

The Economic Impact of Space Weather: Where Do We Stand?

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Space weather describes the way in which the Sun, and conditions in space more generally, impact human activity and technology both in space and on the ground. It is now well understood that space weather represents a significant threat to infrastructure resilience, and is a source of risk that is wide-ranging in its impact and the pathways by which this impact may occur. Although space weather is growing rapidly as a field, work rigorously assessing the overall economic cost of space weather appears to be in its infancy. Here, we provide an initial literature review to gather and assess the quality of any published assessments of space weather impacts and socioeconomic studies. Generally speaking, there is a good volume of scientific peer-reviewed literature detailing the likelihood and statistics of different types of space weather phenomena. These phenomena all typically exhibit “power-law” behavior in their severity. The literature on documented impacts is not as extensive, with many case studies, but few statistical studies. The literature on the economic impacts of space weather is rather sparse and not as well developed when compared to the other sections, most probably due to the somewhat limited data that are available from end-users. The major risk is attached to power distribution systems and there is disagreement as to the severity of the technological footprint. This strongly controls the economic impact. Consequently, urgent work is required to better quantify the risk of future space weather events.

KEY WORDS: Geomagnetic storms; power grids; space weather

1. INTRODUCTION

Space weather is of rising importance both as a scientific discipline in its own right^(1–8) and as a severe source of risk recognized by governmental agencies and corporations at the national and international level.^(9–19) For example, in the United States, it has been the subject of a recent Executive Order issued by President Barack Obama, which directs multiple federal agencies and departments to coordinate their preparation for, and response to, severe space weather.⁹ This highlights the fact that space weather is a fundamentally interdisciplinary risk, and has the potential to affect myriad technologies and activities in space and on the ground.

⁹<https://www.whitehouse.gov/the-press-office/2016/10/13/executive-order-coordinating-efforts-prepare-nation-space-weather-events>

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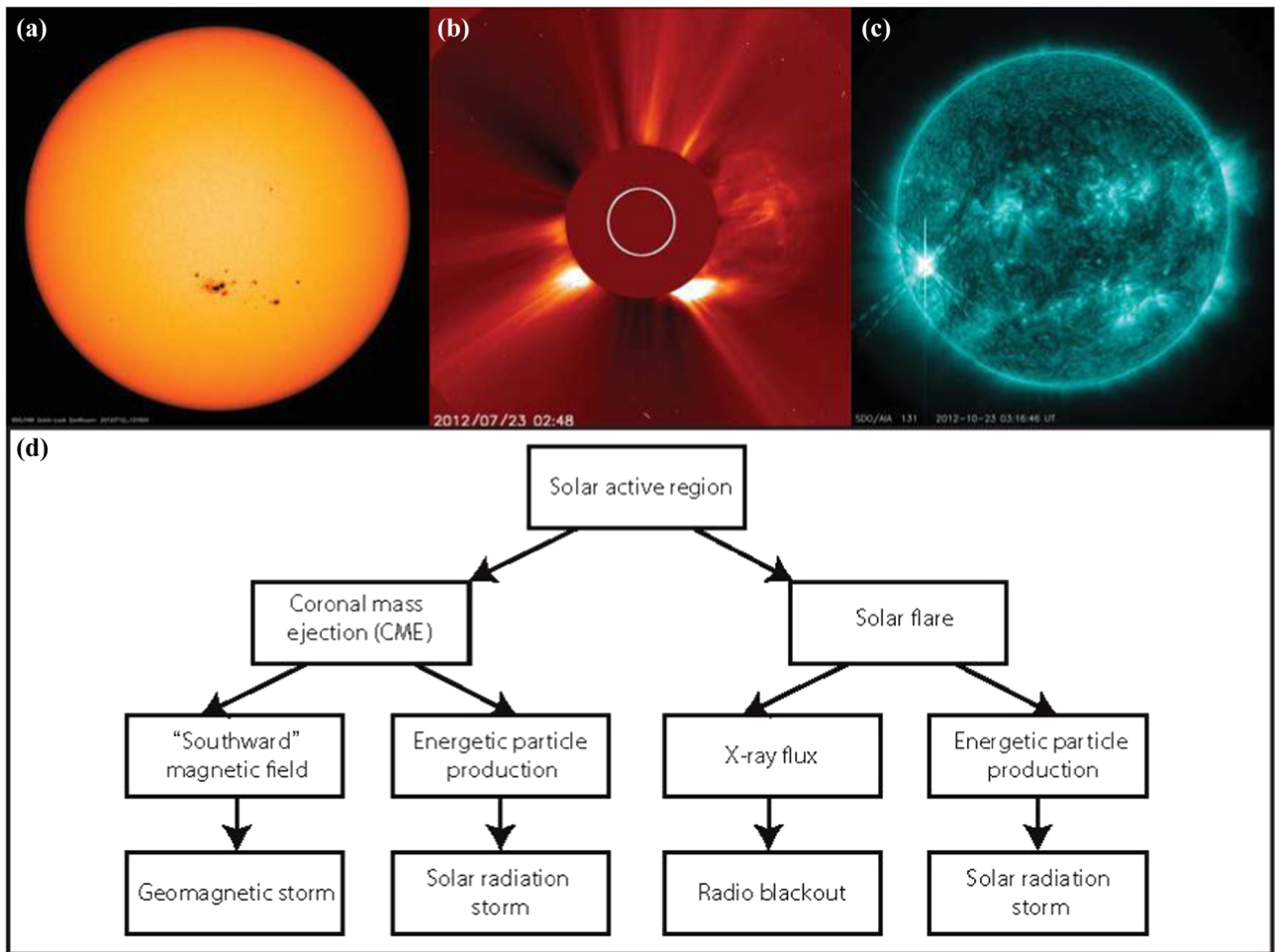


Fig. 1. (a) Image of sunspots on the solar disk from NASA's Solar Dynamics Observatory on July 12, 2012. The grouping of sunspots was associated with active region 1520. (b) CME launched from active region 1520 on July 23, 2012 and observed by the European Space Agency/NASA Solar and Heliospheric Observatory (SOHO) (Image credit ESA/NASA/SOHO). (c) X-class flare observed by NASA's Solar Dynamics Observatory on October 22, 2012 at 131 Angstrom wavelength, which caused an R3 radio blackout (Image credit: NASA/SDO/Goddard). (d) Solar active regions produce CMEs and solar flares, which are subsequently responsible for the three primary categories of space weather as indicated here and in the text.

Although the field is growing rapidly, peer-reviewed work rigorously establishing the economic impact of space weather is still scarce. Here, we aim to briefly summarize what is known to date in this area. We first introduce the physical nature of the risk and notable space weather events, in order to establish the hazard component (event occurrence). We then review documented impacts in several different sectors, (i.e., the loss-severity component), and finally examine attempts to quantify the economic consequences. We conclude with some observations about the current status, the research gaps, and the

main challenges characterizing this area of risk analysis.

1.1. Space Weather

The ultimate source of space weather is the Sun. It can sometimes produce bursts of electromagnetic radiation (*flares*) and eruptions of material (coronal mass ejections, *CMEs*) accompanied by solar energetic particles (*SEPs*). The interaction of CMEs with the Earth's magnetic field can lead to a *geomagnetic storm* (see Fig. 1). Three main types of space weather

