

A Computed Room Temperature Line List for Phosphine

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Abstract

An accurate and comprehensive room temperature rotation-vibration transition line list for phosphine ($^{31}\text{PH}_3$) is computed using a newly refined potential energy surface and a previously constructed *ab initio* electric dipole moment surface. Energy levels, Einstein A coefficients and transition intensities are computed using these surfaces and a variational approach to the nuclear motion problem as implemented in the program TROVE. A ro-vibrational spectrum is computed, covering the wavenumber range 0 to 8000 cm^{-1} . The resulting line list, which is appropriate for temperatures up to 300 K, consists of a total of 137 million transitions between 5.6 million energy levels. Several of the band centres are shifted to better match experimental transition frequencies. The line list is compared to the most recent HITRAN database and other laboratorial sources. Transition wavelengths and intensities are generally found to be in good agreement with the existing experimental data, with particularly close agreement for the rotational spectrum. An analysis of the comparison between the theoretical data created and the existing experimental data is performed, and suggestions for future improvements and assignments to the HITRAN database are made.

1. Introduction

Phosphine (PH_3) is an extremely toxic, rigid molecule found in the lower troposphere of the Earth and is expected to be of great importance for the spectroscopic studies of giant-planets, particularly for probing the deeper layers of the atmospheres [1], and cool stars. Phosphine has for a long time known to be abundant in Jupiter [2, 3, 4], where it is partially responsible for the continuum opacity in the $5\ \mu\text{m}$ region of the atmosphere [5] and where it is a marker for vertical convection zones. It has also been found on Saturn [1, 2], and is expected to be present in extrasolar gas giants.

Further spectroscopic investigations of phosphine in the atmospheres of astrophysical bodies requires an extrapolation of the spectra over a wide range of temperatures, pressures and path lengths. The ability to accurately identify and interpret phosphine features in astrophysical spectra is dependent on the existence of a complete description of the phosphine spectra, for all relevant temperatures.

The current CDMS [53] and HITRAN [7] databases encapsulate decades of accurate laboratorial measurements of phosphine spectra at room temperature. However, between the two databases, only a total of 22230 lines are described, all within the $0 - 3600\ \text{cm}^{-1}$ region. CDMS contains transitions between $0 - 300\ \text{cm}^{-1}$ with J values of up to 34, sensitive to $10^{-36}\ \text{cm}/\text{molecule}$, while HITRAN has transitions between $770 - 3600\ \text{cm}^{-1}$ stronger than $10^{-28}\ \text{cm}/\text{molecule}$ with $J \leq 23$. These databases are designed to be valid for temperatures below 300 K. Other sources (see Table 1 below) mostly overlap with the HITRAN and CDMS databases, and only add little more to these compilations.

The quantity and variety of spectral information required to correctly characterise hot astrophysical bodies is often beyond what can be expected to be delivered experimentally. The ExoMol project [8] (see www.exomol.com) aims to provide a solution to this problem by performing accurate quantum mechanical calculations, with the aim of providing appropriate spectroscopic data on all relevant molecules for the analysis and characterisation of cool stellar and exoplanet atmospheres. This is, however, a challenging alternative to measuring spectra, requiring highly accurate potential energy surfaces (PES), dipole moment surfaces (DMS) and an ability to generate precise wave functions and eigenvalues for the nuclear Schrödinger equation and the associated transition intensities. In practice, at least for the PESs, it is also necessary to incorporate a certain amount of experimental refinement.

Our aim is to develop complete line lists which could be used to accurately model the atmospheres of stars and Jupiter-like planets, for which a high temperature ($\geq 1500\ \text{K}$) line list for PH_3 is of special importance. In order to establish and test the production of such a line list, which will consist of approximately 1 billion transitions, it is first pertinent to perform a simulation of the room-temperature intensities for this molecule, which can be tested against existing reliable experimental sources. In the present paper a cool (300 K) line list is presented.

There have been several previous attempts at *ab initio* studies of the behavior

and characteristics of phosphine spectra. Wang et al [9] computed a PES using coupled cluster CCSD(T) theory and Dunning’s correlation consistent cc-pVTZ basis sets, followed by a vibrational analysis based on second-order perturbation theory. They determined several spectroscopic constants and fundamentals largely within 4 cm^{-1} of the observed results. Zheng et al [10] calculated an *ab initio* three-dimensional P-H stretching DMS using density functional theory and found absolute band intensities agreed with observed ones within a factor of 2. Yurchenko et al [11] presented a calculated, albeit empirically refined, PES and, using a variational method, calculated the rotational energy levels in the vibrational ground state of PH_3 for $J \leq 80$ [12], thus establishing for the first time the existence of six-fold near-degenerate ro-vibrational energy clusters for this molecule. Subsequently, Yurchenko et al [13] computed an entirely *ab initio* six-dimensional dipole moment surface (DMS) at the CCSD(T)/aug-cc-pVTZ level of theory for the electronic ground state of PH_3 . This is the DMS that is used in the present work, and is described in detail below. It gives calculated transition moments within an average of 22.5% of the experimental data. The same paper also presented a PES determined by empirically refining an existing *ab initio* surface. Ovsyannikov et al [14] complemented the variationally calculated PES and DMS of Yurchenko et al [13] by calculating a PES at the CCSD(T) level using aug-cc-pV(Q+d)Z and aug-cc-pVQZ basis sets for P and H, respectively, and then presenting a list of computed vibrational transition moments for the electronic ground state of PH_3 [15]. The PES used here is a refined version of the PES presented by Ovsyannikov et al [14], and is discussed further below. Recently, Nikitin et al produced a new PES [16] and used it to calculate several vibrational energy levels with only a marginal deviation from experiment. Nikitin et al then successfully modelled the lower three polyads of phosphine using a variational approach [17]. The resulting line list is discussed and compared to the present work in Section 4.6. The phosphine line list presented here can be considered the analogue of the ammonia line lists computed by Yurchenko et al [18], using TROVE.

2. Background

Phosphine is a well behaved symmetric top molecule belonging to the group-15 hydrides. The phosphorous atom is positioned on the axis of symmetry, perpendicular to the plane containing the equidistant three hydrogen atoms. As is common with molecules having $C_{3v}(M)$ symmetry [19], there is a splitting of rotational levels with $K = 3n$ (where $n \geq 1$) in non degenerate vibrational states [20]. Phosphorous has only one non-synthetic, stable isotope, ^{31}P , so only $^{31}\text{PH}_3$ is considered in the present work.

Phosphine has near degenerate ν_1 symmetric and ν_3 antisymmetric stretching modes [9] with frequencies of about twice that of the asymmetric bending mode, ν_4 . It is therefore natural to consider its spectrum in terms of polyads, which will be described in section 3.1. There is also a strong Coriolis interaction between the ν_2 and ν_4 bending bands which causes large distortion effects in observations [21]. Recorded phosphine spectra show no inversion splitting. The

tunnelling effect found in ammonia is expected to occur in phosphine but so far attempts to detect it have failed due to its considerably higher potential energy barrier (12300 cm^{-1}) between the two symmetry-related minima [22, 23].

In the absence of the inversion splitting, phosphine is characterized by the $C_{3v}(\text{M})$ molecular group symmetry, i.e. each eigenfunction transforms according to one of the irreducible representations A_1 , A_2 and E , the latter of which is two-fold degenerate.

