpha_{\text{SCH}_1} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
&- \frac{\cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\csc \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} - \frac{\cot \alpha_{\text{SCH}_2} \sin \alpha_{\text{SCH}_1} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} \\
&+ \frac{\csc \alpha_{\text{SCH}_2} \sin \alpha_{\text{SCH}_1} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} - \frac{\cot \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} + \frac{\csc \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} \\
&+ \frac{\sin \alpha_{\text{SCH}_1} \sin \alpha_{\text{SCH}_2} \sin^2(\tau/2)}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} - \frac{\cos \alpha_{\text{SCH}_1}}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} - \frac{\cos \alpha_{\text{SCH}_1} \cos \alpha_{\text{SCH}_2}}{4m_{\text{C}} r_{\text{CH}_1} r_{\text{CH}_2}} - \frac{\cos \alpha_{\text{SCH}_2}}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} \\
&+ \frac{\cot \alpha_{\text{SCH}_1} \csc \alpha_{\text{SCH}_1}}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_1}} + \frac{\cot \alpha_{\text{SCH}_2} \csc \alpha_{\text{SCH}_2}}{4m_{\text{C}} r_{\text{CS}} r_{\text{CH}_2}} - \frac{3}{16\mu_{\text{CS}} r_{\text{CS}}^2} - \frac{3}{16\mu_{\text{CS}} r_{\text{CH}_1}^2} - \frac{3}{16\mu_{\text{CH}} r_{\text{CH}_2}^2}.
\end{aligned}$$

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