Table I. Types of Space Weather and Their Properties

Space Weather Category	Physical Cause	Physical Measure	Timing	Impact	Likelihood of Most Severe Event
Radio blackout R1 least severe R5 most severe	Solar flares: X-ray emission causes increased low altitude ionization (60–90 km). UV emission heats and expands atmosphere.	Solar X-ray flux	The electromagnetic emission travels at light speed—no warning time. Impact is almost immediate, but depends on ionospheric response.	Absorption and disruption of high-frequency (HF) radio (3–30 MHz).	The most severe event (R5) would correspond to a GOES peak X-ray flux exceeding X20 (i.e., $20 \times 10^{-4} \text{ Wm}^{-2}$). Such an event is expected to occur less than once a solar cycle.
Solar radiation storm S1 least severe S5 most severe	Solar energetic particles (SEPs) accelerated by flares and coronal mass ejections (CMEs). ^(20,21)	Energetic particle flux	During “impulsive events” associated with flares, and the launch of CMEs, fluxes can start to rise significantly a few minutes after the arrival of the electromagnetic emission. Gradual events associated with CME propagation through interplanetary space intensify more slowly.	Biological effects on astronauts, aircrew, and airline passengers (if sufficiently energetic). Technological impact on electronic devices on satellites, aircraft, and even on the ground. HF radio blackout in polar regions.	An S5 event is expected to occur less than once a solar cycle. (However, note that S5 is not useful for aviation impact because particles are still insufficiently energetic to penetrate down to altitudes of interest, and so special aviation alerts also exist in addition to S5 warnings.)
Geomagnetic storm G1 least severe G5 most severe	Primarily triggered by earthward-directed coronal mass ejections that travel through space from the Sun. A geomagnetic storm will develop when a CME interacts with the Earth’s magnetic field, provided that the CME magnetic field structure is appropriately oriented “southward.” Solar wind high-speed streams can also drive geomagnetic storms if their magnetic field is similarly oriented. Fast CMEs with strong magnetic fields produce the largest storms. ^(22–24) Although geomagnetic storms are a global phenomenon, their regional impact is highly heterogeneous. Geomagnetic storms cause geomagnetically induced currents (GICs) on the surface of the Earth.	Geomagnetic index ^a	A geomagnetic storm may arise after two to four days depending on the time it takes for the CME to travel from the Sun to the Earth. The severity of each particular geomagnetic storm impact can vary considerably in space and time. It is currently not possible to predict if a CME will be geoeffective until it is measured by satellites just before it reaches the Earth.	Power system failure, spacecraft failure, degraded satellite navigation, severe ionospheric disturbances, GNSS degradation (e.g., affecting aviation augmentation systems) and loss of integrity, and HF communication blackout. Geomagnetic storms are considered a primary risk in most assessments of space weather impact.	Storm severity is typically measured according to the <i>Kp</i> index ($Kp_{\text{max}} = 9$), which is designed to capture the severity of the global perturbation to the quiet magnetic field on the surface of the Earth (time resolution of three hours). The physical measure of a G5 event is $Kp = 9$ and, on average, four such storm days per solar cycle are expected.

^aGeomagnetic storms are difficult to characterize. Geomagnetic indices condense the properties of a complex nonlinear system with a variety of dynamical processes into a single number. They provide information about the occurrence of geomagnetic storms on a global level, and therefore are a necessary part of understanding geomagnetic storm likelihood. They must be used with care:⁽²⁾ there are examples of little or no variation in geomagnetic index even when the space weather is “stormy,”⁽²⁵⁾ and relatively weak storms can have significant effects on technological systems.⁽²⁶⁾ Furthermore, none of the indices currently in use are directly useful for understanding the most important societal impact—the production of GICs that may affect power grids. Such GIC production is highly inhomogeneous in space and time. There is increased interest in understanding the statistics of extreme events on regional scales,⁽²⁷⁾ and in the future, it is likely that a global measure of geomagnetic storms will be rendered obsolete in favor of local warnings of specific storm impact.⁽²⁸⁾

Table II. Notable Space Weather Events

Date	Comment
September 1859	The “Carrington” event is the benchmark for extreme space weather studies. ^(29–33) The solar flare, the geomagnetic storm, and the energetic particle flux associated with this event make it one of the largest on record. ⁽³⁰⁾ Note that many crucial parameters were not measured directly, so its precise properties are subject to uncertainty. In particular, estimating the strength of the geomagnetic storm associated with the Carrington event has attracted some debate; initial estimates ⁽³⁴⁾ should be disregarded in favor of more recent analysis. ^(35,36)
May 1921	This geomagnetic storm has been estimated to be comparable in size to the current best estimate of the Carrington event. ^(37,38) Auroras were seen near the equator in Samoa, ⁽³⁰⁾ and geomagnetically induced currents (GICs) caused fires at several telegraph stations in Sweden. ⁽³⁹⁾
May 1967	An extreme solar flare and coronal mass ejection caused very significant radio blackouts, solar radiation storms, and a major geomagnetic storm. This caused a particularly significant disruption to communications, specifically to the military, and marked the start of a significant U.S. investment in space weather monitoring that continues to this day. ⁽⁴⁰⁾
March 1989	The largest geomagnetic storm of the space age ⁽⁴¹⁾ causing well-known failure of the Quebec power grid ⁽⁴²⁾ and damaging two transformers in the United Kingdom. ⁽⁴³⁾
October–November 2003	Very well-observed and measured complex series of events including one of the largest observed solar flares on record. ^(44,45) The overall technological impact is extremely well documented. ^(46–49) A 90-minute blackout in 2003 affected 50,000 customers in Sweden. (Although it is now widely recognized that this blackout would probably have been avoided if current operational warning systems had been in place). ⁽⁵⁰⁾
July 2012	This CME was not Earth directed, but was measured <i>in situ</i> by the STEREO-A spacecraft. ⁽⁵¹⁾ If this CME had been Earth directed, it would have generated a very severe “Carrington class” geomagnetic storm. ^(52,53) It has been argued that this event should be used to create severe space weather scenarios for planning purposes. ⁽⁵²⁾

are typically recognized: *radio blackouts*, *solar radiation storms*, and *geomagnetic storms* (Table I), and are monitored by various national and international agencies.¹⁰ Since the underlying physical drivers are complex and highly interconnected, all three phenomena can occur multiple times within one space weather “event,” with varying temporal and spatial footprints, as well as levels of severity. Table II summarizes the notable space weather events that will be referred to throughout the article.

1.2. Statistics of Space Weather Events and Severe Event Likelihood

All relevant space weather phenomena—solar flare intensity,^(54–56) CME speeds,^(55,57) and geomagnetic storm strength^(55,58)—typically follow power-law distributions.⁽⁵⁹⁾ Most studies therefore focus on the study of tail indices, although more complex statistical models have also been developed.^(60,61) Solar flare statistics in particular have been interpreted in terms of self-organized criticality. In such models, continual small changes in the evolving magnetic field of the Sun’s atmosphere (the corona) are thought to trigger periodic energy release events (flares) whose size follows a power-law distribution,

analogous to avalanches occurring on the surface of a sand-pile where grains are continually added.^(62–68) Poisson statistics (modulated by the solar cycle) are the standard framework for establishing the waiting time distribution of flares,⁽⁶⁹⁾ CMEs,⁽⁷⁰⁾ and geomagnetic storms.^(71,72) However, there is some evidence for clustering of CME eruptions.⁽⁵⁷⁾

Although there is a growing realization that vulnerability arises not simply due to low-frequency and high-impact events, but also due to continuing degradation as a consequence of many smaller impacts,⁽¹⁸⁾ understanding the *most severe event* that might occur is crucial for disaster planning scenarios.⁽⁷³⁾ The largest *solar flare* ever recorded in satellite data was on November 4, 2003 (see Table II).^(30,45) Given the rarity of very large solar flares, analysis of Sun-like stars using, e.g., NASA’s Kepler spacecraft^(74,75) suggests that superflares ($\sim 10\times$ Carrington) may occur on millennial timescales, but this is still controversial.^(76–79) On this basis, the probability of a flare in the next 30 years whose strength broadly exceeds that observed in 2003 is about 10%.⁽⁵⁶⁾

Direct measurement of *extreme solar radiation storms* is limited to the “space age,”⁽⁸⁰⁾ the largest observed being in August 1972.⁽²⁰⁾ Statistics are therefore very limited. So-called ground-level events in neutron monitor data provide a somewhat longer proxy data set,^(81,82) but their geographic variability is unpredictable. Polar ice core nitrate concentrations are no longer considered to be a reliably proxy of

¹⁰See, for example, the World Meteorological Organization’s list of national and international agencies that provide space weather services: http://www.wmo.int/pages/prog/sat/spaceweather-catalogue_en.php.

event intensity.^(56,83) Alternative proxies, Carbon-14 (in tree rings) and Beryllium-10 (in polar ice cores), are of current interest,^(84,85) strongly suggesting an intense global atmospheric radiation event occurred around AD775^(86–89) due to the Sun.^(87,90–92)

Geomagnetic storm statistics are more complex because likelihood depends on both the solar wind driver (typically, but not exclusively, a CME) and the magnetospheric response. The probability of a Carrington-like event occurring in the next decade is estimated to be 12%⁽⁵⁵⁾ (50% in the next 50 years).⁽¹⁷⁾ The probabilities of a superstorm event (worse than Carrington) and a 1989 event are calculated as $0 < 6.3\% < 23\%$ and $3.4\% < 17.8\% < 38.6\%$ in the next 10 years, respectively (95.4% Bayesian C2 confidence interval).⁽⁵⁸⁾ Finally, although weaker storms are correlated with the strength of the solar cycle, strong storms are not, and so could arise even in epochs where the Sun was quieter, as was the case for both the Carrington event and the 2012 event.⁽⁹³⁾