To fully describe the internal rovibrational motion of phosphine one needs the following minimal set of quantum numbers [24]:

$$n_1, n_2, n_3, n_4, L_3, L_4, L, \Gamma_{\text{vib}}, J, K, \Gamma_{\text{rot}}, \Gamma_{\text{tot}}, \quad (1)$$

where $L_3 = |l_3|$, $L_4 = |l_4|$, $L = |l|$, $K = |k|$. Here the vibrational quantum numbers n_1 (symmetric stretch), n_2 (symmetric bend), n_3 (asymmetric stretch) and n_4 (asymmetric bend) correspond to excitations of, respectively, the ν_1 , ν_2 , ν_3 and ν_4 modes. The doubly degenerate modes ν_3 and ν_4 require additional quantum numbers $L_3 = |l_3|$ and $L_4 = |l_4|$ describing the projections of the corresponding angular momenta (see, for example, Ref. [19]). The vibrational quantum number $L = |l|$, characterizes the coupling of l_3 and l_4 . Γ_{rot} , Γ_{vib} , and Γ_{tot} are, respectively, the symmetry species of the rotational, vibrational and total internal wave-functions in the molecular symmetry group $C_{3v}(\text{M})$, spanning A_1 , A_2 and E . J is the total angular momentum and $K = |k|$, $k = -J, \dots, J$ is the projection of the total angular momentum on the molecule fixed axis z .

Out of these twelve quantum numbers, only two are rigorously conserved quantum numbers (J and Γ_{tot}). This quantum number description is similar to that proposed by Down et al [24] for ammonia. The rigorous dipole selection rules include transitions between states with Γ_{tot} satisfying $A_1 \leftrightarrow A_2$ and $E \leftrightarrow E$ and $\Delta J = 0, \pm 1$ only. Strong transitions obey the dipole selection rules $\Delta J = \pm 1$ and $\Delta K = 0$. There are also weakly allowed transitions which occur when, due to centrifugal distortion, the C_{3v} geometrical symmetry of the molecule is broken and K is no longer a good quantum number [25].

Only approximate selection rules can be associated with the vibrational quantum numbers n_1, n_2, n_3, n_4 .

Table 1 presents an overview of the experimental spectra recorded for phosphine. This spans research starting from 1951 to the most recent body of work, published in 2009. The naming convention for each individual source consist of two digits for the date and first two letters of up to four authors' surnames, e.g. S. Yurchenko, M. Carvajal, P. Jensen, H. Lin, J. Zheng & W. Thiel 2005 becomes 05YuCaJeLi.

Table 1: Experimental sources of phosphine spectra.

Source	No of lines	Range(cm^{-1})	Uncertainties(cm^{-1})	Intensities
71DaNeWoKl[26] ^{a,c}	53	0.0 – 4.8	Yes	No
77HeGo[27] ^{a,c}	27	0.0 – 17.8	No	Calculated
81BeBuPoSh[22] ^{a,c}	68	0.0 – 35.6	No	No
51LoSt[28]	1	0.94	10^{-6}	No
81PiPoCo[29] ^c	1	8.9	3.3×10^{-7}	No
69HeGo[30]	3	8.9 - 17.8	8.3×10^{-6}	No
06CaPu[56] ^a	25	8.9 - 35.6	<i>Yes</i>	No
81BeBuGeKr[58] ^a	52	10 - 35.7	<i>No</i>	No
74ChOk[25] ^{a,c}	12	14.6 – 15.8	2.7×10^{-5}	Calculated
79KrMeSk[31] ^c	4	17.1 – 17.8	1.7×10^{-6}	No
00FuLo[32]	155(+2500)	36 – 125(+750-1400)	0.004	No
88FuCa[57] ^a	118	44.5 – 166.6	0.002	No
02BrSaKlCo[33] ^b	>1060	770 – 1335	0.0002	Some
81TaDaGo[21] ^b	1318	818 – 1340	0.05	Calculated
97AiHaSpKr[34]	62(+250 ^d)	924 – 1085	6×10^{-6}	No
01HeZhHuLi[10]	28	992 – 9040	<0.01	Yes
04SaArBoWa[35]	26	995 – 1093	No	Yes
02UIBeKoZh[36]	≥ 4500	1750 – 9200	≤ 0.0005	No
92TaLaLeGu[1] ^b	6605	1885 – 2445	0.0054	Yes
05WaChChDi[37]	1760	1950 - 2480 3280 - 3580	<6%	Yes
80BaMaNaTa[38]	1244	2184 – 2446	0.005	Approximate
02Suarez[39]	138	2300 – 2381	0.006	Yes
06BuSaKlBr[40] ^b	8075	2721 – 3601	9.2×10^{-4}	Yes
07Kshiraga[41]	400	2730 – 3100	3.9×10^{-4}	No
73MaSaOl[42]	414	2760 – 3050	No	No
04UIBeKoZh[43]	>700	3280 – 3580	0.005	Transmittance

^aUsed in CDMS [53].^bUsed in HITRAN 08 [7].^cUsed in JPL[6].^dCompiled from other sources.

3. Method

The production of the phosphine line list presented here relies on the computational program TROVE (Theoretical ROVibrational Energies) [44], which uses a general variational approach to calculate the ro-vibrational energies for small semi-rigid polyatomic molecules of arbitrary structure. It employs a numerical finite basis representation. The nuclear kinetic energy operator is numerically constructed through a recursive scheme using a Taylor series expansion in terms of the molecule’s internal coordinates, which requires no analytical pre-derivation, making its creation self-contained. This process is an extension of the XY3 program [45]. Although this procedure allows TROVE to simulate the nuclear motion for an arbitrary molecule, it does not calculate an exact kinetic energy operator. As such, the accuracy of the rotation-vibration energies depends on the expansion order of the kinetic energy operator and its associated level of convergence. Similarly, the potential energy operator is expressed as a Taylor-type expansion, and its numerical derivation can lead to an accumulation of round-off errors. For a more detailed description of the TROVE functionality, including its treatment of the kinetic and potential energy operators, the reader is referred elsewhere [44].

3.1. Convergence

To construct the ro-vibrational basis set, the contraction scheme described perviously [46, 47, 48] was followed. This method employs 1-D basis functions $\phi_i(\xi_i^\ell)$ ($i = 1 \dots 6$), where ξ_1^ℓ , ξ_2^ℓ , and ξ_3^ℓ are the linearized versions of the three stretching coordinates Δr_i , $i = 1, 2, 3$ ($\Delta r_i = r_{\text{PH}_i} - r_e$), and ξ_4^ℓ , ξ_5^ℓ , and ξ_6^ℓ are the linearized versions of the three bending coordinates $\Delta \alpha_i$, $i = 1, 2, 3$ associated with the variation of the bond angles α_{23} , α_{13} , and α_{12} , respectively. Here r_e and α_e are, respectively, the equilibrium values for the bond length and the inter bond angle facing the bond length.

The primitive basis functions for each mode ν_i , $i = 1 \dots 6$, are generated by solving the corresponding Schrödinger equation with the Numerov-Cooley method [49, 50]. In order to control the size of the basis sets at all contraction steps the polyad-truncation scheme [44] is used, based on the PH_3 polyad number defined by

$$P = 2(s_1 + s_2 + s_3) + b_1 + b_2 + b_3 \leq P_{\text{max}} \quad (2)$$

where s_i and b_i are the quantum numbers associated with the basis functions, ϕ_{s_i} and ϕ_{b_i} , for the stretching modes and the bending modes, respectively. The basis set is formed only by those basis functions ϕ_s , ϕ_b for which $P \leq P_{\text{max}}$.

The primitive basis functions obtained at this step are then used to solve two 3-D Schrödinger equations for each class of modes, stretching and bending, separately. The corresponding basis sets are formed as direct products of the primitive basis functions satisfying $2(s_1 + s_2 + s_3) \leq P_{\text{max}}$ (stretching) and $b_1 + b_2 + b_3 \leq P_{\text{max}}$ (bending) in agreement with Eq. (2), so the maximal vibrational excitation is taken to be $P_{\text{max}}/2$ and P_{max} for each of the stretching

and bending modes, respectively. The resulting two sets of eigenfunctions Φ_m^{str} and Φ_n^{bend} are then classified using $C_{3v}(\text{M})$ symmetry by analysing their transformational properties and assigned with the normal mode quantum numbers $n_1, n_3^{L_3}$ (stretching modes) and $n_2, n_4^{L_4}$ (bending modes) for future ease of line assignment and comparison to the experimental spectroscopic data. To this end a correlation between the primitive and normal mode quantum numbers is needed to be constructed. At the initial $J = 0$ stage, this is a straightforward process due to the relatively small number of solutions and simple polyad structure of PH_3 energies: only approximately 700 sets of vibrational quantum numbers have to be translated for all energy levels ($J \leq 31$) to be represented by both sets of quantum numbers.