2. SPACE WEATHER: DOCUMENTED IMPACTS

2.1. Power Grids

Geomagnetically induced currents (GICs)⁽⁹⁴⁾ associated with geomagnetic storms may damage physical infrastructure (specifically transformers), introduce voltage instabilities that can lead to a black-out without infrastructure damage, and interfere with protection systems and fault detection.^(17,35,95–97) It is important to note that the ionospheric current systems that couple to GICs are very structured, and are most intense at relatively high latitudes in the vicinity of the auroral ovals. The aurora and associated current systems descend in latitude during strong geomagnetic storms. Consequently, impacts are not restricted to high latitudes^(48,98–101) and have been documented in the United Kingdom,^(17,43,102) Finland,^(95,103) Sweden,⁽⁵⁰⁾ Spain,⁽¹⁰⁴⁾ the United States and Canada,⁽⁴²⁾ South Africa,^(100,105,106) Japan,⁽¹⁰⁷⁾ China,⁽¹⁰⁸⁾ and Brazil.⁽¹⁰⁹⁾ The impact of geomagnetic storms on the North American power grid has been the subject of multiple reviews.^(110–113) For example, 4% of the disturbances between 1992 and 2010 reported to the U.S. Department of Energy are attributable to strong geomagnetic activity.⁽¹¹⁴⁾ Recent technical assessments in the United States and the United Kingdom find that the most likely

impact is system collapse due to voltage instability with some transformer damage.^(17,112)

2.2. Oil and Gas Industry

GICs can cause changes in pipe to soil voltage that drive enhanced corrosion.⁽¹¹⁵⁾ Aeromagnetic surveys and precision drilling are affected by magnetic fluctuations during geomagnetic storms.¹¹ However, it has proven difficult to obtain information on specific documented impacts from anywhere within these industries.

2.3. Communications

Mobile network performance can be affected by solar flare radio noise;⁽¹¹⁶⁾ these effects are hard to discern among various other variables controlling service quality.⁽¹⁷⁾ Certain mobile networks may be affected by the loss of global navigation satellite system (GNSS) timing information. *Short-wave, high-frequency (HF) radio* is used by aviation and shipping, as well as the military.^(17,46,117) During geomagnetic storms, regional and global reductions in the operational HF band occur. Modern HF systems are designed to be resilient, but legacy systems may experience outages. During a Carrington event, HF communication performance could be affected for several days.⁽¹⁷⁾ *Optical fiber networks* require repeater stations to periodically boost the signal; associated power infrastructure is at risk to GICs.^(118,119)

2.4. Ground Transportation

Rail networks are in principle susceptible to GICs.^(120,121) There is potentially considerable economic benefit to the rail industry in the use of space—e.g., for signaling, communications, monitoring, and Earth observation (landslides, etc.).⁽¹²²⁾ However, a substantive issue is how geomagnetic storms could interfere with the electromagnetic environment along the railway, including safety critical systems. *Trams and light railways* may be similarly affected, and all mass transit would be severely impacted by power loss (especially for underground mass transit). Finally, a more speculative space weather impact in the future is that on *driverless cars* and *road charging* based on GNSS.

¹¹See, for example, http://geomag.bgs.ac.uk/data_service/directionaldrilling/home.html.

2.5. Satellite Infrastructure

Satellites are at risk from the space environment.^(123,124) Energetic electrons trapped in the outer radiation belt cause electrostatic charging and discharging, which can damage sensitive electronic equipment and solar panels.^(125–129) SEPs can cause displacement damage (reducing device performance) and single event effects (SEEs),^(130–132) which are a growing issue as devices are miniaturized.⁽¹⁷⁾ During the 2003 Halloween storms, 47 satellites reported anomalies (out of 450 in orbit, i.e., ~10%), one scientific satellite was lost, and 10 satellites lost operational service for more than one day.^(17,46,47) Complete losses have thus been rare since satellites are designed to tolerate a total dose over some lifetime, with good safety margins: temporary outages and fleet aging are both more likely.⁽¹⁷⁾

2.6. Global Navigation Satellite Systems: Disruption to Service

Space weather causes signal distortion (scintillation and loss of lock) in the ionosphere and does not have a significant impact on GNSS satellites themselves.^(17,133) Disruption to *positioning and timing* services would occur during a major space weather event, affecting many sectors (e.g., communications, financial trading, energy networks, etc.). Augmented GNSS systems (e.g., EGNOS and WAAS aviation systems⁽¹⁷⁾) may be particularly vulnerable when very large geomagnetic storms cause signal scintillation and physical differences between the conditions at the receiver and the reference station. During a major storm, complete loss of GNSS service for one day is estimated, with extended loss of service for three days.⁽¹⁷⁾ Although many systems can revert to backup technologies, the impact of the reduced accuracy over a prolonged multiday outage is not well understood or verified.

2.7. Aviation

Solar radiation storms enhance the cosmic-ray-generated radiation environment at flight altitude.^(134–137) A perhaps counterintuitive effect is that energetic particle radiation can diminish during/after a geomagnetic storm (a Forbush decrease) because the CME can block Galactic Cosmic Rays, which leads to a complex balance of effects.⁽¹³⁸⁾ Radiation storms have a technological^(134,139–141) and

biological impact (due to the fact that the radiation is ionizing).^(142,143) While unlikely, mandated crew dose limits could be reached in part due to severe space weather;⁽¹⁴¹⁾ the wider impact has also been examined.⁽¹⁷⁾ Reduced flight time at high altitude may be required should a severe energetic particle event to occur during flight,⁽¹³⁴⁾ and this would have a commercial/operational impact, including delays and increased fuel use,⁽¹⁴¹⁾ since events arrive without warning and may persist for several hours.

A severe loss of HF radio may lead to communications with most aircraft in the north Atlantic being lost. Aircraft already in flight would continue, but those on the ground would probably not be allowed to take off.⁽¹⁷⁾ At high latitudes where satellite communications are unavailable, HF communication is mandatory. Polar routes have been disrupted by space weather and lost HF communications.⁽¹⁶⁾

3. ECONOMIC COST OF SPACE WEATHER

The literature studying the vulnerability of different industry sectors to space weather rarely extends the analysis to the actual quantification of economic losses resulting from space weather events. The few contributions available mainly focus on power grid losses. Some studies either present the views of the insurance sector or rely on its pricing models. Insurers' pricing models offer a robust methodological approach to economic cost quantification,¹² but details on data and methodology used are typically undisclosed. Very few scientific studies go beyond a scenario-based quantification of direct economic losses of specific sectors and exposures. When they do, they usually focus on a specific sector's vulnerability (e.g., power grids), and explore its propagation across other sectors via input–output analysis.⁽¹⁴⁶⁾

3.1. Broader Impact

The National Research Council's Committee on the Social and Economic Impacts of Severe Space

¹²Insurers decouple loss occurrences into a hazard component (event occurrence) and a loss severity component (damage conditional on the hazard event occurrence). The hazard event is translated into *direct* economic/social losses via a vulnerability function, which depends on the characteristics of the risk exposure (rating factors in the language of (re)insurance pricing models). The quantification of *indirect* losses (e.g., business interruption) instead typically relies on econometric models, input–output analysis, or equilibrium models.^(144–146)

Weather Events report summarizes a 2008 workshop and participants' views on current and future risks and vulnerabilities across different industry sectors.⁽¹⁶⁾ Although no holistic quantification of economic costs is attempted, the report collects information useful across a number of sectors, and provides suggestions for sector-specific risk mitigation techniques. It supports quantification based on approaches similar to insurance pricing models and catastrophe risk models. A 2011 OECD report supports the use of a threat-vulnerability-consequence template, but suggests that efforts should be aimed at going beyond insurers' focus on replacement costs to capture broader societal costs.⁽⁹⁾ The latter should rely on estimates of consumers' willingness to pay and opportunity costs. The report also emphasizes the importance of systemic risks arising from interconnected economies and sectors. However, no estimates of economic costs arising from compounding of losses via network interlinkages are provided.