The reassignment from local mode quantum numbers ν_i ($i = 1 \dots 6$) to normal mode quantum numbers n_i ($i = 1 \dots 4$), $L_3 = |l_3|$ and $L_4 = |l_4|$ was performed by applying the following rules :

$$\begin{aligned} s_1 + s_2 + s_3 &= n_1 + n_3, \\ b_1 + b_2 + b_3 &= n_2 + n_4, \\ l_3 &= -n_3, -n_3 + 2, \dots, n_3 - 2, n_3, \\ l_4 &= -n_4, -n_4 + 2, \dots, n_4 - 2, n_4. \end{aligned}$$

It was also assumed that the symmetric modes ν_1 and ν_2 have lower energies than their asymmetric equivalents ν_3 and ν_4 , respectively, and also that the vibrational energy grows when l_3 or l_4 increase.

At the next contraction step, a 6-D vibrational basis is formed as a direct product of the stretching and bending functions Φ_m^{str} and Φ_n^{bend} , contracted using the (normal mode) version of the polyad number truncation given by Eq. (2):

$$P = 2(n_1 + n_3) + n_2 + n_4 \leq P_{\text{max}}. \quad (3)$$

The 6-D functions Φ_m^{str} , Φ_n^{bend} are then symmetrized by the standard reduction technique [19]. At the next stage the vibrational ($J = 0$) Schrödinger equation is solved variationally by diagonalizing the ($J = 0$) Hamiltonian matrix on this 6-D vibrational basis. Using the P_{max} polyad restriction from (3) 1455, 1125, and 2571 basis functions are obtained for the A_1 , A_2 and E symmetries, respectively.

At the last contraction step, the resulting eigenfunctions of the 6-D $J = 0$ problem are used to form our final ro-vibrational basis functions in the so-called $J = 0$ representation [46] through a direct product with the symmetrized rigid rotor wave functions $|J, K, \tau\rangle$, where $\tau = 0, 1$ indicates the rotational parity $(-1)^\tau$ as defined in Ref.[45]. In the $J = 0$ representation building the Hamiltonian matrix is straightforward as the vibrational part of the matrix is diagonal.

The ro-vibrational eigenfunctions obtained can be naturally assigned using a scheme based on the largest contribution of the basis set expansion. The chosen quantum numbers are the vibrational normal mode quantum numbers

$n_1, n_2, n_3^{L_3}$ and $n_4^{L_4}$, the rotational quantum numbers J, K , the vibrational symmetry Γ_{rot} , the vibrational symmetry Γ_{vib} , and the total symmetry Γ_{tot} , where $\Gamma_{\text{rot/vib/tot}} = A_1, A_2, E$. The projection of the total vibrational angular momentum L is also included to reduce ambiguity in the description of the energy levels, with the following rules: L must be a multiple of 3 if $\Gamma_{\text{vib}} = A_1, A_2$ and cannot be a multiple of 3 if $\Gamma_{\text{vib}} = E$. By convention, the lower energy value was assigned to $|L_3 - L_4|$. However, at higher energies, TROVE does not necessarily assign unique quantum labels to every state. In this case we have chosen to assign all states to $|L_3 - L_4|$.

The accuracy with which high ro-vibrational states can be computed depends heavily on the size of the Hamiltonian matrix. To ensure that the calculation of the energy levels was sufficiently complete, convergence tests were performed to facilitate the choice of the polyad number (P_{max}), ranging from $P = 4$ to $P = 18$. A selection of the results from the convergence tests is given in table 2.

Table 2: $J = 0$ energy eigenvalue convergence with polyad number, P_{max} , used to generate the basis set, where $P(n) = (P_{\text{max}} = 18) - (P_{\text{max}} = n)$.

Band	$P_{\text{max}}=18$	P(16)	P(14)	P(12)	P(10)	P(8)
ν_0	0.000	0.000	0.000	0.000	0.000	0.000
ν_2	992.732	-0.010	-0.033	-0.090	-0.243	-0.714
$2\nu_2$	1967.427	-0.133	-0.393	-0.939	-2.197	-5.321
$2\nu_4$	2222.257	-0.009	-0.033	-0.095	-0.283	-0.897
ν_1	2324.077	-0.003	-0.011	-0.036	-0.128	-0.436
$3\nu_2$	2911.866	-1.325	-3.472	-7.230	-14.335	-29.377
$\nu_2 + 2\nu_4$	3204.920	-0.115	-0.361	-0.941	-2.467	-7.220
$\nu_1 + \nu_2$	3322.008	-0.038	-0.132	-0.384	-1.135	-3.820
$3\nu_4$	3341.959	-0.028	-0.098	-0.295	-0.931	-3.434
$\nu_3 + \nu_4$	3439.304	-0.009	-0.034	-0.117	-0.437	-1.843
$4\nu_2$	3803.487	-9.750	-22.284	-39.911	-66.534	-108.521
$2\nu_2 + 2\nu_4$	4159.298	-1.025	-2.839	-6.309	-13.381	-28.564
$\nu_1 + 2\nu_2$	4303.413	-0.352	-1.074	-2.740	-6.742	-16.622
$\nu_2 + 3\nu_4$	4318.966	-0.257	-0.786	-1.981	-4.906	-12.643
$\nu_2 + \nu_3 + \nu_4$	4410.959	-0.131	-0.436	-1.196	-3.221	-9.642
$4\nu_4$	4422.841	-0.098	-0.325	-0.919	-2.707	-7.461

Ideally, the polyad number P_{max} would be chosen to give full convergence, which, as can be seen from table 2, is not achieved even for the highest polyad configuration, $P_{\text{max}} = 18$. However, although the calculations should improve as the polyad number increases, since the ‘spectroscopic’ PES (see Section 3.2)

used was generated with $P_{\max} = 14$, the energy levels move away from the observed values when a higher polyad configuration is used. The PES used in this work relies on a self consistency of parameters, and as such it is only an effective potential for a specific set of parameters, so $P_{\max} = 14$ was chosen.

Convergence tests were performed to test the expansions of the kinetic and potential energy operator and, although accuracy did improve considerably as the expansions grew, so did the computational cost involved. The convergence test presented in Table 2 was done using expansions of the kinetic energy operator and potential energy function both to fourth order, which accounts for the significant discrepancies between the computed band origins and the corresponding experimental values. In the present work, the expansions of the kinetic energy operator and potential energy function are truncated at the 6th and 8th order, respectively, which leads to much more accurate results, as can be seen below. Higher order expansions than these would have made the project computationally prohibitive.

3.2. Potential Energy Surface

The PES of phosphine used here is a refinement of the *ab initio* (CCSD(T)/aug-cc-pV(Q+d)Z) PES [14], done by performing a least square fit to available experimental ro-vibrational energies with $J = 0, 1, 2, 4$ and 10. These refinements followed the fitting procedure introduced in Ref. [47]. The refinement of the PES is represented as a correction ΔV to the *ab initio* PES, V_0 . The ro-vibrational $J = 0, 1, 2, 4, 10$ eigenfunctions of the ‘unperturbed’ Hamiltonian $H_0 = T + V_0$ are used as basis functions when iteratively solving the set of Schrödinger equations for the ro-vibrational behaviour of the molecule to minimise the associated functional in the least-squares fitting. With the final refined PES, a total root-mean-squares (rms) error of 0.03 cm^{-1} was obtained for the fitting energy set. Table 3 shows the change in the band centre values between the pre- and post-refinement PES.