The impact of severe space weather events on global supply chains and the global economy has recently been studied, explicitly considering both direct and indirect losses and adopting the input–output methodology for the first time. Restricted to the systemic effects of the power transmission system failure and interruption, “[f]or a 1989 Quebec-like event, the global economic impacts would range from \$2.4 – \$3.4 tn over a year.”⁽¹⁴⁷⁾ This analysis examines the implications of such an event occurring at different locations around the world (e.g., North America, China, and Western Europe) and assumes an outage of one-year duration based on the presumed long-lead times needed to replace destroyed transformers, as described in the next section.

3.2. Power

The estimated economic impact of the most severe events strongly depends on the assumed technological impact footprint, where there is some controversy. Several studies have assumed that a one in 100-year event (i.e., worst-case Carrington class) would cause catastrophic impact, with major transformer damage/failure and permanent loss of generator step-ups, taking a considerable length of time (4–10 years) to recover from. Generator step-ups are important because of compounding difficulties arising from network effects (loss of output of vital and usually baseload nuclear, coal, and hydroelectric generation resources for the power grid). The consequent economic impact is in the range of trillions of dollars because of the lack

of power for a very prolonged period.^(16,113) In a separate assessment that assumes extended power outages lasting from 16 days to one to two years, and minimum transformer replacement lead times of five months it has been suggested that the estimated total economic cost of a Carrington-level storm is \$0.6–\$2.6 tn⁽¹²⁾ in the United States. This is based on an affected population of 20–40 million. However, data and methodology are not fully disclosed.

A recently completed study also focuses on losses resulting from damages to transformers and associated power outages in the United States, resolved to the level of individual U.S. states taking account of geomagnetic latitude, ground conductivity, and the number of transformers in each state.⁽¹⁴⁸⁾ Three different “stress test” scenarios are presented to help inform the insurance industry about the possible range of impacts a severe space weather event may cause, including a plausible worst-case scenario where there is significant transformer damage causing prolonged power outage. The difficulty of procuring and installing replacement extra high-voltage transformers is discussed in detail. For each scenario, direct and indirect costs are calculated, the latter being quantified via input–output analysis. The total economic loss varies between \$0.5 tn and \$2.7 tn based on calculations examining disruption to the global supply chain. An alternative methodology finds a total loss of \$140–\$613 bn. This is lower as it accounts for the “dynamic response of the global economy.” Losses to U.S. GDP are estimated to range between \$136 bn and \$613 bn over five years following the space weather event, with the worst affected states being Illinois and New York. It also indicates significant knock-on impacts on the global economy with China, Canada, and Mexico being the worst affected (as they are the United States' largest trading partners), but also significant impacts on the United Kingdom, Japan, and Germany. The report⁽¹⁴⁸⁾ also develops analyses of impacts on insurers' payouts and investors' portfolios. The losses to the U.S. insurance industry are estimated as \$55.0–\$333.5 bn, and it is noted that this upper limit is “similar to the total insured losses from all catastrophes in 2015.”

The assumptions that lead to a very severe worse-case-scenario impact footprint are not universally accepted. Both the North American Electric Reliability Corporation (NERC) and the U.K. Royal Academy of Engineering (RAE) specifically find that the more likely impact is system collapse due to voltage instability rather than catastrophic infrastructure destruction.^(17,111,112) The NERC report examines past transformer failures, as well as

experimental data concerning transformer heating as a function of applied direct current. It is argued that during a major geomagnetic storm, design thresholds are unlikely to be exceeded. While transformers that are near end-of-life or employ older designs may be more at risk, it is nevertheless concluded that voltage instability is the most likely primary impact. The RAE report focuses on the United Kingdom in particular, and its conclusion is reached on the basis of studies and assessments undertaken by the National Grid. In particular, it is noted that since 1997, newly installed transformers have employed a more GIC-resistant design, which strengthens resilience. Outages are therefore measured in hours to days, rather than months, but such events still have a considerable economic impact through primary and secondary losses.⁽¹⁴⁹⁾ As examples, the economic impact of Hurricane Katrina was estimated to be \$81–\$125 bn⁽¹⁵⁰⁾ and the August 14, 2003 northeast blackout was \$4–\$10 bn.⁽¹⁵¹⁾ Analyses of historical blackout events in the United States indicate that even short blackouts, which occur several times during a year in the United States, sum up to an annual economic loss between \$104 bn and \$164 bn.⁽¹⁵²⁾ These figures are based on insurance industry pricing models for business interruption insurance. (Details on data and methodology are not publicly available.)

Space weather impacts are not necessarily restricted to catastrophic effects. Insurance claim information suggests that the losses to the U.S. power grid from noncatastrophic disturbances from GICs “may be \$5 – \$10 bn/year.”⁽¹⁵³⁾ The effect of space weather on generation outages, transmission congestion, wholesale real-time electricity prices, and resulting day-ahead prices has been examined,⁽¹⁵⁴⁾ as has its effects on electricity prices and spinning reserves using regression analysis to compute sensitivities.⁽¹⁵⁵⁾ A follow-up study examined GIC impact on different power grids using a variety of metrics, but translation into economic losses, however, was not addressed.⁽¹⁵⁶⁾ Finally, a study of South African power system impacts indicates that interruption costs correlate with business activity levels according to the seasons and time of day, and both this and the cost of interruption can be represented by β probability density functions.⁽¹⁰¹⁾

3.3. Satellites

It is likely that many encountered problems remain undisclosed due to commercial and security sensitivities.⁽¹⁷⁾ An initial attempt to quantify losses

from the bottom up by modeling factors well known to affect satellite resources (solar power erosion, orbit decay, etc.) led to an estimated \$70 bn cost from lost revenue and satellite replacement for a 1859-calibre superstorm.⁽¹⁵⁷⁾ The failure of Intelsat’s Galaxy-15 spacecraft in April 2010, probably due to space weather, provides a useful case study.⁽¹¹⁾ It is indicated that the satellite builder is spending around \$1 m on remedial actions and is facing the loss of payments linked to in-orbit performance worth \$7 m. As it was not even four years into its 10–15-year operational life, the potential total loss is estimated to be \$100 m based on a satellite cost of \$250 m.⁽¹¹⁾ Consequently, direct and indirect economic costs of space weather damage should be recoverable from publicly available information on length of satellite outage or replacement cost in case of total loss.^(17,123)

3.4. Other Sectors: GNSS, Aviation, Pipelines, and Transport

While there is no specific literature available on the economic cost of space weather impacts on GNSS, aviation, pipelines, or transport, initial efforts have been made that may aid analysis: for example, attempts have been made to quantify the fraction of U.S. economic activity dependent on GNSS services.⁽¹⁵⁸⁾ In aviation, it is reported that United Airlines closely monitors this risk dimension given its high numbers of polar routes.⁽¹⁶⁾ However, no estimates of economic/health costs are available beyond anecdotal evidence of operational costs ranging from flight delays and fuel stops resulting from diversion from polar routes following space weather events.⁽¹⁶⁾ Models have been developed to obtain spatial information about the distribution of pipeline GICs,⁽¹⁵⁹⁾ and this framework could be used to determine space weather economic cost from the bottom up. Finally, information on Russian and Swedish railway failures due to space weather events^(120,121,160) should be rich enough to estimate the associated economic costs based on transportation networks literature, and extrapolate to other countries.

4. CONCLUSIONS

Although space weather is now a widely recognized risk, its economic impact remains quite uncertain. Further work is required to comprehensively assess both direct and indirect losses across a diverse range of sectors. While the physical nature of space weather is the best-understood aspect of the problem, there is a lack of agreement in the realistic

technical footprint of the most severe space weather event, and this leads to dramatically divergent cost estimates. Therefore, a second research gap is to accurately quantify the technical impact of space weather in a variety of industries (power, satellite, aviation, GNSS, etc.). Given its rising importance, GNSS may be a particularly important sector to analyze. Two challenges stand out. First, modern society is yet to experience a Carrington-level space weather event, and so projections of economic impact will inevitably be subject to uncertainty. Second, quantifying technical impacts fundamentally relies on the participation of industry by providing appropriate data, and this often conflicts with commercial sensitivities. Since exposure to space weather risk is only likely to increase, urgent effort is required to address this tension.

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REFERENCES

- Hapgood M. Prepare for the coming space weather storm. *Nature*, 2012; 484(7394):311–313.
- Hapgood MA. Towards a scientific understanding of the risk from extreme space weather. *Advances in Space Research*, 2011; 47(12):2059–2072.
- Eastwood JP. The science of space weather. *Philosophical Transactions of the Royal Society A*, 2008; 366:4489–4500.
- Schwenn R. Space weather: The solar perspective. *Living Reviews in Solar Physics*, 2006; 3(2):1–72.
- Knipp D. *Understanding Space Weather and the Physics Behind it*. New York, USA: McGraw-Hill; 2011.
- Moldwin MB. *An Introduction to Space Weather*. Cambridge UK: Cambridge University Press; 2008.
- Schrijver CJ, Siscoe G. *Heliophysics: Space Storms and Radiation: Causes and Effects*. Cambridge UK: Cambridge University Press; 2012.
- Koskinen HEJ. *Physics of Space Storms: from the Solar Surface to the Earth*. Berlin DE: Springer Praxis; 2011.
- OECD. *Future Global Shocks. Improving Risk Governance*. Paris, FR: OECD; 2011.
- Howell L. *Global Risks 2013 Eighth Edition*. Switzerland CH: World Economic Forum; 2013.
- Hapgood M. Space weather: its impact on Earth and implications for business. London UK: Lloyd's, 2010.
- Lloyd's. *Solar Storm Risk to the North American Electric Grid*. London UK: Lloyd's, 2013.
- The Cabinet Office. *Space Weather Preparedness Strategy*. London UK: UK Government Department for Business Innovation and Skills BIS/15/457, 2015.
- The Cabinet Office. *National Risk Register of Civil Emergencies*. London UK: UK Government, 2015.
- SWORM Task Force. *National Space Weather Strategy*. Washington DC USA: United States Government National Science and Technology Council, 2015.
- National Research Council. *Severe Space Weather Events - Understanding Societal and Economic Impacts: A Workshop Report*. Washington DC: The National Academies Press; 2008.
- Cannon P, Angling M, Barclay L, Curry C, Dyer C, Edwards R, Greene G, Hapgood MA, Horne RB, Jackson D, Mitchell C, Owen J, Richards A, Ryden K, Saunders S, Sweeting M, Tanner R, Thomson A, Underwood C. *Extreme space weather: impacts on engineered systems and infrastructure*. London UK: Royal Academy of Engineering; 2013.
- Schrijver CJ, Kauristie K, Aylward AD, Denardini CM, Gibson SE, Glover A, Gopalswamy N, Grande M, Hapgood M, Heynderickx D, Jakowski N, Kalegaev VV, Lapenta G, Linker JA, Liu SP, Mandrini CH, Mann IR, Nagatsuma T, Nandy D, Obara T, O'Brien TP, Onsager T, Oppenorth HJ, Terkildsen M, Valladares CE, Vilmer N. *Understanding space weather to shield society: A global road map for 2015-2025 commissioned by COSPAR and ILWS*. *Advances in Space Research*, 2015; 55(12):2745–807.
- Fry EK. The risks and impacts of space weather: Policy recommendations and initiatives. *Space Policy*, 2012; 28(3):180–184.
- Mewaldt RA. Solar Energetic Particle Composition, Energy Spectra, and Space Weather. *Space Science Reviews*, 2006; 124(1–4):303–316.
- Reames DV. Magnetic topology of impulsive and gradual solar energetic particle events. *The Astrophysical Journal*, 2002; 571:L63–L6.
- Gonzalez WD, Tsurutani BT, Clua de Gonzalez AL. Interplanetary origin of geomagnetic storms. *Space Science Reviews*, 1999; 88:529–562.
- Richardson IG, Webb DF, Zhang J, Berdichevsky DB, Biesecker DA, Kasper JC, Kataoka R, Steinberg JT, Thompson BJ, Wu CC, Zhukov AN. Major geomagnetic storms ($Dst < -100$ nT) generated by corotating interaction regions. *Journal of Geophysical Research*, 2006; 111:A07S09.
- Tsurutani BT, Gonzalez WD, Gonzalez ALC, Guamieri FL, Gopalswamy N, Grande M, Kamide Y, Kasahara Y, Lu G, Mann I, McPherron R, Sorass F, Vasyliunas VM. Corotating solar wind streams and recurrent geomagnetic activity: A review. *Journal of Geophysical Research*, 2006; 111:A07S01.
- Kamide Y. What is an “Intense Geomagnetic Storm”? *Space Weather*, 2006; 4(6):S06008.
- Lanzerotti LJ. Comment on “Great magnetic storms” by Tsurutani et al. *Geophysical Research Letters*, 1992; 19(19):1991–1992.
- Thomson AWP, Dawson EB, Reay SJ. Quantifying extreme behavior in geomagnetic activity. *Space Weather*, 2011; 9(12):S10001.
- Pulkkinen A, Bernabeu E, Eichner J, Viljanen A, Ngwira C. Regional-scale high-latitude extreme geoelectric fields pertaining to geomagnetically induced currents. *Earth Planets and Space*, 2015; 67:1–8.
- Cliver EW. The 1859 space weather event: Then and now. *Advances in Space Research*, 2006; 38(2):119–129.
- Cliver EW, Dietrich WF. The 1859 space weather event revisited: limits of extreme activity. *Journal of Space Weather and Space Climate*, 2013; 3:A31.
- Cliver EW, Svalgaard L. The 1859 solar-terrestrial disturbance and the current limits of extreme space weather activity. *Solar Physics*, 2004; 224(1):407–422.
- Green JL, Boardsen S, Odenwald S, Humble J, Pazamickas KA. Eyewitness reports of the great auroral storm of 1859. *Advances in Space Research*, 2006; 38:145–154.
- Shea MA, Smart DF. Compendium of the eight articles on the “Carrington Event” attributed to or written by Elias Loomis in the *American Journal of Science*, 1859–1861. *Advances in Space Research*, 2006; 38:313–385.