As already mentioned, the kinetic energy operator is not calculated exactly and the basis set is artificially limited. Thus the refined PES must be considered an effective PES; it only gives the accurate results presented in Table 3 when used with TROVE and the parameters described above. The refined PES used here is given in the supplementary material to this paper in the form of a Fortran program.

The PES is a major source of error in the line list creation process, and as such constant updates and refinements are necessary. For example, although the current PES is appropriate for creating an accurate room temperature line list with wavenumbers up to 8000 cm^{-1} and $J \leq 31$, it would not necessarily remain so for a line list with parameters exceeding this. For example, the PES fails to accurately describe the potential of phosphine at very high excitations, showing some artificial minima in the region of dissociation. Consequently, this PES is not suitable for use in dynamical simulations without adjustment.

3.3. Dipole moment surface and transition intensities

An existing six-dimensional *ab initio* electric dipole moment (CCSD(T)/aug-cc-pVTZ) [13] is used to obtain the Einstein coefficients and transition intensities. This was calculated on a large grid of 10080 molecular geometries. The procedure used to compute absolute intensities mirrored that used for ammonia, and is described in detail in Yurchenko et al [18]. A number of vibrational transitional moments were calculated to help characterize the quality of the DMS.

Table 3 compares the empirical vibrational transition moment with values computed by TROVE, both in the present work and in the previous studies [13]. It is clear that the theoretical band intensities are in good agreement with experiment and, although the DMS used here is unchanged from that used in Ref. [13], the use of our more accurate PES (and hence wavefunctions) has meant that our new results have generally reduced the error in the calculations of the transition moments, from an average deviation of 22.5% to 10%.

Table 3: Calculated band centres and their respective transition moments (deviation from experimental values shown as a percentage). Uncertainties of the experimental (Obs) transition moments are given in parentheses (in units of the last digit quoted) where available.

Band	Band Centres (cm ⁻¹)			Transitions Moments (D ²)		
	Obs	Calc '06 [13]	Present	Obs	Calc '06 (%) [13]	Present (%)
ν_0	0.000	0.000	0.000	0.57395(30)	0.583(1.6%)	0.585(2.0%)
ν_2	992.130	992.500	992.152	0.08251(5)	0.085(2.5%)	0.084(2.2%)
ν_4	1118.310	1117.870	1118.322	0.08626(5)	0.087(0.3%)	0.089(3.7%)
$2\nu_2$	1972.550	1972.820	1972.590	0.00299(5)	0.003(9.7%)	0.004(21.4%)
$\nu_2 + \nu_4$	2108.150	2107.170	2108.169	0.01102(6)	0.009(15.6%)	0.014(25.3%)
$2\nu_4$	2226.830	2227.860	2226.835	0.0176(2)	0.006(68.8%)	0.018(1.5%)
$2\nu_4$	2234.930	2234.570	2234.940	0.0176(2)	0.002(91.5%)	0.013(23.5%)
ν_1	2321.120	2322.040	2321.142	0.0690	0.073(6.1%)	0.072(4.3%)
ν_3	2326.870	2329.180	2326.888	0.130	0.139(6.8%)	0.138(6.1%)

3.4. Empirical adjustment of the vibrational band centres

Although using a ‘spectroscopic’ PES improves the values of the energy levels, the TROVE calculations do not completely reproduce the observed transition frequencies. To correct this, an empirical approach was adopted, where an artificial frequency shift is added to calculated band origins as given by the $J = 0$ energies; this procedure is described as an empirical basis set correction scheme (EBSC) [18]. It leads to a rotational energy structure in much better agreement with experimental results for the remaining J values of the band. As a test, the subsequent $J = 1, 2$ and 4 values were compared to those from experiments which, together with $J = 0$ were also the values of J used to refine

the PES. While using this method, care must be taken to only select reliable experimental data, since data of limited accuracy can be a source of error which TROVE cannot compensate for.

Only 11 band centres were manipulated, and several iterative shifts were attempted. The final improvements to the standard deviation, σ , of the whole band from experiment are displayed on table 4. The original average σ value for the set of bands was 0.031 cm^{-1} which was reduced to 0.02 cm^{-1} after this final adjustment.

Table 4: Observed band origins (Obs) and standard deviation with which TROVE reproduces the terms within each band, σ , before and after replacement of the band origins. Observed data is from HITRAN2008 [7].

Band	Obs	Original σ	New σ
ν_2	992.135	0.030	0.020
ν_4	1118.307	0.012	0.005
$2\nu_2$	1972.571	0.011	0.007
$\nu_2 + \nu_4$	2108.152	0.045	0.034
$2\nu_4$	2226.835	0.029	0.010
$2\nu_4$	2234.920	0.032	0.014
ν_1	2321.121	0.028	0.012
ν_3	2326.8667	0.023	0.013
$3\nu_2$	2940.767	0.063	0.046
$\nu_2 + 2\nu_4$	3214.936	0.035	0.024
$\nu_3 + \nu_4$	3440.259	0.037	0.034

4. Results

For the room temperature line list presented here the following thresholds were selected. The chosen ranges of energy eigenvalues are 4000 cm^{-1} for the highest lower energy and 12000 cm^{-1} for the highest upper energy. These choices allow for a range of transitions of 0 to 8000 cm^{-1} at temperatures up to 300 K. The lower energy threshold of 4000 cm^{-1} defines the highest J that had to be taken into account, as there were many more energy values for each symmetry and each J than those within the selected range. By $J = 31$ only 0.0025% of all energy eigenvalues (a total of 8 levels, across all symmetries) were below 4000 cm^{-1} and by $J = 32$ there were none. Figure 1 shows the variation in number of useful (within range) eigenvalues as J increases.

Summing over all the energy levels calculated here gives a partition function equal to 3249.5 at 296 K. This is 0.028% higher than the latest published HITRAN value of 3248.6 [51]. Further work on the partition function and other thermodynamic properties of both phosphine and ammonia can be found in [59].

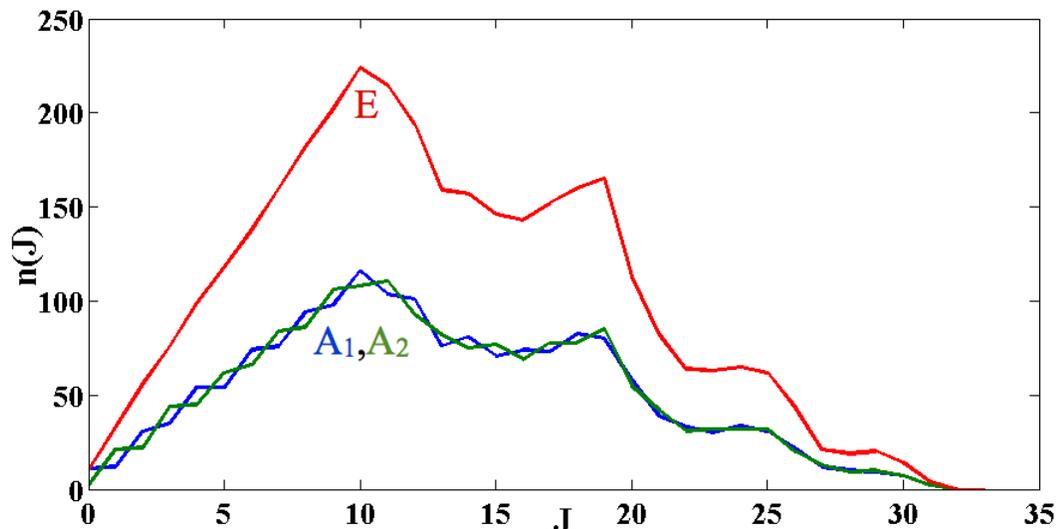


Figure 1: Total number of energy eigenvalues under 4000 cm^{-1} for $J = 0, 1, \dots, 33$. The blue, green and red lines represent, respectively A_1 , A_2 and E symmetry eigenvalues.

For the purposes of comparison with the existing experimental data, only transitions above a minimum intensity were included. In the region $0-300 \text{ cm}^{-1}$, only transitions stronger than $\geq 10^{-13} \text{ cm}^2/\text{mol}$ (or $1.7 \times 10^{-37} \text{ cm}^2/\text{molecule}$) were considered, while for transitions with wavenumbers $\geq 300 \text{ cm}^{-1}$ this threshold was reduced to $10^{-5} \text{ cm}^2/\text{mol}$ ($1.7 \times 10^{-29} \text{ cm}^2/\text{molecule}$). These are approximately one order of magnitude more sensitive than the weakest lines in both the HITRAN 2008 [7] and CDMS [53] databases. With this intensity cut-off,

only 5 488 177 transitions were selected from the total of 137 255 400 computed lines.

Figure 2 shows an overview of our simulation compared to CDMS [53] and HITRAN [7] databases. HITRAN lacks any pure rotational transitions and, although it captures most lines stronger than 5×10^{-25} cm/molecule, it is very incomplete below this value.

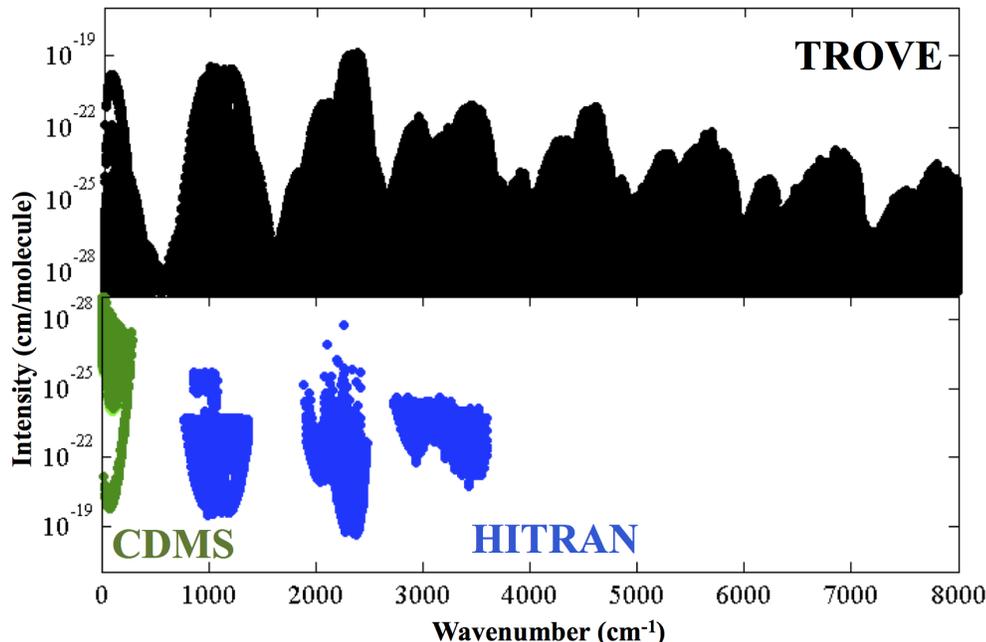


Figure 2: Comparison between simulated absorption spectrum created by TROVE and those from the JPL [6] and HITRAN [7] databases, as a function of wavenumber.

All the information on the 5.6 million rotation-vibration energy levels and the associated list of the 137 million transitions can be found online at www.exomol.com, in the ExoMol format [54]. It is possible to use this line list to generate synthetic spectra as a function of temperature. However, such spectra will become increasingly incomplete as the temperature goes above 300 K. Additionally, the temperature-dependent molecular absorption cross sections [55] for phosphine are also available online. A HITRAN format line list is included as supplementary material to this paper. An intensity cut-off of 10^{-31} cm/molecule is used for the rotational part of the spectrum ($0 - 500$ cm⁻¹) and 10^{-28} cm/molecule for the remainder of the spectrum. Additionally, the refined potential energy surface used here is provided as a fortran program.

Data for the rotational spectrum was obtained from the CDMS database [53], although the strongest of these lines can also be found in the JPL database [6, 52]. The documentation for the CDMS data can be found online and cites

its sources as Cazzoli and Puzzarini [56], Belov et al [58], Fusina and Carlotti [57], Davis, Newman, Wofsy, and Klemperer [26], Chu and Oka [25], Helms and Gordy, [27], and Belov, Burenin, Polyanski, and Shapin [22]. Further information about these papers can be found in Table 1. The CDMS data set contains 2131 transition lines in the region 0 - 300 cm^{-1} with a maximum J quantum number of 34.

To adequately compare the lines calculated here to those in CDMS, lower and upper energy levels were matched between datasets and the difference between the resulting transitions analysed. Due to the ambiguity (between A_1 and A_2) in many of the energy levels with A symmetries, only the E symmetry transitions were matched with sufficient confidence for an accurate analysis. This sample can be expected to give a representative rms for the remaining transitions. An algorithm created with MATLAB matched theoretical lines produced by TROVE with equivalent ones in the CDMS dataset. These matches deviated from those in CDMS with a rms value of 0.076 cm^{-1} . Problems with ambiguous or incorrect labelling of energy levels skew the rms deviation and its true value is expected to be significantly lower. In fact, when the top 1% worst matches are removed from the comparison, the rms deviation lowers to 0.05 cm^{-1} .

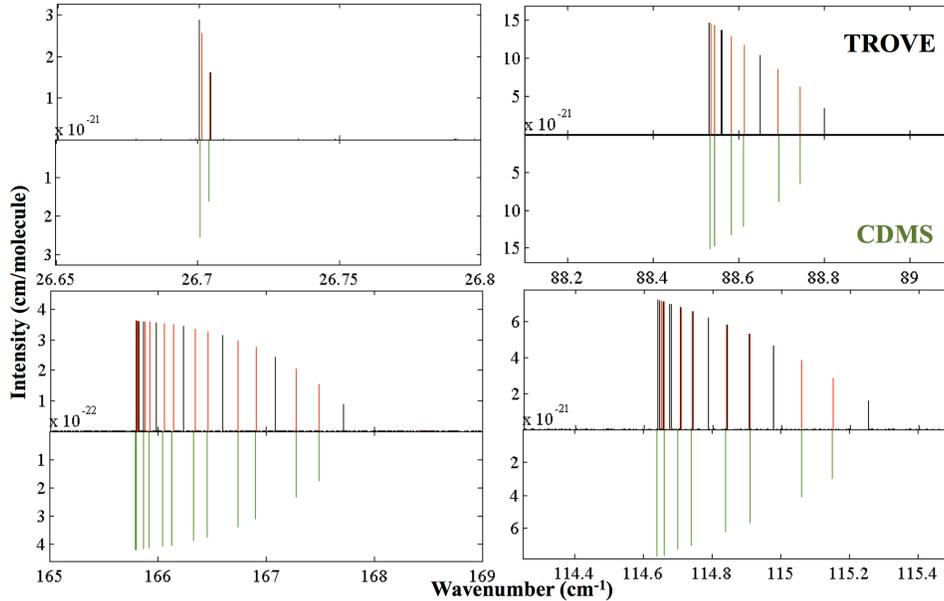


Figure 3: CDMS (lower) versus TROVE (upper). The matched TROVE transitions are highlighted in red.

Figure 3 shows close-ups of the rotational spectrum, where a line by line comparison can be made with the CDMS data mirroring TROVE’s theoretical predictions, and those transitions that could be matched highlighted in red.

The most recent HITRAN data [7] for phosphine contains 20 099 transitions in the region 770 - 3601 cm^{-1} . Of these, only 16 008 are assigned and of these 2011 have no upper vibrational quantum numbers, possibly due to vibrational mixing. There seems to be no consistent description of symmetry and many transitions appear either mislabelled or misassigned.