34. Tsurutani BT, Gonzalez WD, Lakhina GS, Alex S. The extreme magnetic storm of 1-2 September 1859. *Journal of Geophysical Research*, 2003; 108(A7):1268.
35. Akasofu SI, Kamide Y. Comment on "The extreme magnetic storm of 1-2 September 1859" By B. T. Tsurutani, W. D. Gonzalez, G. S. Lakhina, and S. Alex. *Journal of Geophysical Research*, 2005; 110(A9):A09226.
36. Siscoe G, Crooker NU, Clauer CR. Dst of the Carrington storm of 1859. In: Clauer CR, Siscoe G, editors. *Great Historical Geomagnetic Storm of 1859: A Modern Look*. *Advances in Space Research*, 2006; 38:173–179.
37. Silverman SM, Cliver EW. Low-latitude auroras: the magnetic storm of 14-15 May 1921. *Journal of Atmospheric and Solar Terrestrial Physics*, 2001; 63(5):523–535.
38. Kappenman JG. Great geomagnetic storms and extreme impulsive geomagnetic field disturbance events – An analysis of observational evidence including the great storm of May 1921. *Advances in Space Research*, 2006; 38(2):188–199.
39. Boteler DH, Pirjola RJ, Nevanlinna H. The effects of geomagnetic disturbances on electrical systems at the earth's surface. *Advances in Space Research*, 1998; 22:17–27.
40. Knipp DJ, Ramsay AC, Beard ED, Boright AL, Cade WB, Hewins IM, McFadden RH, Denig WF, Kilcommons LM, Shea MA, Smart DF. The May 1967 great storm and radio disruption event: Extreme space weather and extraordinary responses. *Space Weather*, 2016; 14(9):614–633.
41. Gonzalez WD, Joselyn JA, Kamide Y, Kroehl HW, Rostoker G, Tsurutani BT, Vasyliunas VM. What is geomagnetic storm? *Journal of Geophysical Research*, 1994; 99(A4):5771–5792.
42. Bolduc L. GIC observations and studies in the Hydro-Quebec power system. *Journal of Atmospheric and Solar Terrestrial Physics*, 2002; 64:1793–1802.
43. Erimmez IA, Kappenman JG, Radasky WA. Management of the geomagnetically induced current risks on the national grid company's electric power transmission system. *Journal of Atmospheric and Solar Terrestrial Physics*, 2002; 64(5–6):743–756.
44. Gopalswamy N, Barbieri L, Cliver EW, Lu G, Plunkett SP, Skoug RM. Introduction to violent Sun-Earth connection events of October-November 2003. *Journal of Geophysical Research*, 2005; 110:A09S00.
45. Brodrick D, Tingay S, Wieringa M. X-ray magnitude of the 4 November 2003 solar flare inferred from the ionospheric attenuation of the galactic radio background. *Journal of Geophysical Research*, 2005; 110:A09S36.
46. Balch C. Intense space weather storms October 19–November 07, 2003. Service assessment. Silver Spring, MD USA: National Oceanic and Atmospheric Administration, 2004.
47. Barbieri LP, Mahmot RE. October-November 2003's space weather and operations lessons learned. *Space Weather*, 2004; 2(9):S09002.
48. Kappenman JG. An overview of the impulsive geomagnetic field disturbances and power grid impacts associated with the violent Sun-Earth connection events of 29–31 October 2003 and a comparative evaluation with the other contemporary storms. *Space Weather*, 2005; 3:S08C01.
49. Tsurutani BT, Verkhoglyadova OP, Mannucci AJ, Lakhina GS, Huba JD. Extreme changes in the dayside ionosphere during a Carrington-type magnetic storm. *Journal of Space Weather and Space Climate*, 2012; 2:A05.
50. Pulkkinen A, Lindahl S, Viljanen A, Pirjola R. Geomagnetic storm of 29-31 October 2003: Geomagnetically induced currents and their relation to problems in the Swedish high-voltage power transmission system. *Space Weather*, 2005; 3:S08C03.
51. Liu YD, Luhmann JG, Kajdič P, Kilpua EKJ, Lugaz N, Nitta NV, Möstl C, Lavraud B, Bale SD, Farrugia CJ, Galvin AB. Observations of an extreme storm in interplanetary space caused by successive coronal mass ejections. *Nature Communications*, 2014; 5:3481.
52. Baker DN, Li X, Pulkkinen A, Ngwira CM, Mays ML, Galvin AB, Simunac KDC. A major solar eruptive event in July 2012: Defining extreme space weather scenarios. *Space Weather*, 2013; 11(10):585–591.
53. Ngwira CM, Pulkkinen A, Mays ML, Kuznetsova MM, Galvin AB, Simunac K, Baker DN, Li X, Zheng Y, Gloer A. Simulation of the 23 July 2012 extreme space weather event: What if this extremely rare CME was Earth directed? *Space Weather*, 2013; 11(12):671–679.
54. Crosby N, Aschwanden M, Dennis B. Frequency distributions and correlations of solar X-ray flare parameters. *Solar Physics*, 1993; 143(2):275–299.
55. Riley P. On the probability of occurrence of extreme space weather events. *Space Weather*, 2012; 10:S02012.
56. Schrijver CJ, Beer J, Baltensperger U, Cliver EW, Güdel M, Hudson HS, McCracken KG, Osten RA, Peter T, Soderblom DR, Usoskin IG, Wolff EW. Estimating the frequency of extremely energetic solar events, based on solar, stellar, lunar, and terrestrial records. *Journal of Geophysical Research*, 2012; 117(A8):A08103.
57. Ruzmaikin A, Feynman J, Stoev SA. Distribution and clustering of fast coronal mass ejections. *Journal of Geophysical Research*, 2011; 116(A4):A04220.
58. Love JJ. Credible occurrence probabilities for extreme geophysical events: Earthquakes, volcanic eruptions, magnetic storms. *Geophysical Research Letters*, 2012; 39:L10301.
59. Balasis G, Papadimitriou C, Daglis IA, Anastasiadis A, Sandberg I, Eftaxias K. Similarities between extreme events in the solar-terrestrial system by means of nonextensivity. *Nonlinear Processes in Geophysics*, 2011; 18(5):563–572.
60. Lepreti F, Carbone V, Veltri P. Solar flare waiting time distribution: varying-rate Poisson or Levy function? *The Astrophysical Journal*, 2001; 555:L133–L136.
61. Stanislavsky AA, Burnecki K, Magdziarz M, Weron A, Weron K. Farima Modeling of Solar Flare Activity From Empirical Time Series of Soft X-Ray Solar Emission. *The Astrophysical Journal*, 2009; 693(2):1877.
62. Bak P, Tang C, Wiesenfeld K. Self-organized criticality: An explanation of the $1/f$ noise. *Physical Review Letters*, 1987; 59(4):381–384.
63. Lu ET, Hamilton RJ. Avalanches and the distribution of solar flares. *The Astrophysical Journal*, 1991; 380:L89–L92.
64. Georgoulis MK, Vilmer N, Crosby NB. A comparison between statistical properties of solar X-ray flares and avalanche predictions in cellular automata statistical flare models. *Astronomy & Astrophysics*, 2001; 367(1):326–338.
65. Georgoulis MK, Vlahos L. Variability of the occurrence frequency of solar flares and the statistical flare. *Astronomy & Astrophysics*, 1998; 336:721–724.
66. Baiesi M, Paczuski M, Stella AL. Intensity Thresholds and the Statistics of the Temporal Occurrence of Solar Flares. *Physical Review Letters*, 2006; 96(5):051103.
67. Hughes D, Paczuski M, Dendy RO, Helander P, McClements KG. Solar Flares as Cascades of Reconnecting Magnetic Loops. *Physical Review Letters*, 2003; 90(13):131101.
68. Paczuski M, Boettcher S, Baiesi M. Interoccurrence Times in the Bak-Tang-Wiesenfeld Sandpile Model: A Comparison with the Observed Statistics of Solar Flares. *Physical Review Letters*, 2005; 95(18):181102.
69. Wheatland MS, Litvinenko YE. Understanding solar flare waiting-time distributions. *Solar Physics*, 2002; 211:255–274.
70. Wheatland MS. The coronal mass ejection waiting time distribution. *Solar Physics*, 2003; 214:361–373.
71. Tsubouchi K, Omura Y. Long-term occurrence probabilities of intense geomagnetic storm events. *Space Weather*, 2007; 5(12):S12003.

72. Kataoka R. Probability of occurrence of extreme magnetic storms. *Space Weather*, 2013; 11(5):214–218.
73. Schrijver CJ, Beer J. Space weather from explosions on the Sun: how bad could it be? *EOS Transactions of the American Geophysical Union*, 2014; 95(24):201–202.
74. Maehara H, Shibayama T, Notsu S, Notsu Y, Nagao T, Kusaba S, Honda S, Nogami D, Shibata K. Superflares on solar-type stars. *Nature*, 2012; 485(7399):478–481.
75. Shibayama T, Maehara H, Notsu S, Notsu Y, Nagao T, Honda S, Ishii TT, Nogami D, Shibata K. Superflares on solar-type stars observed with Kepler. I. Statistical properties of superflares. *Astrophysical Journal Supplement Series*, 2013; 209(1):5.
76. Maehara H, Shibayama T, Notsu Y, Notsu S, Honda S, Nogami D, Shibata K. Statistical properties of superflares on solar-type stars based on 1-min cadence data. *Earth Planets and Space*, 2015; 67:1–10.
77. Nogami D, Notsu Y, Honda S, Maehara H, Notsu S, Shibayama T, Shibata K. Two sun-like superflare stars rotating as slow as the Sun. *Publications of the Astronomical Society of Japan*. 2014; 66(2):L4.
78. Shibata K, Isobe H, Hillier A, Choudhuri AR, Maehara H, Ishii TT, Shibayama T, Notsu S, Notsu Y, Nagao T, Honda S, Nogami D. Can Superflares Occur on Our Sun? *Publications of the Astronomical Society of Japan*. 2013;65(3):49.
79. Kitze M, Neuhaeuser R, Hambaryan V, Ginski C. Superflares on the slowly rotating solar-type stars KIC10524994 and KIC07133671? *Monthly Notices of the Royal Astronomical Society*, 2014; 442(4):3769–76.
80. Shea MA, Smart DF. A summary of major solar proton events. *Solar Physics*, 1990; 127(2):297–320.
81. Gold T, Palmer DR. The solar outburst, 23 February 1956. Observations by the Royal Greenwich Observatory. *Journal of Atmospheric and Solar Terrestrial Physics*, 1956; 8:287–291.
82. Rishbeth H, Shea MA, Smart DF. The solar-terrestrial event of 23 February 1956. *Advances in Space Research*, 2009; 44(10):1096–1106.
83. Wolff EW, Bigler M, Curran MAJ, Dibb JE, Frey MM, Legrand M, McConnell JR. The Carrington event not observed in most ice core nitrate records. *Geophysical Research Letters*, 2012; 39:L08503.
84. Usoskin IG, Solanki SK, Kovaltsov GA, Beer J, Kromer B. Solar proton events in cosmogenic isotope data. *Geophysical Research Letters*, 2006; 33:L08107.
85. Steinhilber F, Abreu JA, Beer J, Brunner I, Christl M, Fischer H, Heikkilä U, Kubik PW, Mann M, McCracken KG, Miller H, Miyahara H, Oerter H, Wilhelms F. 9,400 years of cosmic radiation and solar activity from ice cores and tree rings. *Proceedings of the National Academy of Sciences*, 2012; 109(16):5967–5971.
86. Miyake F, Nagaya K, Masuda K, Nakamura T. A signature of cosmic-ray increase in AD 774–775 from tree rings in Japan. *Nature*, 2012; 486(7402):240–242.
87. Usoskin IG, Kromer B, Ludlow F, Beer J, Friedrich M, Kovaltsov GA, Solanki SK, Wacker L. The AD775 cosmic event revisited: the Sun is to blame. *Astronomy & Astrophysics*, 2013; 552:L3.
88. Sigl M, Winstrup M, McConnell JR, Welten KC, Plunkett G, Ludlow F, Buentgen U, Caffee M, Chellman N, Dahl-Jensen D, Fischer H, Kipfstuhl S, Kostick C, Maselli OJ, Mekhaldi F, Mulvaney R, Muscheler R, Pasteris DR, Pilcher JR, Salzer M, Schuepbach S, Steffensen JP, Vinther BM, Woodruff TE. Timing and climate forcing of volcanic eruptions for the past 2,500 years. *Nature*, 2015; 523(7562):543–549.
89. Guettler D, Adolphi F, Beer J, Bleicher N, Boswijk G, Christl M, Hogg A, Palmer J, Vockenhuber C, Wacker L, Wunder J. Rapid increase in cosmogenic C-14 in AD 775 measured in New Zealand kauri trees indicates short-lived increase in C-14 production spanning both hemispheres. *Earth and Planetary Science Letters*, 2015; 411:290–297.
90. Thomas BC, Melott AL, Arkenberg KR, Snyder BR, II. Terrestrial effects of possible astrophysical sources of an AD 774–775 increase in C-14 production. *Geophysical Research Letters*, 2013; 40(6):1237–1240.
91. Mekhaldi F, Muscheler R, Adolphi F, Aldahan A, Beer J, McConnell JR, Possnert G, Sigl M, Svensson A, Synal HA, Welten KC, Woodruff TE. Multiradionuclide evidence for the solar origin of the cosmic-ray events of AD 774/5 and 993/4. *Nature Communications*, 2015; 6:8611.
92. Melott AL, Thomas BC. Causes of an AD 774–775 C-14 increase. *Nature*, 2012; 491(7426):E1–E2.
93. Kilpua EKJ, Olsper N, Grigorievskiy A, Kämpylä MJ, Tanskanen EI, Miyahara H, Kataoka R, Pelt J, Liu YD. Statistical Study of Strong and Extreme Geomagnetic Disturbances and Solar Cycle Characteristics. *The Astrophysical Journal*, 2015; 806(2):272.
94. Knipp DJ. Foreword to space weather collection on geomagnetically induced currents: Commentary and research. *Space Weather*, 2015; 13(11):742–746.
95. Juusola L, Viljanen A, van de Kamp M, Tanskanen EI, Vanhamaki H, Partamies N, Kauristie K. High-latitude ionospheric equivalent currents during strong space storms: Regional perspective. *Space Weather*, 2015; 13(1):49–60.
96. Pulkkinen A, Viljanen A, Pirjola R. Estimation of geomagnetically induced current levels from different input data. *Space Weather*, 2006; 4(8):S08005.
97. Showstack R. Threat of severe space weather to the US electrical grid explored at conference. *EOS Transactions American Geophysical Union*, 2011; 92:374–375.
98. Carter BA, Yizengaw E, Pradipta R, Halford AJ, Norman R, Zhang K. Interplanetary shocks and the resulting geomagnetically induced currents at the equator. *Geophysical Research Letters*, 2015; 42(16):6554–6559.
99. Kappenman JG. Storm sudden commencement events and the associated geomagnetically induced current risks to ground-based systems at low-latitude and midlatitude locations. *Space Weather*, 2003; 1(3):1016.
100. Lotz SI, Cilliers PJ. A solar wind-based model of geomagnetic field fluctuations at a mid-latitude station. *Advances in Space Research*, 2015; 55(1):220–230.
101. Gaunt CT. Reducing uncertainty – responses for electricity utilities to severe solar storms. *Journal of Space Weather and Space Climate*, 2014; 4:A01.
102. Thomson AWP, McKay AJ, Clarke E, Reay SJ. Surface electric fields and geomagnetically induced currents in the Scottish Power grid during the 30 October 2003 geomagnetic storm. *Space Weather*, 2005; 3:S11002.
103. Pirjola R, Kauristie K, Lappalainen H, Viljanen A, Pulkkinen A. Space weather risk. *Space Weather*, 2005; 3(2):S02A02.
104. Torta JM, Serrano L, Regue JR, Sanchez AM, Roldan E. Geomagnetically induced currents in a power grid of northeastern Spain. *Space Weather*, 2012; 10(11):S06002.
105. Matandirotya E, Cilliers PJ, Van Zyl RR. Modeling geomagnetically induced currents in the South African power transmission network using the finite element method. *Space Weather*, 2015; 13(3):185–195.
106. Gaunt CT, Coetzee G. Transformer failures in regions incorrectly considered to have low GIC-risk. *Lausanne CH: 2007 IEEE Lausanne Power Tech*, 2007; 807–812.
107. Watari S. Estimation of geomagnetically induced currents based on the measurement data of a transformer in a Japanese power network and geoelectric field observations. *Earth Planets and Space*, 2015; 67(12):77.
108. Wang K-R, Lian-guang LIU, Yan LI. Preliminary Analysis on the Interplanetary Cause of Geomagnetically Induced