The energy levels of the assigned lines were compared to those calculated in TROVE and this information was used to match lines between experiment and theory. Similarly to the CDMS dataset, ambiguity between energy levels with A symmetry meant that only E symmetry transitions were considered in the comparison. Again, this sample can be expected to give representative rms for each band. Assigned E symmetry transitions correspond to 7838 of the total HITRAN lines. 13 of these transitions had ambiguous or incorrect K values, so only 7825 E transitions were matched. These were found to have a rms deviation from their experimental equivalent of 0.23 cm^{-1} . The wavenumbers of a few transitions involving energy levels which are defined by a single transition in the HITRAN database disagree significantly with that of those calculated with TROVE. These skew the rms deviation and again, its true value is expected to be significantly lower. When the top 1% worst matches are removed from the comparison, the rms deviation lowers to 0.19 cm^{-1} .

The first region (Polyad number = 1) is located between 770 - 1372 cm^{-1} and is dominated by the fundamental bending bands ν_2 and ν_4 and the "hot" band $2\nu_2 - \nu_2$. TROVE's rms deviation from HITRAN in this region is 0.23 cm^{-1} . Its rms is 0.37 cm^{-1} for the ν_2 band, 0.11 cm^{-1} for the ν_4 band and 0.11 cm^{-1} for the $2\nu_2 - \nu_2$ band. HITRAN references Brown, Sams and Kleiner [33] as the source for this region. Figure 4 shows close-ups of a selection of representative sub-regions within the $P = 1$ region.

The second region is located between 1885 - 2478 cm^{-1} and is dominated by the fundamental stretching bands ν_1 and ν_3 , the $2\nu_2$ and $2\nu_4$ fundamental overtones and the combination band $\nu_2 + \nu_4$. TROVE's rms deviation from HITRAN in this region is 0.20 cm^{-1} , with 0.28 cm^{-1} for the ν_1 band, 0.22 cm^{-1} for ν_3 , 0.05 cm^{-1} for $2\nu_2$, 0.15 cm^{-1} for $2\nu_4$ and 0.08 cm^{-1} for $\nu_2 + \nu_4$. HITRAN references Tarrago et al [1] as the source for this region. Figure 5 shows close-ups of a selection of representative sub-regions within the $P = 2$ region.

The third region is located between 2733 - 3600 cm^{-1} and is dominated by the $4\nu_2 - \nu_2$ hot band, the $3\nu_2$ overtone and the $\nu_1 + \nu_2$, $\nu_1 + \nu_4$, $\nu_2 + \nu_3$, $\nu_3 + \nu_4$, $\nu_2 + 2\nu_4$ and $2\nu_2 + \nu_4$ combination bands. TROVE's rms deviation from HITRAN in this region is 0.27 cm^{-1} , with 0.46 cm^{-1} for $4\nu_2 - \nu_2$, 0.37 cm^{-1} for $3\nu_2$, 0.50 cm^{-1} for $\nu_1 + \nu_2$, 0.21 cm^{-1} for $\nu_1 + \nu_4$, 0.23 cm^{-1} for $\nu_2 + \nu_3$, 0.19 cm^{-1} for $\nu_3 + \nu_4$, 0.11 cm^{-1} for $\nu_2 + 2\nu_4$ and 0.28 cm^{-1} for $2\nu_2 + \nu_4$. HITRAN references Butler et al [40] as the source for this region. Figure 6 shows close-ups of a selection of representative sub-regions within the $P = 3$ region.

Nikitin et al [16] computed 55 223 transitions in the range 700 - 3500 cm^{-1}

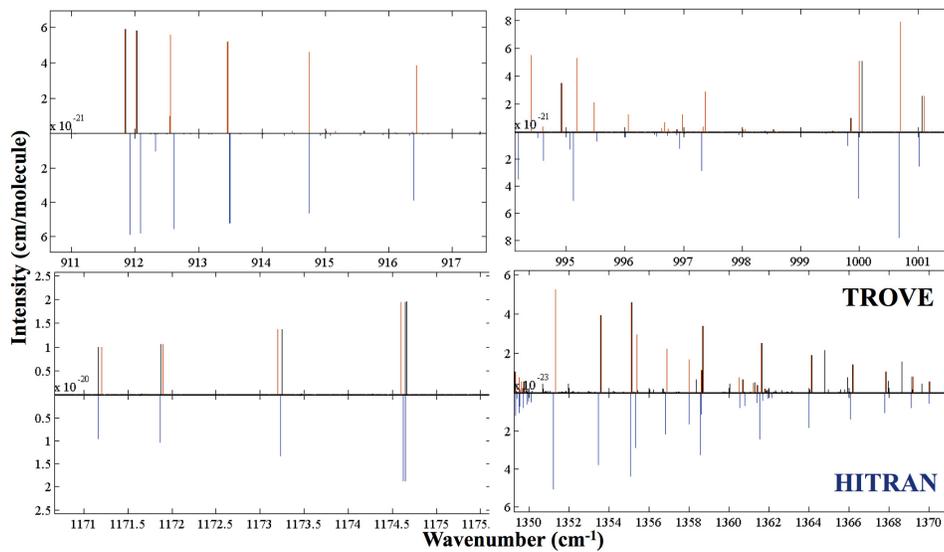


Figure 4: HITRAN (lower) versus TROVE (upper), with matched transitions highlighted in red, for the region $P = 1$.

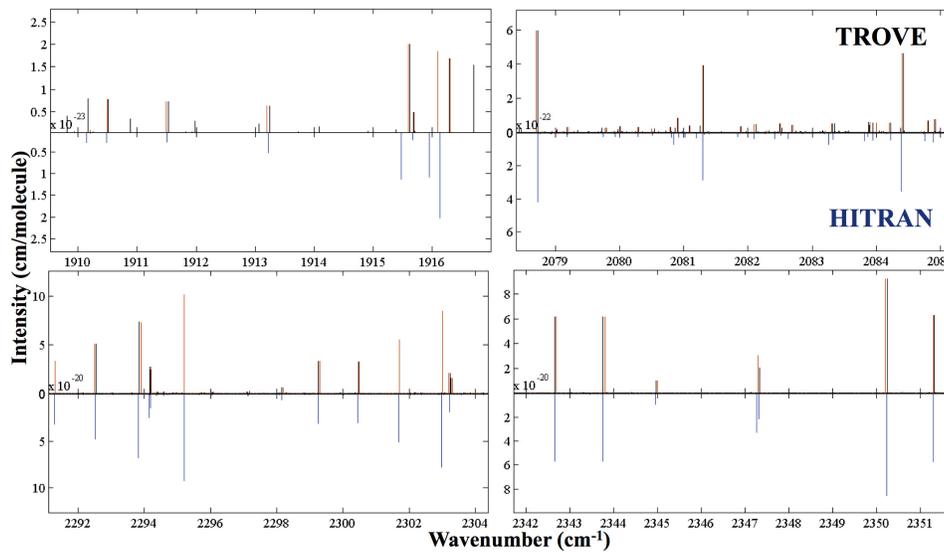


Figure 5: HITRAN (lower) versus TROVE (upper), with matched transitions highlighted in red, for the region $P = 2$.

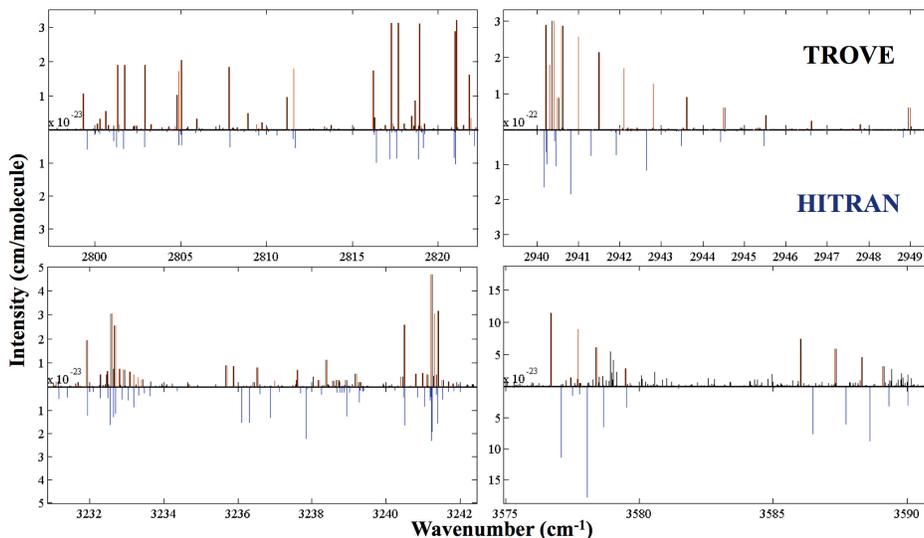


Figure 6: HITRAN (lower) versus TROVE (upper), with matched transitions highlighted in red, for the region $P = 3$.

(the HITRAN range [7]). This work does not aim for completeness and even within the HITRAN regions appears to have significant omissions (e.g. maximum $J = 20$). The work presented here improves on their root mean squares deviation from observed band centres of 1.4 cm^{-1} .

Experimental data (1768 lines) from Wang et al [37] in the $1950\text{--}2480 \text{ cm}^{-1}$ and $3280\text{--}3580 \text{ cm}^{-1}$ regions was compared to the theoretical lines created by TROVE. The results presented here deviated from those of Wang et al's with an overall rms deviation of 0.17 cm^{-1} , for those lines that could be matched with confidence. The rms deviation for the first region was 0.11 cm^{-1} , compared to 0.20 cm^{-1} for the equivalent HITRAN region and 0.21 cm^{-1} for the second, compared to 0.27 cm^{-1} for HITRAN.

Additionally, our intensities agree significantly better with the Wang data than with HITRAN's data, as can be seen by the intensity plots in Figure 7, suggesting that Wang et al's intensities should be used in a future release of HITRAN.

The JPL data [52, 6] agrees very well with our calculations. It contains 729 transitions between $0 - 188 \text{ cm}^{-1}$ with $J \leq 27$. Energy level comparison of the E symmetry transitions showed that the data presented here deviated from JPL with a rms of 0.066 cm^{-1} . When the top 1% worst matches are removed from the comparison, the rms deviation lowers to 0.031 cm^{-1} . Overall, the JPL transition wavenumbers agree more closely with those presented here than CDMS'. However, the hundred strongest transitions in the rotational spectrum from CDMS match ours much better (0.003 cm^{-1} compared to JPL's 0.07

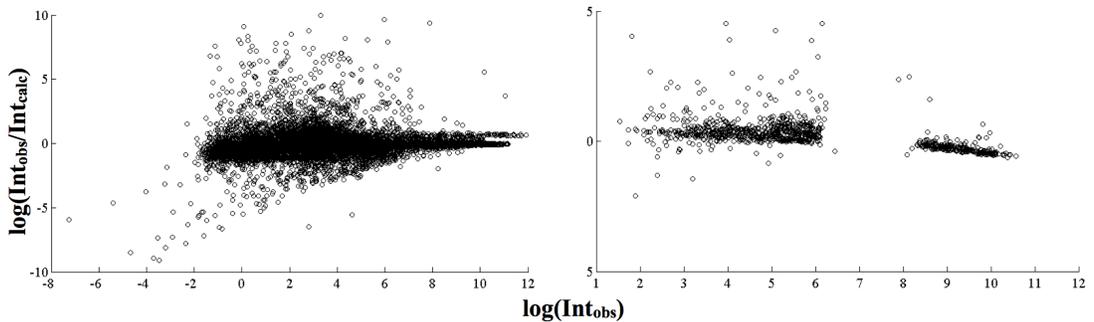


Figure 7: Plot of $\log(I_{obs}/I_{calc})$ versus $\log(I_{obs})$ for the HITRAN data [7] (left) and the Wang data [37].

cm^{-1}). Additionally, CDMS is considerably more complete and its intensities are in closer agreement with ours. This should at least partially be due to the fact that some of the $K=3n$ transition doublets are very close so their intensities are combined in the JPL database, while the equivalent TROVE lines are left unresolved. This difference in data handling leads to an apparent factor of two disagreement in the intensities.

5. Discussion and Conclusion

In this work a phosphine line list is produced with 137 million transitions between 5.6 million energy levels for ro-vibrational states up to $J_{max} = 31$ and energies up to 8000 cm^{-1} . This work replicates very well the observed phosphine spectra at room temperature, with a maximum rms deviation from CDMS of 0.076 cm^{-1} for the rotational spectrum and of 0.23 cm^{-1} from HITRAN. This is a valid line list for any phosphine analysis below 300 K, however our ultimate aim is to construct a hot line list capable of simulating observed spectra from astronomical bodies at higher temperatures, e.g. 2000 K. This line list will complement the one already available for ammonia [46]. Work on this is currently in progress.

The tunnelling effect present in the ammonia molecule is predicted to be found in phosphine [22] but, due to its much higher barrier ($12\,300 \text{ cm}^{-1}$), yet to be observed. The value of splitting in various vibrational states as well as the intensity of the inversion-rotation, and inversion-rovibrational lines can be computed by adapting the procedure given here to work with $D_{3h}(M)$ symmetry. Predictions for transitions which can be used to resolve the doublet splitting will be presented in future work.

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References

- [1] G. Tarrago, N. Lacome, A. Levy, G. Guelachvili, B. Bezaud, P. Drossart, *J. Mol. Spectrosc.* 154 (1992) 30–42.
- [2] L. N. Fletcher, G. S. Orton, N. A. Teanby, P. G. J. Irwin, *Icarus* 202 (2009) 543–564.
- [3] R. G. Prinn, J. S. Lewis, *Science* 190 (1975) 274–276.
- [4] S. G. Edgington, S. K. Atreya, L. M. Trafton, J. J. Caldwell, R. F. Beebe, A. A. Simon, R. A. West, C. Barnet, *Icarus* 133 (1998) 192–209.
- [5] R. Beer, F. W. Taylor, *Icarus* 40 (1979) 189–192.
- [6] H. M. Pickett, R. L. Poynter, E. A. Cohen, M. L. Delitsky, J. C. Pearson, H. S. P. Müller, *J. Quant. Spectrosc. Radiat. Transf.* 60 (1998) 883–890.
- [7] L. S. Rothman, I. E. Gordon, A. Barbe, D. C. Benner, P. F. Bernath, M. Birk, V. Boudon, L. R. Brown, A. Campargue, J. P. Champion, K. Chance, L. H. Coudert, V. Dana, V. M. Devi, S. Fally, J. M. Flaud, R. R. Gamache, A. Goldman, D. Jacquemart, I. Kleiner, N. Lacome, W. J. Lafferty, J. Y. Mandin, S. T. Massie, S. N. Mikhailenko, C. E. Miller, N. Moazzen-Ahmadi, O. V. Naumenko, A. V. Nikitin, J. Orphal, V. I. Perevalov, A. Perrin, A. Predoi-Cross, C. P. Rinsland, M. Rotger, M. Simeckova, M. A. H. Smith, K. Sung, S. A. Tashkun, J. Tennyson, R. A. Toth, A. C. Vandaele, J. Vander Auwera, *J. Quant. Spectrosc. Radiat. Transf.* 110 (2009) 533–572.
- [8] J. Tennyson, S. N. Yurchenko, *Mon. Not. R. Astr. Soc.* 425 (2012) 21–33.
- [9] D. Wang, Q. Shi, Q.-S. Zhu, *J. Chem. Phys.* 112 (2000) 21.
- [10] S. G. He, J. J. Zheng, S. M. Hu, H. Lin, Y. Ding, X. H. Wang, Q. S. Zhu, *J. Chem. Phys.* 114 (2001) 7018–7027.
- [11] S. N. Yurchenko, M. Carvajal, P. Jensen, F. Herregodts, T. R. Huet, *Chem. Phys.* 290 (2003) 59–67.