- Current and Its Effect on Power Systems. *Chinese Astronomy and Astrophysics*, 2015; 39(1):78–88.
109. Trivedi NB, Vitorello I, Kabata W, Dutra SLG, Padilha AL, Bologna MS, de Padua MB, Soares AP, Luz GS, Pinto FA, Pirjola R, Viljanen AI. Geomagnetically induced currents in an electric power transmission system at low latitudes in Brazil: A case study. *Space Weather*, 2007; 5(4): S04004.
 110. Forbes KF, St Cyr OC. Did geomagnetic activity challenge electric power reliability during solar cycle 23? Evidence from the PJM regional transmission organization in North America. *Space Weather*, 2012; 10:S05001.
 111. NERC. High-impact, low frequency event risk to the north american power system. Atlanta GA USA: North America Electric Reliability Cooperation, 2010.
 112. NERC. Effects of geomagnetic disturbances on the bulk power system. Atlanta GA USA: North America Electric Reliability Corporation, 2012.
 113. Kappenman J. Geomagnetic storms and their impact on the US power grid. Goleta CA USA: Metatech Corp., 2010.
 114. Schrijver CJ, Mitchell SD. Disturbances in the US electric grid associated with geomagnetic activity. *Journal of Space Weather and Space Climate*, 2013; 3:A19.
 115. Viljanen A, Pulkkinen A, Pirjola R, Pajunpaa K, Posio P, Koistinen A. Recordings of geomagnetically induced currents and a nowcasting service of the Finnish natural gas pipeline system. *Space Weather*, 2006; 4(10):S10004.
 116. Kintner PM, O'Hanlon B, Gary DE, Kintner PMS. Global Positioning System and solar radio burst forensics. *Radio Science*, 2009; 44:RS0A08.
 117. Kelly MA, Comberiate JM, Miller ES, Paxton LJ. Progress toward forecasting of space weather effects on UHF SATCOM after Operation Anaconda. *Space Weather*, 2014; 12(10):601–611.
 118. Lanzerotti LJ, Medford LV, MacLennan CG, Kraus JS, Kappenman J, Radasky W. Trans-atlantic geopotentials during the July 2000 solar event and geomagnetic storm. *Solar Physics*, 2001; 204(1–2):351–359.
 119. Medford LV, Lanzerotti LJ, Kraus JS, MacLennan CG. Transatlantic Earth potential variations during the March 1989 magnetic storms. *Geophysical Research Letters*, 1989; 16(10):1145–1148.
 120. Eroshenko EA, Belov AV, Boteler D, Gaidash SP, Lobkov SL, Pirjola R, Trichtchenko L. Effects of strong geomagnetic storms on Northern railways in Russia. *Advances in Space Research*, 2010; 46(9):1102–1110.
 121. Wik M, Pirjola R, Lundstedt H, Viljanen A, Wintoft P, Pulkkinen A. Space weather events in July 1982 and October 2003 and the effects of geomagnetically induced currents on Swedish technical systems. *Annales Geophysicae*, 2009; 27(4):1775–1787.
 122. Brunstrom A. Stargazing. *Rail Technology Magazine*, 2011; (Apr/May):188–189.
 123. Koons HC, Mazur JE, Selesnick RS, Blake JB, Fennell JF, Roeder JL, Anderson PC. The impact of the space environment on space systems. Contract No.: TR-99(1670)-1. El Segundo, CA: The Aerospace Corporation, 1999.
 124. Choi HS, Lee J, Cho KS, Kwak YS, Cho IH, Park YD, Kim YH, Baker DN, Reeves GD, Lee DK. Analysis of GEO spacecraft anomalies: Space weather relationships. *Space Weather*, 2011; 9:S06001.
 125. Fennell JF, Koons HC, Roeder JL, Blake JB. Spacecraft charging: Observations and relationship to satellite anomalies. Pp. 279–285 in: Harris RA (ed). *Proceedings of the 7th Spacecraft Charging Technology Conference: 2001: A Spacecraft Charging Odyssey*. Noordwijk NL: Esa Special Publication SP476.
 126. Wrenn GL, Rodgers DJ, Ryden KA. A solar cycle of spacecraft anomalies due to internal charging. *Annales Geophysicae*, 2002; 20(7):953–956.
 127. Hapgood MA. Report on demonstration of correlation of spacecraft anomalies with environment data. Rutherford Appleton Laboratory, Harwell UK: STFC, 2001 Contract No.: RAL-SED-RP-0303.
 128. Fazakerley AN, Carter PJ, Watson G, Spencer A, Sun YQ, Coker J, Kataria DO, Fontaine D, Liu ZX, Gilbert L, He L, Lahiff AD, Mihaljcic B, Szita S, Taylor MGGT, Wilson RJ, Dedieu M, Schwartz SJ. The Double Star plasma electron and current experiment. *Annales Geophysicae*, 2005; 23(8):2733–2756.
 129. Loto'aniu TM, Singer HJ, Rodriguez JV, Green J, Denig W, Biesecker D, Angelopoulos V. Space weather conditions during the Galaxy 15 spacecraft anomaly. *Space Weather*, 2015; 13(8):484–502.
 130. Campbell A, Buchner S, Petersen E, Blake B, Mazur J, Dyer C. SEU measurements and predictions on MPTB for a large energetic solar particle event. *IEEE Transactions on Nuclear Science*, 2002; 49(3):1340–1344.
 131. Lohmeyer WQ, Cahoy K. Space weather radiation effects on geostationary satellite solid-state power amplifiers. *Space Weather*, 2013; 11(8):476–488.
 132. Yearby KH, Balikhin M, Walker SN. Single-event upsets in the Cluster and Double Star Digital Wave Processor instruments. *Space Weather*, 2014; 12(1):24–28.
 133. Kintner PM, Ledvina BM, de Paula ER. GPS and ionospheric scintillations. *Space Weather*, 2007; 5:S09003.
 134. Dyer CS, Lei F, Clucas SN, Smart DF, Shea MA. Solar particle enhancements of single-event effect rates at aircraft altitudes. *IEEE Transactions on Nuclear Science*, 2003; 50(6):2038–2045.
 135. Mishev AL. Computation of radiation environment during ground level enhancements 65, 69 and 70 at equatorial region and flight altitudes. *Advances in Space Research*, 2014; 54(3):528–535.
 136. Spurny F, Kudela K, Dachev T. Airplane radiation dose decrease during a strong Forbush decrease. *Space Weather*, 2004; 2(5):S05001.
 137. Getley IL. Observation of solar particle event on board a commercial flight from Los Angeles to New York on 29 October 2003. *Space Weather*, 2004; 2(5):S05002.
 138. Getley IL, Duldig ML, Smart DF, Shea MA. Radiation dose along North American transcontinental flight paths during quiescent and disturbed geomagnetic conditions. *Space Weather*, 2005; 3(1):S01004.
 139. Dyer CS, Truscott PR. Cosmic radiation effects on avionics. *Microprocessors and Microsystems*, 1999; 22(8):477–483.
 140. Stassinopoulos EG, Stauffer CA, Brucker GJ. A systematic global mapping of the radiation field at aviation altitudes. *Space Weather*, 2003; 1(1):1005.
 141. Jones JBL, Bentley RD, Hunter R, Iles RHA, Taylor GC, Thomas DJ. Space weather and commercial airlines. *Advances in Space Research*, 2005; 36(12):2258–2267.
 142. Lindborg L, Bartlett D, Beck P, McAulay I, Schnuer K, Schraube H, Spurny F. Cosmic radiation exposure of aircraft crew: Compilation of measured and calculated data. *Radiation Protection Dosimetry*, 2004; 110(1–4):417–422.
 143. Sigurdson AJ, Ron E. Cosmic radiation exposure and cancer risk among flight crew. *Cancer Investigation*, 2004; 22(5):743–761.
 144. Rose A. Input-output economics and computable general equilibrium models. *Structural Change and Economic Dynamics*, 1995; 6:295–304.
 145. Rose A. Model validation in estimating higher-order economic losses from natural hazards. Pp. 105–131 in: Taylor C, van Marcke E, editors. *Acceptable risk to lifeline systems*

- from natural hazard threats. New York: American Society of Civil Engineers; 2002.
146. Miller RE, Blair PD. Input-output analysis: foundations and extensions. Cambridge UK: Cambridge University Press; 2009.
147. den Baeumen HSI, Moran D, Lenzen M, Cairns I, Steenge A. How severe space weather can disrupt global supply chains. *Natural Hazards and Earth System Sciences*, 2014; 14(10):2749–2759.
148. Cambridge Centre for Risk Studies. Helios Solar Storm Scenario. Cambridge UK: University of Cambridge Centre for Risk Studies, 2016.
149. Hertzfeld HR, Williamson RA, Sen A. Weather satellites and the economic value of forecasts: evidence from the electric power industry. *Acta Astronautica*, 2004; 55:791–802.
150. Blake ES, Rappaport EN, Landsea CW. The deadliest, costliest and most intense United States tropical cyclones from 1851 to 2006 (and other frequently requested hurricane facts). In: NOAA (ed). Miami, FL: National Weather Service National Hurricane Center, 2007.
151. US - Canada Power System Outage Task Force. Final report on the implementation of the task force recommendations. Washington DC USA: Natural Resources Canada and US Department of Energy, 2006.
152. Bruch M, Munch V, Aichinger M, Kuhn M, Weymann M. Power Blackout Risks. Amsterdam, The Netherlands: CRO Forum, 2011.
153. Schrijver CJ, Dobbins R, Murtagh W, Petrinec SM. Assessing the impact of space weather on the electric power grid based on insurance claims for industrial electrical equipment. *Space Weather*. 2014; 12(7):487–498.
154. Forbes KF, St Cyr OC. Space weather and the electricity market: An initial assessment. *Space Weather*, 2004; 2:S10003.
155. Forbes KF, St Cyr OC. Solar Activity and Economic Fundamentals: The Case of the Electricity Market in Texas. 86th AMS Annual Meeting. Third Symposium on Space Weather. Atlanta, GA: American Meteorological Society, 2006.
156. Forbes KF, St. Cyr OC. Solar activity and economic fundamentals: Evidence from 12 geographically disparate power grids. *Space Weather*, 2008; 6(10):S10003.
157. Odenwald S, Green J, Taylor W. Forecasting the impact of an 1859-calibre superstorm on satellite resources. *Advances in Space Research*, 2006; 38:280–297.
158. Pham ND. The economic benefits of commercial GPS use in the US and the costs of potential disruption. Washington, DC: NDP Consulting, 2011.
159. Trichtchenko L, Boteler DH. Specification of geomagnetically induced electric fields and currents in pipelines. *Journal of Geophysical Research*, 2001; 106(A10):21039–21048.
160. Ptitsyna NG, Kasinskii VV, Villoresi G, Lyahov NN, Dorman LI, Iucci N. Geomagnetic effects on mid-latitude railways: A statistical study of anomalies in the operation of signaling and train control equipment on the East-Siberian Railway. *Advances in Space Research*, 2008; 42(9):1510–1514.