- [12] S. N. Yurchenko, W. Thiel, S. Patchkovskii, P. Jensen, *Phys. Chem. Chem. Phys.* 7 (2005) 573–582.
- [13] S. N. Yurchenko, M. Carvajal, W. Thiel, P. Jensen, *J. Mol. Spectrosc.* 239 (2006) 71–87.
- [14] R. I. Ovsyannikov, W. Thiel, S. N. Yurchenko, M. Carvajal, P. Jensen, *J. Chem. Phys.* 129 (2008) 044309.
- [15] R. I. Ovsyannikov, W. Thiel, S. N. Yurchenko, M. Carvajal, P. Jensen, *J. Mol. Spectrosc.* 252 (2008) 121–128.
- [16] A. V. Nikitin, F. Holka, V. G. Tyuterev, J. Fremont, *J. Chem. Phys.* 130 (2009) 244312.
- [17] A. V. Nikitin, J. P. Champion, R. A. H. Butler, L. R. Brown, I. Kleiner, *J. Mol. Spectrosc.* 256 (2009) 4–16.
- [18] S. N. Yurchenko, R. J. Barber, A. Yachmenev, W. Thiel, P. Jensen, J. Tennyson, *J. Phys. Chem. A* 113 (2009) 11845–11855.
- [19] P. R. Bunker, P. Jensen, *Molecular Symmetry and Spectroscopy*, NRC, Canada, 1998.
- [20] M. Carlotti, G. Di Lonardo, A. Trombetti, *J. Chem. Phys.* 78 (1983) 1670–1673.
- [21] G. Tarrago, M. Dang-Nhu, A. Goldman, *J. Mol. Spectrosc.* 88 (1981) 311–322.
- [22] S. P. Belov, A. V. Burenin, O. L. Polyansky, S. M. Shapin, *J. Mol. Spectrosc.* 90 (1981) 579–589.
- [23] P. Schwerdtfeger, L. J. Laakkonen, P. Pyykkö, *J. Chem. Phys.* 96 (1992) 6807–6819.
- [24] M. J. Down, C. Hill, S. N. Yurchenko, J. Tennyson, L. R. Brown, I. Kleiner, *J. Quant. Spectrosc. Radiat. Transf.* (2013). (submitted).
- [25] F. Y. Chu, T. Oka, *J. Chem. Phys.* 60 (1974) 4612–4618.
- [26] P. B. Davies, R. M. Neumann, S. C. Wofsy, W. Klemperer, *J. Chem. Phys.* 55 (1971) 3564–&.
- [27] D. A. Helms, W. Gordy, *J. Mol. Spectrosc.* 66 (1977) 206–218.
- [28] C. C. Loomis, M. W. P. Strandberg, *Phys. Rev.* 81 (1951) 798–807.
- [29] H. M. Pickett, R. L. Poynter, E. A. Cohen, *J. Quant. Spectrosc. Radiat. Transf.* 26 (1981) 197–198.
- [30] P. Helminger, W. Gordy, *Phys. Rev.* 188 (1969) 100–107.

- [31] A. F. Krupnov, A. A. Melnikov, V. A. Skvortsov, *Optika I Spektroskopiya* 46 (1979) 1012–1013.
- [32] L. Fusina, G. Di Lonardo, *J. Mol. Spectrosc.* 517 (2000) 67–78.
- [33] L. R. Brown, R. L. Sams, I. Kleiner, C. Cottaz, L. Sagui, *J. Mol. Spectrosc.* 215 (2002) 178–203.
- [34] A. Ainetschian, U. Haring, G. Spiegl, W. A. Kreiner, *J. Mol. Spectrosc.* 181 (1997) 99–107.
- [35] J. Salem, H. Aroui, J. Bouanich, J. Walrand, G. Blanquet, *J. Mol. Spectrosc.* 225 (2004) 174–181.
- [36] O. N. Ulenikov, E. S. Bekhtereva, V. A. Kozinskaia, J. J. Zheng, S. G. He, S. M. Hu, Q. S. Zhu, C. Leroy, L. Pluchart, *J. Mol. Spectrosc.* 215 (2002) 295–308.
- [37] L. Wang, P. Chen, G. S. Cheng, Y. Ding, S. M. Hu, *Spectroscopy and Spectral Analysis* 25 (2005) 1221.
- [38] A. Baldacci, V. Malathy Devi, K. Narahari Rao, G. Tarrago, *J. Mol. Spectrosc.* 81 (1980) 179–206.
- [39] C. B. Suarez, *Spectroscopy letters* 35 (2002) 757–767.
- [40] R. A. H. Butler, L. Sagui, I. Kleiner, L. R. Brown, *J. Mol. Spectrosc.* 238 (2006) 178–192.
- [41] R. J. Kshirsagar, *J. Mol. Spectrosc.* 241 (2007) 116–118.
- [42] A. G. Maki, R. L. Sams, W. B. Olson, *J. Chem. Phys.* 58 (1973) 4502.
- [43] O. N. Ulenikov, E. S. Bekhtereva, V. A. Kozinskaia, J. J. Zheng, S. G. He, S. M. Hu, Q. S. Zhu, C. Leroy, L. Pluchart, *J. Quant. Spectrosc. Radiat. Transf.* 83 (2004) 599–618.
- [44] S. N. Yurchenko, W. Thiel, P. Jensen, *J. Mol. Spectrosc.* 245 (2007) 126–140.
- [45] S. N. Yurchenko, M. Carvajal, P. Jensen, H. Lin, J. Zheng, W. Thiel, *Mol. Phys.* 103 (2005) 359–378.
- [46] S. N. Yurchenko, R. J. Barber, J. Tennyson, *Mon. Not. R. Astr. Soc.* 413 (2011) 1828–1834.
- [47] S. N. Yurchenko, R. J. Barber, J. Tennyson, W. Thiel, P. Jensen, *J. Mol. Spectrosc.* 268 (2011) 123–129.
- [48] A. Yachmenev, S. N. Yurchenko, P. Jensen, W. Thiel, *J. Chem. Phys.* 134 (2011) 244307.

- [49] B. Numerov, *Mon. Not. R. Astron. Soc.* 84 (1924) 592–xxx.
- [50] J. W. Cooley, *Math. Comp.* 15 (1961) 363–374.
- [51] M. Šimečková, D. Jacquemart, L. S. Rothman, R. R. Gamache, A. Goldman, *J. Quant. Spectrosc. Radiat. Transf.* 98 (2006) 130–155.
- [52] J. C. Pearson, H. S. P. Müller, H. M. Pickett, E. A. Cohen, B. J. Drouin, *J. Quant. Spectrosc. Radiat. Transf.* 111 (2010) 1614–1616.
- [53] H. S. P. Müller, F. Schlöder, J. Stutzki, G. Winnewisser, *J. Molec. Struct. (THEOCHEM)* 742 (2005) 215–227.
- [54] J. Tennyson, C. Hill, S. N. Yurchenko, *AIP Conference Proceedings*, AIP, New York, 2013.
- [55] C. Hill, S. N. Yurchenko, J. Tennyson, *Icarus* (2013).
- [56] G. Cazzoli, C. Puzzarini, *J. Mol. Spectrosc.* (2006).
- [57] L. Fusina, M. Carlotti, *J. Mol. Spectrosc.* (1988).
- [58] S. P. Belov, A. V. Burenin, L. I. Gershtein, A. F. Krupnov, V. N. Markov, A. V. Maslovsky, S. M. Shapin, *J. Mol. Spectrosc.* (1981).
- [59] C. Sousa-Silva, N. Hesketh, S. N. Yurchenko, C. Hill, J. Tennyson, *J. Quant. Spectrosc. Radiat. Transf.*(submitted) (2013).